

CAPVT II.
DE
MICROSCOPIIS HVIVS GENERIS
MAGIS COMPOSITIS.

Problema I.

250.

Microscopium huius generis ex quinque lentibus
construere, quae ita sint dispositae, vt prior
imago realis inter lentem secundam et tertiam, po-
sterior vero inter tertiam et quartam cadat.

Solutio.

Cum igitur prior imago in interuallum secun-
dum, posterior vero in tertium cadere debeat, litte-
rarum P, Q, R, S secunda et tertia Q et R debent
esse negativae. Statuatur ergo $Q = -k$ et $R = -k'$,
vt sit $P k k' S = M$, existente $M = \frac{m a}{b}$. Deinde
vero sit $M = \frac{q+r+s+t}{M-1}$, vt fiat spatii in obiecto
conspicui femidiameter $= z = M a \xi$. Quare vt
campus euadat maximus, efficiendum est, vt littera-
rum q, r, s, t tot fiant vnitati aequales, quam reli-

quae circumstantiae permittunt, quod cum de omnibus statui nequeat, saltem pro postremis ponamus

$$s = 1 \text{ et } t = 1, \text{ ut sit } M = \frac{q+r+a}{m-1}.$$

Tum vero habebuntur sequentes aequationes

$$1^{\text{ma}} \quad Bq = (P - r)M$$

$$2^{\text{da}} \quad Cr = -(Pk + 1)M - q$$

$$3^{\text{tia}} \quad D = (Pkk' - 1)M - q - r$$

quibus adiungatur aequatio, qua margo coloratus destruitur,

$$0 = \frac{q}{P} - \frac{r}{Pk} + \frac{r}{Pkk'} + \frac{1}{Pkk's}$$

ex qua deducitur

$$k' = \frac{1 + \frac{q}{s}}{r - kq}$$

unde patet, esse debere $r > kq$. Ante quam autem hinc quidquam definire valeamus, lentium intervalla considerare debemus, quae sunt

$$1^{\text{mum}} = Aa \left(1 - \frac{1}{P}\right)$$

$$2^{\text{dum}} = -\frac{AB}{P} a \left(1 + \frac{1}{k}\right)$$

$$3^{\text{tium}} = -\frac{ABG}{Pk} a \left(1 + \frac{1}{k}\right)$$

$$4^{\text{tum}} = -\frac{ABCD}{Pkk'} a \left(1 - \frac{1}{s}\right)$$

quae omnia debent esse positiva, quibus adiungatur distantia focalis ultimae lentis

$$t = \frac{ABCD}{Pkk's} a = \frac{ABCD}{m} a;$$

quae

quae etiam debet esse positua, vt prodeat distantia oculi post eam, $O = \frac{r}{M}$, positua, vnde ob $t = r$ euidens est, esse debere t posituum, ideoque $ABCD > 0$. Quocirca vt quartum interuallum etiam fiat posituum, necesse est, vt sit $r - \frac{r}{s} < 0$, siue $S < r$. vnde euidens est, productum Pkk' fore $> M$ ideoque numerum praemagnu; vnde tertia illa aequatio, si loco M eius valor substituatur, dabit

$$D = \frac{(Pkk' - 1)(q + r + 2)}{M - 1} - q - r$$

vbi cum Pkk' et M sint numeri praemagni et $Pkk' > M$, fiat proxime

$$D = \frac{Pkk'}{M} (q + r + 2) - q - r \\ = \frac{Pkk'}{M} + (q + r) \left(\frac{Pkk'}{M} - 1 \right)$$

siue ob

$$Pkk' = \frac{M}{s}, \text{ erit } D = \frac{2}{s} + (q + r) \left(\frac{1}{s} - 1 \right);$$

vnde cum $q + r$ certe sit < 2 , euidens est, fore $D > r$. ideoque D negatiuum. Erit ergo $ABC < 0$. hinc tertium interuallum sponte fit posituum. Quare cum ob secundum interuallum esse debeat $AB < 0$, oportet esse $C > 0$ hincque $C > 0$ et $C < r$. Vnde si fuerit $A > 0$, ideoque $P > r$ ob interuallum primum; debeat esse $B < 0$ hincque vel $B < 0$ vel $B > r$; vnde sequitur fore priori casu $q < 0$ altero casu $q > 0$; vnde ob $q + Cr < 0$ fieret $r < 0$ ideoque $k' < s$ quod est absurdum. Sin autem esset $A < 0$

$A < 0$ hincque $P < 1$, debent esse $B > 0$ ideoque $\mathfrak{B} > 0$ et $\mathfrak{B} < 1$ unde iterum fieret $q < 0$. Neque igitur etiamnum constat, vtrum ambo isti casus consistere queant. Quia vero in vtroque fit $q < 0$, statuamus, vt ante, $q = -\omega$, vt fit $\omega = \frac{(1-P)M}{\mathfrak{B}}$ et ob secundam aequationem esse debet $\omega - \mathfrak{C}r > 0$, tum vero erit

$$k' = \frac{1 + \frac{1}{\mathfrak{B}}}{r + k\omega};$$

unde ob $Pk' = \frac{\mathfrak{M}}{\mathfrak{B}}$, fiet $Pk = \frac{\mathfrak{M}(r+k\omega)}{1+\mathfrak{B}}$.

Nunc igitur litteras ω et r ex calculo eliminemus et cum fit $\omega = \frac{1-P}{\mathfrak{B}} M$ ponamus breuitatis gratia $\frac{1-P}{\mathfrak{B}} = \zeta$ vt fit $\omega = \zeta M$; deinde fit etiam breuitatis gratia

$$1 + Pk = \eta \text{ fietque } r = \frac{(\zeta - \eta)M}{\mathfrak{C}}.$$

Ergo porro

$$r - \omega = \frac{(\zeta(1-\mathfrak{C}) - \eta)M}{\mathfrak{C}} \text{ et}$$

$$2 - \omega + r = \frac{2\mathfrak{C} + (\zeta(1-\mathfrak{C}) - \eta)M}{\mathfrak{C}}$$

at cum fit $M = \frac{2+r-\omega}{\mathfrak{M}-1}$ erit

$$2 + r - \omega = M(\mathfrak{M} - 1);$$

unde concludimus fore

$$M = \frac{2\mathfrak{C}}{\mathfrak{C}(\mathfrak{M} + \zeta - 1) - \zeta + \eta}$$

quo valore inuento erit

$$\omega = \frac{2\mathfrak{C}\zeta}{\mathfrak{C}(\mathfrak{M} + \zeta - 1) - \zeta + \eta} \text{ et } r = \frac{2(\zeta - \eta)}{\mathfrak{C}(\mathfrak{M} + \zeta - 1) - \zeta + \eta}$$

ex

ex quibus valoribus porro conficitur

$$r + k\omega = \frac{2(\zeta - \eta) + 2k\mathcal{E}\zeta}{\mathcal{E}(\mathfrak{M} + \zeta - 1) - \zeta + \eta}$$

hincque

$$Pk = \frac{2\mathfrak{M}(\zeta - \eta + k\mathcal{E}\zeta)}{(1+S)(\mathcal{E}(\mathfrak{M} + \zeta - 1) - \zeta + \eta)} = \eta - 1$$

ex quo definitur

$$k = \frac{(\eta - 1)(1+S)\mathcal{E}(\mathfrak{M} + \zeta - 1) - 2\mathfrak{M}(\zeta - \eta) - (\eta - 1)(1+S)(\zeta - \eta)}{2\mathfrak{M}\mathcal{E}\zeta}$$

vnde porro inuenitur $P = \frac{\eta - 1}{k}$. Quia autem k debet esse positium, in eius numeratore coefficiens ipsius \mathfrak{M} debet esse positius, vnde sequitur $\mathcal{E} > \frac{2(\zeta - \eta)}{(\eta - 1)(1+S)}$ at vero vidimus esse $\mathcal{E} < 1$, sicque necesse est, vt sit

$$\frac{2(\zeta - \eta)}{(\eta - 1)(1+S)} < 1 \text{ feu } 2(\zeta - \eta) < (\eta - 1)(1+S);$$

ex qua conditione concludimus

$$\zeta < \eta + \frac{(\eta - 1)(1+S)}{2}.$$

Nouimus autem esse debere $\zeta > \eta$; vnde littera ζ capi debet intra limites η et $\eta + \frac{(\eta - 1)(1+S)}{2}$, vbi manifestum est esse debere $\eta > 1$. Nostrum igitur problema sequenti modo resolui conueniet: Pro lubitu capi possunt litterae S et η , dummodo obseruetur, esse debere $S < 1$ et $\eta > 1$, quia $Pk = \eta - 1$. Deinde littera ζ capiatur intra limites η et $\eta + \frac{(\eta - 1)(1+S)}{2}$. At \mathcal{E} capiatur intra hos limites

$$\mathcal{E} < 1 \text{ simulque } \mathcal{E} > \frac{2(\zeta - \eta)}{(\eta - 1)(1+S)}$$

vnde simul C definitur: Tum vero capiatur

$$k = \frac{(\eta-1)(1+s)\mathcal{C}(\mathcal{M}+\zeta-1) - 2\mathcal{M}(\zeta-\eta) - (\eta-1)(1+s)(\zeta-\eta)}{2\mathcal{M}\mathcal{C}\zeta}$$

ex quo habebitur

$$P = \frac{\eta-1}{k} \text{ et } k' = \frac{\mathcal{M}}{(\eta-1)s}$$

Postea capiatur $\mathcal{B} = \frac{1-P}{\zeta}$; vnde definitur B. Denique ob $Pkk' = \frac{\mathcal{M}}{s}$ formula supra pro D inuenta dabit

$$\mathcal{D} = \left(\frac{\mathcal{M}}{s} - 1\right)M - (\mathcal{M} - 1)M + 2 \\ = M\mathcal{M}\left(\frac{1}{s} - 1\right) + 2;$$

ubi substituto valore ipsius M prodit

$$\mathcal{D} = \frac{2\mathcal{M}\mathcal{C} + 2\mathcal{C}(\zeta-1) - 2(\zeta-\eta)}{\mathcal{C}(\mathcal{M} + \zeta - 1) - \zeta + \eta}$$

qui valor, cum sit \mathcal{M} numerus praemagnus, erit proxime $\mathcal{D} = \frac{2}{s}$; hincque accuratius

$$\mathcal{D} = \frac{2}{s} + \frac{\frac{2}{s}(\zeta-\eta) - \mathcal{C}(\zeta-1) + 2\mathcal{C}(\zeta-1) - 2(\zeta-\eta)}{\mathcal{C}(\mathcal{M} + \zeta - 1) - \zeta + \eta}$$

sive

$$\mathcal{D} = \frac{2}{s} + \frac{2\left(\frac{1}{s} - 1\right)(\zeta - \eta - \mathcal{C}(\zeta - 1))}{\mathcal{C}(\mathcal{M} + \zeta - 1) - \zeta + \eta}$$

vel

$$\mathcal{D} = \frac{2}{s} - \frac{2\left(\frac{1}{s} - 1\right)(\mathcal{C}(\zeta - 1) - \zeta + \eta)}{\mathcal{C}(\mathcal{M} + \zeta - 1) - \zeta + \eta}$$

vnde saltem patet, certe fore $\mathcal{D} > 1$. ideoque D negati-

gatiuum, vt supra iam notauimus. Iam prouti fuerit P siue > 1 siue < 1 , capi debet vel $A > 0$ vel $A < 0$; ita, vt nunc omnia elementa sint determinata; habebuntur enim distantiae focales

$$p = \mathcal{A}a; \quad q = -\frac{AB}{P}a; \quad r = -\frac{ABC}{Pk}a;$$

$$s = -\frac{ABCD}{Pkk'}a \quad \text{et} \quad t = \frac{ABCD.a}{m};$$

deinde interualla

$$\text{I}^{\text{um}} = Aa \left(1 + \frac{1}{P}\right);$$

$$\text{II}^{\text{um}} = -\frac{AB.a}{P} \left(1 + \frac{1}{k}\right);$$

$$\text{III}^{\text{um}} = -\frac{ABC.a}{Pk} \left(1 + \frac{1}{k'}\right);$$

$$\text{IV}^{\text{um}} = -\frac{ABCD}{Pkk'}a \left(1 - \frac{1}{s}\right);$$

tum vero erit

$$M = 2: \left(\mathcal{M} + \zeta - 1 - \frac{(\zeta - \eta)}{c} \right) \quad \text{et} \quad O = \frac{c}{\mathcal{M}}$$

et apertura primae lentis ex aequatione notissima $\frac{1}{k^2} = \dots$ definietur.

Coroll. I.

251. Conditio, quam inuenimus $\zeta > \eta$ hic plus non inuoluit, quam ne ζ minus, quam η , accipiatur. Nihil ergo impedit, quominus statuamus $\zeta = \eta$, etsi enim hic valor campum aliquantillum diminuit; tamen is adhuc satis prodit notabilis. Tam autem fiet $r = 0$, sicque tertia lens quam minimam requireret aperturam, ita, vt simul officium diaphragmatis angustissimi praestet

X x 2

Coroll.

COROLL. 2.

252. Quodsi vero statuamus $\zeta = \eta$, sufficit capere \mathcal{C} intra limites 0 et 1. vnde simul \mathcal{C} fit positium. Tum vero capiatur

$$k = \frac{(\eta-1)(1+S)(\mathfrak{M}+\eta-1)}{2\mathfrak{M}\eta}$$

ex quo habebitur

$$P = \frac{2\mathfrak{M}\eta}{(1+S)(\mathfrak{M}+\eta-1)^2}$$

ita, vt pro maioribus multiplicationibus fit proxime

$$k = \frac{(\eta-1)(1+S)}{2\eta} \text{ et } P = \frac{2\eta}{1+S};$$

vnde patet, esse $P > 1$ ideoque A positium. Tum vero erit porro $\mathfrak{B} = -\frac{(P-1)}{\eta}$, ita, vt sit tam $\mathfrak{B} < 0$, quam $B < 0$. Denique hoc casu fiet

$$\mathfrak{D} = \frac{2\mathfrak{M} + 2S(\eta-1)}{S(\mathfrak{M}+\eta-1)} \text{ hincque } D = -\frac{2\mathfrak{M} - 2S(\eta-1)}{\mathfrak{M}(2-S) + S(\eta-1)}$$

$$\text{atque } M = \frac{2}{\mathfrak{M}+\eta-1}$$

Scholion.

253. Praeterquam quod hic casus $\zeta = \eta$ ad praxin imprimis est accommodatus, etiam hanc prerogativam complectitur, vt a littera \mathcal{C} reliqua elementa prorsus non pendeant, ita, vt quomocumque \mathcal{C} accipiamus intra limites scilicet 0 et 1, reliqua elementa nullam inde mutationem patiantur. Hoc autem modo facillime euitari poterit, ne lens obiectiva nimis fiat exigua, quod vero insuper commodius per litteram A praestatur, nisi forte de telescopiis sit sermo.

fermo, vbi $Aa = p$; superfluum igitur foret hic alios casus praeter istum $\zeta = \eta$ euoluere, atque nunc imprimis operae pretium erit aliquot valores pro η considerare, vt inde intelligere queamus, quinam inde ad praxin maxime futuri sint idonei. Pro littera vero S , quam vnitatem minorem esse debere vidimus, statuamus semper $S = \frac{1}{2}$, quia hinc satis idoneum intervallum inter binas lentes postremas oritur. Tum autem nostrae conditiones sequenti modo exprimentur:

$$1^{\circ}. k = \frac{2(\eta-1)(\eta+\eta-1)}{+\eta\eta}$$

$$2^{\circ}. P = \frac{+\eta\eta}{2(\eta+\eta-1)}$$

$$3^{\circ}. B = -\frac{(+\eta-1)\eta + 2(\eta-1)}{2\eta(\eta+\eta-1)}$$

4^o. C , vt iam notauimus, arbitrio nostro permittitur, dummodo sit intra 0 et 1.

$$5^{\circ}. D = \frac{+\eta + 2(\eta-1)}{\eta + \eta-1} \text{ et } D = -\frac{+\eta - 2(\eta-1)}{2\eta + \eta-1}$$

Hinc itaque distantiae focales erunt

$$p = Aa;$$

$$q = -\frac{A\eta}{P} \cdot a = \frac{(+\eta-1)\eta - 2(\eta-1)}{+\eta\eta} \cdot Aa;$$

$$r = -\frac{B\eta}{Pk} \cdot Aa = -\frac{B\eta}{\eta-1} \cdot Aa;$$

$$s = -\frac{BCD}{2\eta} \cdot Aa \text{ et } t = \frac{BCD}{\eta} \cdot Aa.$$

Deinde lentium intervalla

$$I^{num} = Aa \left(1 - \frac{1}{P}\right) = \frac{(+\eta-1)\eta - 2(\eta-1)}{+\eta\eta} \cdot Aa.$$

$$II^{dum} = -B \left(\frac{\eta(\eta-1) + 2\eta(\eta-1) + 2}{+\eta\eta(\eta-1)} \right) Aa.$$

$$\text{III}^{\text{tum}} = BC \left(\frac{1}{\eta-1} + \frac{1}{2M} \right) A a.$$

$$\text{IV}^{\text{tum}} = \frac{BCD}{2M} A a.$$

Deinde ob

$$M = \frac{2}{M+\eta-1} \text{ erit: } Q = \frac{1}{2} \left(1 + \frac{\eta-1}{2M} \right).$$

Pro prima lente femidiameter aperturae erit $= x$ ex aequatione notissima definiendus; pro secunda autem ob $\omega = \frac{2M}{M+\eta-1}$ erit femidiameter aperturae $= \frac{2M}{2(M+\eta-1)} q = \frac{2M}{P}$ et ob $r = 0$ femidiameter aperturae tertiae lentis $= \frac{2M}{Pk} = \frac{2M}{\eta-1}$; duabus autem reliquis lentibus apertura maxima tribuitur.

Exempl. I.

254. Sumamus: $\eta = 2$ eritque $k = \frac{3(M+1)}{2M}$ et quia tantum de maioribus multiplicationibus agitur, sumi poterit $k = \frac{3}{2}$ hinc $P = \frac{3M}{3(M+1)} = \frac{2}{3}$ proxime.

Porro

$$B = -\frac{5}{7} \text{ et } B = -\frac{5}{11};$$

$$D = 4 \text{ et } D = -\frac{4}{3};$$

vnde distantiae focales erunt

$$p = 2a; \quad q = \frac{5}{10} A a;$$

$$r = \frac{5}{11} C A a; \quad s = \frac{10}{11} \frac{C}{M} A a;$$

$$t = \frac{20}{11} \frac{C}{M} A a.$$

et lentium intervalla

$$1^{\text{um}} = \frac{5}{8} A a.$$

2^{um}

$$2^{dum} = \frac{s}{r}. A a.$$

$$3^{tium} = \frac{s}{r}. C (r + \frac{r}{m}) A a.$$

$$4^{tium} = \frac{10}{33} \frac{c^2}{m}. A a. \text{ et}$$

$$M = \frac{r^2}{m+1}; \text{ hinc } z = \frac{r}{a} \frac{a}{m+1} \text{ et}$$

$$O = \frac{1}{2} r (r + \frac{r}{m})$$

rum vero semidiameter aperturae lentis:

$$2^{dae} = \frac{q}{m+1} + \frac{s}{r} x \text{ et } 3^{tiae} = x.$$

Pro x autem inveniendō satisfiat huic aequationi:

$$\frac{x^3}{m^2 a^3} = \left\{ \begin{array}{l} \lambda + \nu A (r - A) \\ - \frac{(r-A)^2}{B^2 P} (\lambda' + \nu B (r - B)) \\ + \frac{(r-A)^2}{B^2 C^2 P k} (\lambda'' + \nu C (r - C)) \\ - \frac{(r-A)^2}{B^2 C^2 D^2 P k k'} (\lambda''' + \nu D (r - D)) \\ + \frac{(r-A)^2}{B^2 C^2 D^2} \cdot \lambda'''' \end{array} \right.$$

quae aequatio commodius ita repraesentabitur:

$$\frac{x^3}{m^2 a^3} = \left\{ \begin{array}{l} \frac{\lambda + \nu A (r - A)}{P^2} \\ - \frac{r}{P^2 q s} (\lambda' + \nu B (r - B)) \\ + \frac{r}{(P k)^2 r^3} (\lambda'' + \nu C (r - C)) \\ - \frac{r}{(P k k')^2 s^3} (\lambda''' + \nu D (r - D)) \\ + \frac{\lambda''''}{(P k k' s)^2 P} \end{array} \right.$$

hinc ergo inveniendū est

in A D

Coroll.

Coroll.

255. Si haec ad telescopia transferantur ponendo $\mathcal{A}a = Aa = p$ et $\mathfrak{M} = m$, fient distantiae focales

$$q = \frac{5}{16} \cdot p; \quad r = \frac{5}{11} \mathfrak{C} \cdot p;$$

$$s = \frac{10}{11} \cdot \frac{\mathfrak{C}}{\mathfrak{M}} \cdot p. \quad \text{et} \quad t = \frac{20}{11} \cdot \frac{\mathfrak{C}}{\mathfrak{M}} \cdot p.$$

interualla

$$1^{um} = \frac{5}{8} p;$$

$$2^{um} = \frac{5}{11} p;$$

$$3^{um} = \frac{5}{11} \mathfrak{C} \left(1 + \frac{1}{2\mathfrak{M}} \right) p \quad \text{et}$$

$$4^{um} = \frac{10}{11} \cdot \frac{\mathfrak{C}}{\mathfrak{M}} \cdot p.$$

et femidiameter campi $\Phi = \frac{1713}{\mathfrak{M} + 1}$ min. Longitudo ergo erit propemodum $= \frac{95}{88} p + \frac{5}{11} \mathfrak{C} \cdot p$. Nunc autem p ex aequatione ante data definiri poterit, ubi erit $\mathcal{A} = 0$ ob $a = \infty$.

Exemplum II.

256. Sit nunc $\eta = 3$ erit

$$k = \frac{\mathfrak{M} + 2}{2\mathfrak{M}}, \quad \text{fumi ergo poterit } k = \frac{1}{2}$$

$$P = \frac{4\mathfrak{M}}{\mathfrak{M} + 2} = 4$$

$$\mathfrak{B} = -\frac{3\mathfrak{M} + 2}{2(\mathfrak{M} + 2)} = -1; \quad \text{ergo } B = -\frac{2}{3}$$

$$\mathfrak{D} = \frac{4\mathfrak{M} + 4}{\mathfrak{M} + 2} = 4; \quad \text{ergo } D = -\frac{1}{4}.$$

Hinc ergo fient distantiae focales

$$p = \mathcal{A}a; \quad q = \frac{1}{2} Aa; \quad r = \frac{1}{2} \mathfrak{C} \cdot Aa;$$

$$r = \frac{C}{M} \cdot A \cdot a. \text{ et } t = \frac{2}{3} \cdot \frac{C}{M} \cdot A \cdot a.$$

et lentium interualla

$$1^{mum} = \frac{3}{4} \cdot A \cdot a.$$

$$2^{dum} = \frac{3}{8} A \cdot a.$$

$$3^{tium} = \frac{1}{4} C \left(1 + \frac{1}{M} \right) A \cdot a.$$

$$4^{tum} = \frac{1}{3} \cdot \frac{C}{M} \cdot A \cdot a.$$

hinc porro erit

$$M = \frac{2}{M+2} \text{ hinc } z = \frac{1}{2} \cdot \frac{2}{M+2} \text{ et}$$

$$O = \frac{1}{2} t \left(1 + \frac{2}{M} \right)$$

tum vero semidiameter aperturae lentis

$$2^{dae} = \frac{3q}{2(M+2)} + \frac{1}{4} x; \text{ et } 3^{tiae} = \frac{1}{2} \cdot x.$$

Cetera se habent, vt ante.

Coroll.

257. Facta applicatione ad telescopia, vbi fit $A a = p$ omnia elementa facile determinantur, vt ante; tum vero longitudo instrumenti omiſſis partibus per M diuisis erit $= \frac{2}{3} p + \frac{1}{4} C \cdot p$, quae minor est, quam casu praecedente.

Exemplum III.

258. Statuamus $\eta = 6$, vt fiat $M = \frac{2}{M+5}$ eritque

$$k = \frac{5(M+5)}{8M} = \frac{5}{8} \text{ proxime.}$$

Tom. III.

Y y

P=8.

$$P = 8. \quad \mathfrak{B} = \frac{-7\mathfrak{M} + 5}{6(\mathfrak{M} + 5)} = -\frac{7}{6} \text{ proxime}$$

vnde fit

$$B = -\frac{7}{18}; \quad \mathfrak{D} = 4 \text{ et } D = -\frac{4}{3}.$$

vnde distantiae focales prodibunt

$$p = 2a; \quad q = \frac{7}{48} \cdot Aa; \quad r = \frac{7}{35} \cdot C \cdot Aa;$$

$$s = \frac{14}{18} \cdot \frac{C}{\mathfrak{M}} \cdot Aa; \quad t = \frac{28}{35} \cdot \frac{C}{\mathfrak{M}} \cdot Aa.$$

et lentium interualla

$$1^{um} = \frac{7}{8} \cdot Aa.$$

$$2^{dum} = \frac{7}{48} \cdot Aa.$$

$$3^{tiam} = \frac{7}{18} \cdot C \cdot \left(\frac{1}{5} + \frac{1}{2\mathfrak{M}} \right) \cdot Aa.$$

$$4^{tum} = \frac{14}{35} \cdot \frac{C}{\mathfrak{M}} \cdot Aa.$$

Praeterea

$$z = \frac{1}{2} \cdot \frac{a}{\mathfrak{M} + 5} \text{ et } O = \frac{1}{2} t \left(1 + \frac{5}{\mathfrak{M}} \right).$$

tum vero semidiameter aperturae lentis

$$2^{dae} = \frac{3q}{\mathfrak{M} + 5} + \frac{1}{5} x \text{ et } 3^{tiae} = \frac{1}{5} x.$$

Coroll.

259. Translatione igitur ad telescopia facta prodiret hoc casu eorum longitudo $= \frac{2}{18} p + \frac{7}{35} C \cdot p$. quae longitudo satis est exigua, vt etiam in aliis generibus vix minor sperari queat.

Scho-

Scholion.

260. Etsi iste casus $\zeta = \eta$ in praxi summum vsum praestare videtur, tamen etiam considerari conueniet quempiam casum, quo $\zeta > \eta$, quandoquidem hoc modo campo quodpiam augmentum adferatur. Manente autem $S = \frac{1}{2}$, alter limes pro ζ erat $\zeta < \eta + \frac{3(\eta-1)}{4}$ siue $\zeta < \frac{7}{4}\eta - \frac{3}{4}$. Huic autem limiti ipsi aequari nequit, quia alioquin \mathcal{C} deberet esse $= 1$ hincque $C = \infty$. Sumamus igitur

$$\zeta = \eta + \frac{2}{4}(\eta - 1) = \frac{3}{2}\eta - \frac{1}{2}$$

ac reperietur $\mathcal{C} > \frac{2}{3}$ et $\mathcal{C} < 1$. Sumatur igitur $\mathcal{C} = \frac{2}{3}$ vt fiat $C = 3$. hincque fiet

$$k = \frac{(\eta-1)(-2M+15(\eta-1))}{12M(3\eta-1)} \text{ et } P = \frac{12M(3\eta-1)}{2M+15(\eta-1)}$$

hincque porro

$$\mathfrak{B} = -\frac{4M(18\eta-7)+30(\eta-1)}{(3\eta-1)(2M+15(\eta-1))} \text{ et}$$

$$\mathfrak{D} = \frac{24M+10(\eta-1)}{6M+5(\eta-1)} \text{ seu proxime } \mathfrak{D} = 4.$$

Tum vero prodibit

$$M = \frac{12}{6M+5(\eta-1)} = \frac{2}{M + \frac{5}{6}(\eta-1)},$$

cum antea fuisset $M = \frac{2}{M+\eta-1}$. Quodsi iam sumamus, vt in exemplo 2^{do}, $\eta = 3$ fient haec elementa

$$k = \frac{M+15}{24M} \text{ et } P = \frac{48M}{M+15} \text{ hinc}$$

$$\mathfrak{B} = \frac{-47M+15}{4(M+15)}, \quad \mathfrak{B} = \frac{-47M+15}{51M+45}$$

Y y 2

tum

tum.

$$G = \frac{3}{4}; C = 3; D = 4; \text{ et } D = -\frac{4}{3};$$

vnde singula momenta pro constructione definiiri possunt. Quia vero hic tanti occurrunt numeri, quos prae M negligere non amplius licet; in adplicatione ad exempla statim quoque pro M determinatum valorem assumi conueniet. Praeterea vero hic ad specialiora non progredimur, quia adhuc lente obiectiua simplice utimur, ita, ut confusio aliter tolli nequeat, nisi aperturam lentis obiectiuae diminuendo; quod remedium cum praxis sponte offerat, non opus erit, litteram x . molesto illo calculo definire, si quis enim microscopium secundum huiusmodi mensuras construxerit, ipse vsus aperturam declarabit; quando autem in sequente capite per multiplicationem lentis obiectiuae omnem confusionem ad nihilum redigemus, tum demum necesse erit completas determinationes pro singulis momentis, uti haecenus fecimus, exhibere.

Problema 2.

261. Microscopium huius generis ex quinque lentibus construere, quae ita sint dispositae, ut prior imago realis in interuallum secundum, posterior vero in interuallum quartum incidat.

Solutio.

Quatuor ergo litterarum P, Q, R, S secunda et quarta erunt negatiuae; vnde ponatur $Q = -k$ et $S =$

$S = -k'$, vt fit: $P k R k' = \frac{ma}{b} = M$, hinc erit vlti-
mae lentis distantia focalis

$$t = \frac{ABCD}{P k R k'} a = \frac{ABCD}{M} a,$$

quae debet esse positua aequae ac lentium interualla,
quae sunt:

$$1^{um} = A a \left(1 - \frac{1}{P}\right);$$

$$2^{um} = -\frac{AB'}{P} a \left(1 + \frac{1}{k}\right);$$

$$3^{um} = -\frac{ABC}{Pk} a \left(1 - \frac{1}{R}\right);$$

$$4^{um} = +\frac{ABCD}{P k R} a \left(1 + \frac{1}{k'}\right);$$

ergo vt tam vltima lens, quam vltimum interual-
lum, fiant positua, debet esse $ABCD > 0$. Hinc
vt tertium quoque fiat posituum, debet esse

$$-D \left(1 - \frac{1}{R}\right) > 0.$$

ficque circa D nihil definitur. Ob secundum autem
interuallum debet esse $-AB > 0$ et ob primum
 $A \left(1 - \frac{1}{P}\right) > 0$. Tum ergo debebit esse $CD < 0$.
Statuatur nunc $M = \frac{a+r+s+t}{M-i}$ et quia campus maxi-
mus desideratur, statim sumi poterit $s = 1$ et $t = 1$,
vt fiat $M = \frac{a+r+1+1}{M-i}$ hincque distantia oculi $O = \frac{t}{M M'}$
quae cum fit positua, destructio marginis praebet:

$$O = \frac{a}{P} - \frac{r}{Pk} - \frac{1}{PkR} + \frac{1}{P k R k'}$$

unde colligitur

$$\frac{1}{k'} = -a k R + r R + 1$$

Y y 3

hinc

hinc erit

$$P k R = M (1 + r R - q k R)$$

ficque patet esse $1 + r R > q k R$. Praeterea vero considerare debemus sequentes aequationes:

$$1. \mathfrak{B} q = (P - 1) M$$

$$2. \mathfrak{C} r = -(1 + P k) M - q$$

$$3. \mathfrak{D} = -(1 + P k R) M - q - r$$

Ponatur hic, vt ante, breuitatis gratia

$$\frac{1-P}{\mathfrak{B}} = \zeta \text{ et } 1 + P k = \eta, \text{ fietque}$$

$$q = -\zeta M \text{ et } r = \frac{(\zeta - \eta)M}{\mathfrak{C}}$$

unde colligitur

$$2 + q + r = \frac{2\mathfrak{C} + (\zeta(1-\mathfrak{C}) - \eta)M}{\mathfrak{C}} = (M - 1) M;$$

ex qua aequatione deducitur

$$M = \frac{2\mathfrak{C}}{\mathfrak{C}(M-1) - \zeta(1-\mathfrak{C}) + \eta} = \frac{2\mathfrak{C}}{\mathfrak{C}(M+\zeta-1) - \zeta + \eta}$$

ex quo valore viciffim erit

$$q = -\frac{2\zeta\mathfrak{C}}{\mathfrak{C}(M+\zeta-1) - \zeta + \eta} \text{ et}$$

$$r = \frac{2(\zeta - \eta)}{\mathfrak{C}(M+\zeta-1) - \zeta + \eta}$$

Nunc vt marginis colorati rationem habeamus, erit statim

$$1 + r R - q k R = \frac{\mathfrak{C}(M+\zeta-1) - \zeta + \eta + 2(\zeta - \eta)R + 2\zeta\mathfrak{C}kR}{\mathfrak{C}(M+\zeta-1) - \zeta + \eta}$$

Et cum ob

$$P k = \eta - 1 \text{ fit } P k R = (\eta - 1) R,$$

erit

erit

$$\begin{aligned} & \mathfrak{C}(\eta-1)R(\mathfrak{M}+\zeta-1) - (\eta-1)(\zeta-\eta)R - \mathfrak{M}\mathfrak{C}(\mathfrak{M}+\zeta-1) \\ & + \mathfrak{M}(\zeta-\eta) - 2\mathfrak{M}(\zeta-\eta)R \\ & - 2\mathfrak{M}\zeta\mathfrak{C}kR = 0. \end{aligned}$$

Ante quam autem hinc vel k vel R determinemus, considerare debemus rationem litterae \mathfrak{D} ex superiori tertia aequatione; cum igitur PkR sit sine dubio numerus magnus inuoluens \mathfrak{M} , facile intelligitur litteram \mathfrak{D} esse negatiuam; unde etiam erit \mathfrak{D} negatiuum, adeoque concludimus fore $\mathfrak{C} > 0$, hincque $\mathfrak{C} < 1$. Ob eandem vero rationem debet esse $R > 1$; ita, ut haec littera aliquatenus tanquam nota spectari possit; quare ex illa aequatione colligimus

$$k = \left\{ \begin{array}{l} \mathfrak{C}(\eta-1)R(\mathfrak{M}+\zeta-1) - (\eta-1)(\zeta-\eta)R \\ - \mathfrak{M}\mathfrak{C}(\mathfrak{M}+\zeta-1) + \mathfrak{M}(\zeta-\eta) \\ - 2\mathfrak{M}(\zeta-\eta)R \end{array} \right\} : 2\mathfrak{M}\zeta\mathfrak{C}R$$

hincque $P = \frac{(\eta-1)}{k}$; ita, ut sit $\eta > 1$. Quare ut valor ipsius k fiat positius, debet esse

$$\begin{aligned} & R(\mathfrak{C}(\eta-1)(\mathfrak{M}+\zeta-1) - (\eta-1)(\zeta-\eta) - 2\mathfrak{M}(\zeta-\eta)) \\ & > \mathfrak{M}(\mathfrak{C}\mathfrak{M} + \mathfrak{C}(\zeta-1) - \zeta + \eta) \end{aligned}$$

ad quod primo requiritur, ut quantitas litteram R multiplicans sit positua, et quia \mathfrak{M} est numerus praemagnus, ipsius coefferens ante omnia debet esse positius, unde colligimus $\mathfrak{C}(\eta-1) > 2(\zeta-\eta)$ unde
con-

concluditur $\mathcal{E} > \frac{2(\zeta - \eta)}{\eta - 1}$; quia igitur $\mathcal{E} < 1$, erit

$$2(\zeta - \eta) < \eta - 1; \text{ sicque } \zeta < \frac{\eta - 1}{2}.$$

Qua conditione impleta debet esse

$$R > \frac{\mathfrak{M}(\mathcal{E}\mathfrak{M} + \mathcal{E}(\zeta - 1) - \zeta + \eta)}{\mathcal{E}(\eta - 1)(\mathfrak{M} + \zeta - 1) - 2\mathfrak{M}(\zeta - \eta) - (\eta - 1)(\zeta - \eta)}$$

atque hinc retrogrediendo omnia elementa determinabuntur. Reliqua vero expediuntur, ut in praecedente problemate.

COROLL.

262. Hic igitur littera R denotabit numerum magnum \mathfrak{M} involuentem; deinde conditio $\zeta < \frac{\eta - 1}{2}$ instituto nostro maxime fauet, cum campi conditio imprimis postulet, ne ζ ultra necessitatem augeatur. Quare cum semper sit $\eta > 1$; commodissime videtur statui posse $\zeta = \eta$, uti in praecedente problemate; ita, ut tertiae lentis apertura iterum fiat minima prodeatque

$$M = \frac{2\mathcal{E}}{\mathcal{E}\mathfrak{M} + \mathcal{E}\eta - \mathcal{E}} = \frac{2}{\mathfrak{M} + \eta - 1}.$$

COROLL. 2.

263. Sumto autem $\zeta = \eta$, pro \mathcal{E} limites erunt $\mathcal{E} < 1$ et $\mathcal{E} > 0$. Porro capi debet $R > \frac{\mathfrak{M}}{\eta - 1}$; indeque fiet

$$k = \frac{((\eta - 1)R - \mathfrak{M})(\mathfrak{M} + \eta - 1)}{2\mathfrak{M}\eta R} \text{ et } P = \frac{2\mathfrak{M}\eta(\eta - 1)R}{((\eta - 1)R - \mathfrak{M})(\mathfrak{M} + \eta - 1)}$$

Prae-

Praeterea vero erit

$$\mathfrak{B} = - \frac{(\eta-1)R((2\eta-1)\mathfrak{M}-\eta+1) - \mathfrak{M}(\mathfrak{M}+\eta-1)}{\eta((\eta-1)R - \mathfrak{M})(\mathfrak{M}+\eta-1)}$$

Denique vero reperietur

$$\mathfrak{D} = - \frac{2(1+(\eta-1)R-\eta)}{\mathfrak{M}+\eta-1}$$

sive cum \mathfrak{M} et R sint numeri praemagni, erit proxime $\mathfrak{D} = - \frac{2(\eta-1)R}{\mathfrak{M}}$, qui valor certo est negativus, ut supra iam posuimus.

Coroll. 3.

264. Quin etiam statui poterit $\zeta = 0$; unde pro \mathfrak{E} limites erunt $\mathfrak{E} < 1$ et $\mathfrak{E} > - \frac{2\eta}{\eta-1}$; cui satisfit, dummodo \mathfrak{E} intra limites 1 et 0 contineatur. Tum vero sumi debet

$$R > \frac{\mathfrak{M}(\mathfrak{E}\mathfrak{M} - \mathfrak{E} + \eta)}{\mathfrak{E}(\eta-1)(\mathfrak{M}-1) + \mathfrak{M}\eta + \eta(\eta-1)}$$

sive ob \mathfrak{M} numerum praemagnum $R > \frac{\mathfrak{E}\mathfrak{M}}{\mathfrak{E}(\eta-1) + 2\eta}$

Verum hinc sequitur porro $k = \infty$ et $P = 0$. ita, ut fit $Pk = \eta - 1$. Praeterea vero prodit $\mathfrak{B} = \infty$ et $B = -1$. Denique vero ob

$$q = 0 \text{ et } r = - \frac{\eta\mathfrak{M}}{\mathfrak{E}} \text{ erit}$$

$$\mathfrak{D} = - (1 + (\eta-1)R - \frac{\eta}{\mathfrak{E}}) \mathfrak{M}$$

ideoque ob $M = \frac{\epsilon}{\epsilon(M-1)+\eta}$, qui valor ipsius M aliquanto minor est, quam casu praecedente, fiet $\mathfrak{D} = -\frac{2(\eta-1)R}{\mathfrak{M}}$.

Scholion.

265. Quantumvis hic casus paradoxus videatur, cum ob $\mathfrak{B} = \infty$ tum vero ob $P = 0$, tamen reuera est realis et ad casum in praecedente capite exposi- tum reducitur, quia enim $\mathfrak{B} = \infty$; secundae lentis di- stantia focalis est infinita, sicque res eodem redit ac si secunda lens plane abesset, ita, vt non amplius quae- stio sit de eius loco, quare etsi primum interuallum prodeat $= A a (1 - \frac{1}{P}) = -\infty$, et secundum

$$A a (\frac{1}{P} + \frac{1}{\eta-1}) = +\infty,$$

tamen horum summa, quae sola nunc est spectanda, sit finita

$$= A a (1 + \frac{1}{\eta-1}) = \frac{\eta}{\eta-1} \cdot A a.$$

Cum igitur tantum quatuor lentes hic habeantur, hic casus ad praecedens caput est referendus. Interim tamen hinc incommodum nasci debet, quando ζ prope ad 0 accedit, quia tum P etiam erit vnitae minus, ita, vt A debeat esse numerus negatiuus et $B > 0$. Cum autem sit $\mathfrak{B} = \frac{1-P}{\zeta}$, erit quidem $\mathfrak{B} > 0$, verum insuper necesse est, vt sit $1-P < \zeta$, vel $P > 1-\zeta$, siue P contineri debet intra limites 1 et $1-\zeta$ seu esse debet $k < \frac{\eta-1}{1-\zeta}$, quare cum \mathfrak{M} et

et R sint numeri praemagni, debeat esse

$$R \left(\frac{\mathbb{C}(\eta-1)(1-\zeta)}{1-\zeta} - 2 \right) < \mathbb{C} \mathbb{M}$$

quod sponte euenit, si fuerit

$$\frac{\mathbb{C}(\eta-1)(1-\zeta)}{1-\zeta} < 2 \text{ siue } \mathbb{C} < \frac{2(1-\zeta)}{(\eta-1)(1-\zeta)}$$

Sin autem sit

$$\frac{\mathbb{C}(\eta-1)(1-\zeta)}{1-\zeta} > 2 \text{ debet esse } R < \frac{(1-\zeta)\mathbb{C}\mathbb{M}}{\mathbb{C}(\eta-1)(1-\zeta) - 2(1-\zeta)}$$

quibus obseruatis aliquot casus fusius euoluamus.

C A S V S I.

quo $\zeta = \eta$.

266. Hoc casu iam vidimus, lentem tertiam nostro arbitrio relinqui, dummodo pro ea capiatur $\mathbb{C} < 1$ et $\mathbb{C} > 0$, vt C fiat numerus positius, vnde si circumstantiae postulent, vt C sit numerus satis magnus, tum \mathbb{C} parum ab vnitade deficere debeat; deinde etiam notauimus, capi debere $R > \frac{\mathbb{M}}{\eta-1}$, vnde cum semper sit $\eta > 1$, si etiam fuerit > 2 , tunc commode sumi poterit $R = \mathbb{M}$. Notetur autem, litteram η non nimis magnam sumi conuenire, quia pro campo fit $M = \frac{2}{\mathbb{M} + \eta - 1}$. Deinde vero prodit

$$k = \frac{((\eta-1)R - \mathbb{M})(\mathbb{M} + \eta - 1)}{2\mathbb{M}\eta R}$$

quare pro maioribus multiplicationibus tuto sumi poterit $k = \frac{(\eta-1)R - \mathbb{M}}{2\eta R}$ vnde patet, litteram k eo fore minorem, quo minus R limitem praescriptum $\frac{\mathbb{M}}{\eta-1}$

Z z 2

supe-

superet; vnde fit

$$P = \frac{\eta - 1}{k} = \frac{2\eta(\eta - 1)k}{(\eta - 1)R - \mathfrak{M}}$$

Pro reliquis elementis primo prodit

$$\mathfrak{B} = - \frac{(\eta - 1)R((2\eta - 1)\mathfrak{M} - \eta + 1) - \mathfrak{M}(\mathfrak{M} + \eta - 1)}{\eta((\eta - 1)R - \mathfrak{M})(\mathfrak{M} + \eta - 1)}$$

hincque proxime

$$\mathfrak{B} = - \frac{(\eta - 1)(2\eta - 1)R - \mathfrak{M}}{\eta(\eta - 1)R - \eta\mathfrak{M}}$$

qui valor manifesto est negatiuus, ex quo etiam B erit negatiuum. Deinde cum fit $P > 1$; ob eandem rationem littera A debet esse positiua; ex quo productum AB ob

$$B = - \frac{(\eta - 1)(2\eta - 1)R - \mathfrak{M}}{(\eta - 1)(2\eta - 1)R - (\eta - 1)\mathfrak{M}}$$

sponte fit negatiuum, prorsus, vti conditiones praescriptae postulant. Denique vero reperitur

$$\mathfrak{D} = - \frac{2(1 + (\eta - 1)R - \eta)}{\mathfrak{M} + \eta - 1}$$

ideoque proxime

$$\mathfrak{D} = - \frac{2(\eta - 1)R}{\mathfrak{M}}; \text{ vnde fit } D = - \frac{2(\eta - 1)R}{\mathfrak{M} + 2(\eta - 1)R}$$

His definitis erunt primo distantiae focales

$$p = \mathfrak{A} a, \quad q = - \frac{\mathfrak{A}\mathfrak{B}}{P} a = \frac{(\eta - 1)(2\eta - 1)R + \mathfrak{M}}{2\eta^2(\eta - 1)R} \cdot \mathfrak{A} a,$$

$$r = - \frac{\mathfrak{A}\mathfrak{C}}{\eta - 1} \cdot a; \quad s = + \frac{\mathfrak{A}\mathfrak{B}\mathfrak{C}\mathfrak{D}}{(\eta - 1)R} \cdot a,$$

$$\text{et } t = \frac{\mathfrak{A}\mathfrak{B}\mathfrak{C}\mathfrak{D}}{\mathfrak{M}} \cdot a.$$

Deinde

Deinde vero interualla

$$I^{mum} = A a \left(1 - \frac{1}{P} \right);$$

$$II^{dum} = - A B. a \left(\frac{1}{P} + \frac{1}{\eta - 1} \right);$$

$$III^{tium} = - \frac{ABC a}{\eta - 1} \left(1 - \frac{1}{R} \right);$$

$$IV^{tum} = A B C D. a \left(\frac{1}{(\eta - 1)R} + \frac{1}{M} \right).$$

Distantia vero oculi fiet

$$O = \frac{t}{M \eta} = \frac{t}{2} \cdot \frac{M + \eta - 1}{M} = \frac{1}{2} t \left(1 + \frac{\eta - 1}{M} \right) = \frac{1}{2} t \text{ proxime.}$$

Pro aperturis vero lentium primae quidem apertura tutissime per experientiam definitur; vnde reperitur littera x , ex eaque mensura claritatis $= \frac{20 x}{M}$, si scilicet x in digitis exprimatur.

Pro secunda vero lente cum sit

$$q = - \eta M = - \frac{2 \eta}{M + \eta - 1} = - \frac{2 \eta}{M},$$

erit eius aperturæ semidiameter

$$= \frac{1}{2} q q + \frac{x}{P} = - \frac{1}{2} \cdot \frac{\eta}{M} \cdot q + \frac{x}{P}.$$

Pro tertia vero lente ob $r = 0$ sufficit aperturæ semidiameter $= \frac{x}{\eta - 1}$; reliquas vero lentes vtrinque aequæ conexas confici conuenit.

Coroll.

267. Si sumatur $R = \frac{M}{\eta - 1}$, vt fiat $k = 0$ et $P = \infty$, existente $P k = \eta - 1$; tum fiet $B = - \frac{P}{\eta} = \infty$

Z z 3

et

et $B = -1$. Tum igitur secunda lens in ipsam imaginem realem priorem incidet ob primum intervallum $= Aa = a$, eiusque distantia focalis erit $q = \frac{Aa}{\eta}$. At vero secundum intervallum hoc casu euadet $= -\frac{ABa}{\eta-1} = \frac{Aa}{\eta-1}$. Reliqua vero definiuntur, ut in genere, si modo notetur, fore $\mathfrak{D} = -2$. et $D = -\frac{2}{3}$.

Exempl. I.

Ponamus $\eta = 2$ et capi debet $R > \mathfrak{M}$. Nihil vero impedit, quominus secundum praecedens Coroll. sumamus $R = \mathfrak{M}$ ita, ut tum fiat $k = 0$ et $P = \infty$ quare primo distantiae focales ita exprimentur

$$p = 2a; \quad q = \frac{Aa}{2};$$

$$r = AC \cdot a; \quad s = \frac{2AC}{\mathfrak{M}} \cdot a;$$

$$\text{et } t = \frac{2}{3} \cdot \frac{C}{\mathfrak{M}} \cdot Aa = \frac{1}{3} s.$$

Deinde intervalla ita se habebunt:

$$1^{um} = Aa; \quad 2^{um} = Aa;$$

$$3^{um} = C \cdot (1 - \frac{1}{\mathfrak{M}}) Aa;$$

$$4^{um} = \frac{4}{3} \cdot \frac{C}{\mathfrak{M}} \cdot Aa.$$

distantia vero oculi $= \frac{1}{2} t (1 + \frac{1}{\mathfrak{M}})$.

Pro valore ipsius x siue per experientiam siue per formulam notam definito, fiat secundae lentis semidiameter aperturæ $= \frac{q}{\mathfrak{M}} = \frac{1}{2} \cdot \frac{Aa}{\mathfrak{M}}$ et tertiae lentis $= x$.

$= x$. semidiameter spatii conspicui erit $z = \frac{a}{2M}$ et
 mensura claritatis $= \frac{20 \cdot x}{M}$.

Corollarium.

268. Hae formulae quoque ad telescopia accommodari poterunt, sumendo $A a = p$ et $M = m$.
 Tum vero sumi debet ipsa distantia focalis

$$p = m \sqrt{\mu m (\lambda * + \frac{\lambda''}{C^3} + \frac{v}{C^2})}$$

omissis terminis sequentibus per M diuisis.

Exempl. II.

269. Sit nunc $\eta = 3$, vt esse debeat $R > \frac{20}{3}$
 sumaturque $R = M$, vnde fiet $k = \frac{1}{3}$ et $P = 12$;
 quare reliqua elementa fient $B = -\frac{11}{3}$; hincque
 $B = -\frac{11}{3}$ et $D = -4$ et $D = -\frac{4}{3}$ vnde distantiae
 focales euadent

$$p = 2a; q = \frac{11}{36} A a.$$

$$r = \frac{11}{24} C. A a. s = \frac{11}{7} \cdot \frac{C}{M}. A a. \text{ et}$$

$$t = \frac{22}{33} C. A a. = \frac{2}{3} \cdot s$$

atque interualla lentium

$$1^{um} = \frac{11}{12} A a;$$

$$2^{um} = \frac{11}{24} A a;$$

$$3^{tum} = \frac{11}{24} C (1 - \frac{1}{M}) A a.$$

$$4^{tum} = \frac{11}{25} \cdot \frac{C}{M} \cdot A a.$$

oculi-

oculique distantia $O = \frac{1}{2} t (1 + \frac{2}{\mathfrak{M}})$, spatii vero in
 obiecto conspicui erit semidiameter $x = \frac{a}{2(\mathfrak{M} + 2)}$. De-
 finito x siue per experientiam siue per formulam
 notam erit semidiameter aperturæ

$$\text{secundæ lentis} = \frac{2}{3} \cdot \frac{q}{\mathfrak{M}} + \frac{\infty}{12} \text{ et tertiæ} = \frac{\infty}{2}.$$

$$\text{et gradus claritatis} = \frac{20 \cdot \infty}{\mathfrak{M}}.$$

Coroll.

270. Hæ formulæ etiam ad telescopia accom-
 modari possunt; erit enim $A a = p$ et $\mathfrak{M} = m$. Tum
 vero lentis obiectivæ distantia focalis definitur per
 hanc formulam

$$p = m \sqrt[3]{\mu m (\lambda + \frac{1}{12} (\frac{3}{11^2} \lambda' - \frac{3 \cdot 14 \nu}{11^2}) + \frac{14^3}{11^2 \cdot 2} (\frac{\lambda''}{\mathfrak{C}^3} + \frac{\nu}{\mathfrak{C}\mathfrak{E}}))}$$

$$\text{Longitudo huius telescpii erit circiter} = (\frac{11}{9} + \frac{11}{28} \mathfrak{C}) p.$$

Exempl. III.

271. Sit nunc $\eta = 6$, vt esse debeat $R > \frac{300}{5}$
 et sumatur $R = \frac{300}{2}$ ac reperitur $k = \frac{1}{2}$ et $P = 20$.
 Hinc porro fiet $\mathfrak{B} = -\frac{10}{9}$ et $B = -\frac{10}{12}$. Tum vero
 $\mathfrak{D} = -5$ et $D = -\frac{5}{2}$. Distantiæ ergo focales ita
 se habebunt:

$$p = 21 a; q = \frac{10}{120} A a;$$

$$r = \frac{30}{125} \mathfrak{C} A a; s = \frac{30}{25} \frac{\mathfrak{C}}{\mathfrak{M}} A a, \text{ et}$$

$$t = \frac{10}{10} \cdot \frac{\mathfrak{C}}{\mathfrak{M}} A a \text{ seu } t = \frac{5}{12} \cdot s.$$

et

et interualla lentium

$$1^{mum} = \frac{10}{20} A a;$$

$$2^{dum} = \frac{10}{100} A a;$$

$$3^{tum} = \frac{11}{128} C \left(1 - \frac{2}{21} \right) A a. \text{ et}$$

$$4^{tum} = \frac{133}{130} \cdot \frac{C}{21} \cdot A a.$$

localique distantia $O = \frac{1}{2} t \left(1 + \frac{5}{21} \right)$.

Spatii conspicui semidiameter $x = \frac{a}{2(21+5)}$.

Definito denique x , vt ante, erit semidiameter aperturæ lentis

$$\text{secundæ} = \frac{3}{21} q + \frac{30}{20} \text{ et tertiæ} = \frac{1}{3} x;$$

$$\text{gradus autem claritatis manet} = 20 \cdot \frac{30}{21}.$$

Exempl. IV.

272. Sit vt ante $\eta = 6$, sumatur vero $R = 21$, ac reperitur $k = \frac{1}{3}$ et $P = 15$. Hinc porro fit

$$\mathfrak{B} = -\frac{7}{3} \text{ et } \mathfrak{B} = -\frac{7}{10}, \text{ at}$$

$$\mathfrak{D} = -10. \text{ et } \mathfrak{D} = -\frac{10}{11}.$$

Distantiæ ergo focales erunt

$$p = 21 a; q = \frac{7}{15} A a;$$

$$r = \frac{7}{30} C \cdot A a; s = \frac{7}{3} \cdot \frac{C}{21} \cdot A a;$$

$$t = \frac{7}{11} \cdot \frac{C}{21} \cdot A a \text{ seu } t = \frac{5}{11} s.$$

Tom. III.

A a a

et

et lentium interualla

$$1^{um} = \frac{14}{15} \cdot A a;$$

$$2^{um} = \frac{14}{75} \cdot A a;$$

$$3^{um} = \frac{7}{50} \cdot C \cdot \left(1 - \frac{1}{50}\right) A a;$$

$$4^{um} = \frac{42}{55} \cdot \frac{C}{M} \cdot A a.$$

et distantia Oculi $= \frac{1}{2} t \left(1 + \frac{5}{50}\right)$ pariter, ac reliqua momenta, se habet, vt ante.

Corollarium.

273. Si haec ad telescopia transferantur, ponendo $A a = p$ et $M = m$, casus hic posterior antecedenti praesferendus videtur, siquidem praebet longitudinem parumper minorem, quippe quae neglectis terminis per M diuisis erit $= \left(1 \frac{5}{50} + \frac{7}{50} C\right) p$ cum ex exemplo praecedente fuerit $= \left(1 \frac{7}{50} + \frac{12}{125} C\right) p$ Casu ergo ultimi exempli lentis obiectiuae distantia focalis ita definietur, vt fit

$$p = m \sqrt{\mu} m \left(\lambda + \frac{1}{15} \left(\frac{55}{75} \lambda' - \frac{30 \cdot \nu}{7^2} \right) - \frac{200}{545} \left(\frac{\lambda''}{65} + \frac{\nu}{65} \right) \right)$$

CASVS II.

quo $\zeta = 1$.

274. Cum fit $\zeta = 1$, limites pro C erunt $C < 1$ et $C > -2$, ita, vt C aequae arbitrio nostro permittatur, ac ante. Tum vero esse debet

$$R > \frac{m(Cm + \eta - 1)}{(\eta - 1)(Cm + 2M + \eta - 1)}$$

feru

feu neglectis minoribus partibus $R > \frac{\mathcal{E}\mathcal{M}}{(\eta-1)(2+\mathcal{E})}$.

Statuatur ergo $R = i\mathcal{M}$, fumendo scilicet $i > \frac{\mathcal{E}}{(\eta-1)(2+\mathcal{E})}$

Tum ergo fiet

$$k = \frac{\mathcal{E}(\eta i - 1) + 2i(\eta - 1)}{2i\mathcal{E}} \text{ et } P = \frac{2i(\eta - 1)\mathcal{E}}{\mathcal{E}((\eta - 1)i - 1) + 2i(\eta - 1)}$$

Hinc ergo fiet

$$P - 1 = \frac{\mathcal{E}(i(\eta - 1) + 1) - 2i(\eta - 1)}{\mathcal{E}((\eta - 1)i - 1) + 2i(\eta - 1)}$$

qui valor erit positivus, seu $P > 1$, si fuerit

$$i < \frac{\mathcal{E}}{(\eta - 1)(2 - \mathcal{E})}$$

Hoc ergo casu prodit

$$\mathcal{B} = \frac{-\mathcal{E}(i(\eta - 1) + 1) + 2i(\eta - 1)}{\mathcal{E}((\eta - 1)i - 1) + 2i(\eta - 1)}$$

qui valor cum sit negativus, etiam \mathcal{B} erit negativum et ob $P > 1$ debet \mathcal{A} esse positivum, ut superiores conditiones postulant; sin autem esset

$$i > \frac{\mathcal{E}}{(\eta - 1)(2 - \mathcal{E})};$$

tum foret $P < 1$ sumique deberet \mathcal{A} negativum, ac proderet $\mathcal{B} > 0$. unde ut etiam \mathcal{B} fiat positivum, insuper necesse est, ut sit $\mathcal{B} < 1$, quod manifestum est, cum sit $\mathcal{B} = 1 - P$. Postea vero pro \mathcal{D} inveniendone notetur esse

$$M = \frac{2\mathcal{E}}{\mathcal{E}\mathcal{M} + \eta - 1} \text{ et}$$

$$q = -M \text{ et } r = -\frac{(\eta - 1)M}{\mathcal{E}};$$

ex quibus prodit

$$\mathcal{D} = -i(\eta - 1)M \quad \mathcal{M} = -2i(\eta - 1).$$

illi vero valores abibunt in hos:

$$q = -\frac{z}{\mathfrak{M}}; \text{ et } r = -\frac{z(\eta-1)}{\mathfrak{C}\mathfrak{M}}.$$

His valoribus inuentis, considerentur primo distantiae focales, quae sunt:

$$p = \mathfrak{A}a; \quad q = -\frac{\mathfrak{A}\mathfrak{B}}{\mathfrak{P}} \cdot a;$$

$$r = -\frac{\mathfrak{A}\mathfrak{B}\mathfrak{C}}{\eta-1} \cdot a; \quad s = \frac{\mathfrak{A}\mathfrak{B}\mathfrak{C}\mathfrak{D}}{z(\eta-1)\mathfrak{M}} \cdot a;$$

$$\text{et } t = \frac{\mathfrak{A}\mathfrak{B}\mathfrak{C}\mathfrak{D}}{\mathfrak{M}} \cdot a.$$

Tum vero interualla:

$$1^{um} = \mathfrak{A}a \left(1 - \frac{1}{\mathfrak{P}}\right);$$

$$2^{um} = -\mathfrak{A}\mathfrak{B} \cdot a \left(\frac{1}{\mathfrak{P}} + \frac{1}{\eta-1}\right);$$

$$3^{um} = -\frac{\mathfrak{A}\mathfrak{B}\mathfrak{C} \cdot a}{\eta-1} \left(1 - \frac{1}{z\mathfrak{M}}\right);$$

$$4^{um} = \frac{\mathfrak{A}\mathfrak{B}\mathfrak{C}\mathfrak{D} \cdot a}{\mathfrak{M}} \left(\frac{1}{z(\eta-1)} + 1\right);$$

distantia vero oculi

$$O = \frac{t}{\mathfrak{M}\mathfrak{M}} = \frac{1}{z} \cdot \frac{1}{\mathfrak{C}} \left(1 + \frac{1}{\mathfrak{C}\mathfrak{M}}\right).$$

Deinde simili modo, ut iam ante notauimus, littera x siue per experientiam siue ex formula nota defini poterit. Tum vero erit lentis secundae semidiameter aperturae:

$$= \frac{1}{z} \cdot q \cdot q + \frac{z}{\mathfrak{P}} = \frac{1}{z} \cdot \frac{q}{\mathfrak{M}} + \frac{z}{\mathfrak{P}}.$$

tertia vero lentis.

$$= \frac{1}{z} \cdot r \cdot r + \frac{z}{\eta-1} = \frac{1}{z} \cdot \frac{(\eta-1)r}{\mathfrak{C}\mathfrak{M}} + \frac{z}{\eta-1}.$$

reliquae vero lentes quia sunt vtrinque aequae convexae, maximam aperturam admittunt. Pro spatio deni-

denique conspicuo erit

$z = \frac{1}{2} \cdot \frac{\epsilon \cdot a}{\epsilon m + \eta - 1} = \frac{a}{2m}$ proxime
et mensura claritatis $= 20 \cdot \frac{x}{m}$.

COROLL. I.

275. Si littera i contineatur intra hos limites, scilicet

$$i > \frac{\epsilon}{(\eta-1)(2+\epsilon)} \text{ et } i < \frac{\epsilon}{(\eta-1)(2-\epsilon)}$$

tum fit $P > 1$ et litterae \mathfrak{B} et B negatiuae, littera vero A sumi debet positiue; vnde omnia elementa eiusdem sunt naturae, vti in casu praecedente.

COROLL. 2.

276. Sin autem adeo fuerit $i > \frac{\epsilon}{(\eta-1)(2-\epsilon)}$; tum littera P vnitatis fit minor, hincque tam littera \mathfrak{B} , quam B fiunt positiuae; at vero littera A esse debet negatiua, id quod duplici modo euenire potest, altero quo $\mathfrak{A} > 1$; altero vero, quo $\mathfrak{A} < 0$. quo posteriore casu lens prima euaderet concaua et instrumentum multis incommodis foret obnoxium.

COROLL. 3.

277. Sin autem fuerit $i = \frac{\epsilon}{(\eta-1)(2-\epsilon)}$; tum fiet $P = 1$. hincque $\mathfrak{B} = 0$ et $B = 0$. Tum igitur ne fiat $q = 0$, necesse erit sumi $A = \infty$, ita, tamen, vt fit $AB = A\mathfrak{B} = -\frac{q}{a}$. Vnde cum fit $1 - P = \mathfrak{B}$

A a a 3

erit.

erit primum interuallum

$$= A a \left(1 - \frac{1}{p} \right) = - A \mathfrak{B}. a = q$$

quare ob $\mathfrak{A} = 1$. fient distantiae focales

$$p = a; q = q; r = \frac{Cq}{\eta - 1};$$

$$s = - \frac{CD}{i(\eta - 1)\mathfrak{M}} q \text{ seu } s = \frac{2C}{\mathfrak{M}} \cdot q \text{ et}$$

$$t = - \frac{CD}{\mathfrak{M}} \cdot q \text{ seu } t = \frac{2i(\eta - 1)C}{2i(\eta - 1) + 1} \cdot \frac{q}{\mathfrak{M}}.$$

Interualla vero lentium

$$1^{um} = q;$$

$$2^{um} = \frac{\eta}{\eta - 1} \cdot q;$$

$$3^{um} = \frac{C}{\eta - 1} \left(1 - \frac{1}{i\mathfrak{M}} \right) \cdot q;$$

$$4^{um} = - \frac{CD}{\mathfrak{M}} \left(1 + \frac{1}{i(\eta - 1)} \right) q.$$

$$= \frac{2i(\eta - 1) + 2}{2i(\eta - 1) + 1} \cdot \frac{C}{\mathfrak{M}} \cdot q.$$

In reliquis vero momentis nihil mutandum occurrit.

Scholion.

278. Probe autem hic est notandum, casus in his duobus postremis corollariis contentos neutiquam ad telescopia transferri posse. Pro telescopiis enim ob $a = \infty$ necessario sumi debet $\mathfrak{A} = 0$ et $A = 0$, cum in his casibus debeat esse \mathfrak{A} vel infinitum vel negativum.

Exempl. I.

279. Sumamus $\eta = 2$ et quia C in iis tantum terminis occurrit, qui per \mathfrak{M} sunt diuisi ideoque

que semper numerum praemagnum significare debet, pro \mathcal{C} recte unitatem assumere poterimus, hinc ergo pro littera R primus limes erit $i > \frac{1}{2}$, ut nostra instrumenta ad casum corollarii primi pertineant, sumi quoque debet $i < 1$ hinc ergo capiatur $i = \frac{1}{2}$, ut fit $R = \frac{1}{2} \mathcal{M}$; unde colligemus $k = \frac{1}{2}$ et $P = 2$. Deinde $\mathcal{B} = -1$ et $B = -\frac{1}{2}$ et $\mathcal{D} = -1$. hinc $D = -\frac{1}{2}$. Hinc distantiae focales erunt

$$\begin{aligned} p &= 2a; \quad q = \frac{1}{2} A a; \\ r &= \frac{\mathcal{C}}{2} A a; \quad s = \frac{\mathcal{C} A a}{\mathcal{M}} \text{ et} \\ t &= \frac{1}{4} \cdot \frac{\mathcal{C}}{\mathcal{M}} A a. \text{ seu } t = \frac{1}{4} s. \end{aligned}$$

Interualla vero lentium erunt

$$\begin{aligned} 1^{mum} &= \frac{1}{2} A a; \\ 2^{dum} &= \frac{3}{4} A a; \\ 3^{tium} &= \frac{\mathcal{C}}{2} \left(1 - \frac{2}{\mathcal{M}}\right) A a; \\ 4^{tum} &= \frac{3}{4} \cdot \frac{\mathcal{C}}{\mathcal{M}} A a. \end{aligned}$$

ac distantia oculi $O = \frac{1}{2} t \left(1 + \frac{1}{\mathcal{M}}\right)$.

Semidiametri denique aperturarum lentis

$$\begin{aligned} 1^{mae} &= x; \\ 2^{dae} &= \frac{1}{2} \frac{q}{\mathcal{M}} + \frac{1}{2} x; \\ 3^{tae} &= \frac{1}{2} \frac{r}{\mathcal{M}} + x. \end{aligned}$$

Exempl.

Exemplum 2.

280. Maneat $\eta = 2$, fumatur vero $i = 1$,
 fue $R = \mathfrak{M}$, atque erit $k = 1$ et $P = 1$, tum vero
 $\mathfrak{B} = 0 = B$ et $\mathfrak{D} = -2$, et $D = -\frac{2}{3}$; vnde ex co-
 rollario tertio nanciscimur

$$p = a; q = q; r = q;$$

$$s = \frac{2C}{\mathfrak{M}} q; t = \frac{2}{3} \cdot \frac{C}{\mathfrak{M}} \cdot q. \text{ feu } t = \frac{1}{3} s.$$

Intervalla vero lentium erunt

$$1^{mum} = q;$$

$$2^{dum} = 2q;$$

$$3^{tium} = C \left(1 - \frac{1}{\mathfrak{M}} \right) q;$$

$$4^{tum} = \frac{4}{3} \cdot \frac{C}{\mathfrak{M}} q;$$

reliqua vero momenta perinde ac in praecedente
 exemplo se habebunt.

Exempl. 3.

281. Maneat $\eta = 2$, et fumatur $i = 2$, vt fit
 $R = 2 \mathfrak{M}$; erit ergo $k = \frac{3}{2}$ et $P = \frac{4}{3}$, vnde $\mathfrak{B} = \frac{1}{3}$,
 hinc $B = \frac{1}{3}$; tum vero $\mathfrak{D} = -4$ et $D = -\frac{4}{3}$. Hinc
 cum A debeat esse negatium, statuatur

$$A = -\alpha, \text{ vt fit } \mathfrak{A} = -\frac{\alpha}{1-\alpha} = \frac{\alpha}{\alpha-1};$$

ex quo distantiae focales erunt

$$p = \frac{\alpha}{\alpha-1} \cdot a; q = \frac{1}{2} \alpha a;$$

$$r = \frac{1}{2} C \alpha a; s = \frac{1}{2} \cdot \frac{C}{\mathfrak{M}} \alpha a$$

et

$$\text{et } t = \frac{1}{5} \frac{C}{\mathcal{M}} \alpha a \text{ siue } t = \frac{2}{5} s.$$

Intervalla vero erunt

$$1^{mum} = \frac{1}{4} \alpha a;$$

$$2^{dum} = \frac{2}{10} \alpha a.$$

$$3^{tium} = \frac{1}{4} C \left(1 - \frac{1}{2\mathcal{M}} \right) \alpha a.$$

$$4^{tum} = \frac{3}{10} \frac{C}{\mathcal{M}} \alpha a.$$

Reliqua vero momenta sunt iterum, vt in exemplo primo. Hic autem probe notandum est, si capiatur $\alpha = 1$. prodire $p = \infty$ ideoque primam lentem plane reici posse, ita, vt microscopium ex solis lenticulis posterioribus componatur. Quia autem tum confusio prodiret enormis, cum in formulis capitae praecedentis sumi deberet $\mathcal{M} = \frac{1}{2}$, hoc instrumentum merito reuicimus multoque magis ea, quae prodirent, si esset $\alpha < 1$. Ac si α unitatem haud parum superet, haec instrumenta in praxi usum habere posse videntur.

Exempl. 4.

282. Sit nunc $\eta = 4$ et cum primus limes sit $i > \frac{1}{5}$, pro casu corollari primi sumamus $i < \frac{1}{5}$; sit igitur $i = \frac{1}{5}$ sumto $C = 1$, critque $k = \frac{1}{2}$; $P = 2$; hinc $\mathcal{B} = -1$; $B = -\frac{1}{2}$; $\mathcal{D} = -1$; $D = -\frac{1}{2}$; unde distantiae focales erunt

$$p = \mathcal{A} a; \quad q = \frac{1}{2} A a;$$

Tom. III.

B b b

$$r = \frac{1}{5}$$

$$r = \frac{1}{2} A a; s = \frac{C}{2R} A a; \text{ et}$$

$$t = \frac{1}{4} \cdot \frac{C}{2R} A a. \text{ seu } t = \frac{1}{4} s.$$

Intervalla vero lentium

$$1^{mum} = \frac{1}{2} A a;$$

$$2^{dum} = \frac{5}{12} A a;$$

$$3^{tium} = \frac{C}{2} \left(1 - \frac{C}{2R} \right) A a;$$

$$4^{tum} = \frac{3}{4} \frac{C}{2R} A a.$$

$$\text{Oculi distantia } O = \frac{1}{2} t \left(1 + \frac{3}{2R} \right).$$

$$\text{Porro } z = \frac{a}{2 \left(\frac{C}{2R} + 1 \right)}.$$

Semidiametri porro aperturarum erunt lentis

$$\text{primae} = x;$$

$$\text{secundae} = \frac{1}{2} \frac{q}{2R} + \frac{1}{2} x;$$

$$\text{tertia} = \frac{3}{2} \cdot \frac{r}{2R} + \frac{1}{2} x.$$

Exempl. 5.

283. Maneat $\eta = 4$, at sumatur $i = \frac{1}{2}$ secundum Coroll. tertium; eritque $k = 1$ et $P = 1$; vnde colligantur distantiae focales

$$p = a; q = q; r = \frac{1}{2} q;$$

$$s = \frac{2C}{2R} q. \text{ et } t = \frac{2}{3} \cdot \frac{C}{2R} \cdot q = \frac{1}{3} s.$$

et lentium intervalla

$$1^{mum} = q; 2^{dum} = \frac{4}{3} q;$$

3^{tium}

$$3^{tium} = \frac{C}{5} \left(1 - \frac{3}{5R} \right) q;$$

$$4^{tum} = \frac{4}{5} \cdot \frac{C}{5R} q = 2 t.$$

Distancia oculi $O = \frac{1}{2} t \left(1 + \frac{5}{5R} \right)$ et reliqua momenta omnia sunt, vt ante.

Exempl. 6.

284. Manente $\eta = 4$, sumatur $i = 1$. eritque $k = 4$ et $P = \frac{3}{4}$; tum vero $\mathfrak{B} = \frac{1}{4}$ et $B = \frac{1}{5}$ et $\mathfrak{D} = -6$; $D = -\frac{6}{7}$. Cum igitur B sit positium, littera A negatiua esse debet. Sit igitur $A = -a$ fietque $\mathfrak{A} = \frac{a}{a-1}$, vnde prodibunt distantiae focales

$$p = \frac{a}{a-1} \cdot a; q = \frac{1}{5} a a;$$

$$r = \frac{1}{5} C \cdot a a; s = \frac{2}{5} \cdot \frac{C}{5R} a a. \text{ et}$$

$$t = \frac{2}{7} \cdot \frac{C}{5R} a a; t = \frac{3}{7} s.$$

et interualla lentium

$$1^{mum} = \frac{1}{5} a a.$$

$$2^{dum} = \frac{5}{9} a a.$$

$$3^{tium} = \frac{1}{9} C \cdot \left(1 - \frac{5}{5R} \right) a a.$$

$$4^{tum} = \frac{8}{27} \cdot \frac{C}{5R} a a.$$

Distancia oculi O cum reliquis momentis eadem, ac ante, manet.

Problema 3.

285. Microscopium ex sex lentibus construere, quae ita sint dispositae, vt prior imago realis in in-

B b b 2

terual-

ternallum secundum, posterior vero in quartum incidat.

Solutio.

Quinque igitur litterarum P, Q, R, S, T secunda et quarta debent esse negatiuae; quare ponatur $Q = -k$ et $S = -k'$, vt sit $PkRk'T = M = \frac{ma}{b}$. Hinc erit vltimae lentis distantia focalis

$$u = -\frac{ABCDE}{\mathfrak{M}} \cdot a,$$

quae debet esse positua aeque ac lentium interualla, quae sunt

$$1^{mum} = A a \left(1 - \frac{1}{P}\right);$$

$$2^{dum} = -\frac{AB}{P} a \left(1 + \frac{1}{k}\right);$$

$$3^{tium} = -\frac{ABC}{Pk} a \left(1 - \frac{1}{R}\right);$$

$$4^{tum} = \frac{ABCD}{Pkk'} a \left(1 + \frac{1}{k'}\right);$$

$$5^{tum} = \frac{ABCDE}{Pkk'R} a \left(1 - \frac{1}{T}\right).$$

Ob quintum ergo debet esse $T < 1$. ob quartum vero $E < 0$. hincque $ABCD > 0$. Ob secundum vero esse debet $AB < 0$. hincque etiam CD negatiuum. Statuatur nunc $M = \frac{q+r+s+t+u}{\mathfrak{M}-1}$ et quia campus maximus desideratur, sumi poterit $s = 1$; $t = 1$ et $u = 1$, vt fiat

$$M = \frac{q+r+3}{\mathfrak{M}-1} \text{ hincque } z = M a \xi = \frac{1}{4} M a.$$

et distantia oculi $O = \frac{u}{\mathfrak{M}M}$ quae cum sit positua, destructio marginis colorati praebet

o =

$$0 = \frac{q}{P} - \frac{r}{Pk} - \frac{1}{PER} + \frac{1}{Pkk'R} + \frac{1}{Pkk'R'T}$$

vnde colligitur

$$\frac{1}{k'}(1 + \frac{1}{T}) = -qkR + rR + 1$$

et quia constat, esse $T < 1$, statuatur breuitatis gratia $1 + \frac{1}{T} = \mathfrak{D}$, vt fit $\mathfrak{D} > 2$, hincque $T = \frac{1}{\mathfrak{D}-1}$; ex quo habebitur

$$\frac{1}{k'} = -\frac{qkR + rR + 1}{\mathfrak{D}}; \text{ ergo ob } \frac{Pkk'R}{\mathfrak{D}-1} = \mathfrak{M}$$

fiet statim

$$PkR = \frac{(\mathfrak{D}-1)\mathfrak{M}}{k'} = \frac{\mathfrak{M}(\mathfrak{D}-1)(1+rR-qkR)}{\mathfrak{D}}$$

Praeterea iam considerandae sunt aequationes sequentes:

- 1°. $\mathfrak{B}q = (P-1)M$;
- 2°. $\mathfrak{C}r = -(1+Pk)M - q$;
- 3°. $\mathfrak{D} = -(1+Pkk'R)M - q - r$;
- 4°. $\mathfrak{E} = (Pkk'Rk' - 1)M - q - r - 1$;

pro quarum resolutione statuamus breuitatis gratia

$$\frac{1-P}{\mathfrak{B}} = \zeta \text{ et } 1+Pk = \eta, \text{ vt fiat}$$

$$q = -\zeta M \text{ et } r = \frac{(\zeta-\eta)M}{\mathfrak{C}} \text{ ergo}$$

$$3 + q + r = \frac{3\mathfrak{E} + (\zeta(1-\mathfrak{C})-\eta)M}{\mathfrak{C}} = M(\mathfrak{M}-1);$$

vnde concluditur

$$M = \frac{3\mathfrak{E}}{\mathfrak{C}(\mathfrak{M}+\zeta-1)-\zeta+\eta}; \text{ vnde vicissim}$$

$$q = -\frac{3\zeta\mathfrak{E}}{\mathfrak{C}(\mathfrak{M}+\zeta-1)-\zeta+\eta} \text{ et } r = \frac{3(\zeta-\eta)}{\mathfrak{C}(\mathfrak{M}+\zeta-1)-\zeta+\eta}$$

Bbb 3

Nunc

Nunc ergo habebimus $P k R = (\eta - 1) R$, seu

$$(\eta - 1) R = \mathfrak{M} \frac{\theta - 1}{\theta} \left(1 + \frac{s R (\zeta + \zeta \mathfrak{E} k - \eta)}{\mathfrak{E} (\mathfrak{M} + \zeta - 1) - \zeta + \eta} \right)$$

vnde ob rationes ante allegatas litteram k definire conuenit; quem in finem notasse iuuabit litteras ζ et η vna cum \mathfrak{E} semper, prae multiplicatione \mathfrak{M} fore valde exiguas; alioquin enim campus praeter necessitatem diminueretur; contra vero R etiam fore numerum praemagnum; vnde superior illa aequatio inducet hanc formam

$$(\eta - 1) R = \frac{\theta - 1}{\theta} \left(\frac{\mathfrak{E} \mathfrak{M} + s R (\zeta - \eta) + s R \zeta \mathfrak{E} k}{\mathfrak{E}} \right)$$

ex qua sequitur

$$k = \frac{(\eta - 1) \mathfrak{E} \theta R - (\theta - 1) \mathfrak{E} \mathfrak{M} - s (\theta - 1) (\zeta - \eta) R}{s (\theta - 1) \zeta \mathfrak{E} R}$$

qui valor cum debeat esse positiuus, debeat esse

$$R > \frac{(\theta - 1) \mathfrak{E} \mathfrak{M}}{(\eta - 1) \mathfrak{E} \theta - s (\theta - 1) (\zeta - \eta)}$$

existente

$$(\eta - 1) \mathfrak{E} \theta > s (\theta - 1) (\zeta - \eta), \text{ seu } \mathfrak{E} > \frac{s (\theta - 1) (\zeta - \eta)}{(\eta - 1) \theta}.$$

Cum vero vt in problemate praecedente esse debeat $\mathfrak{E} < 1$; numerator illius limitis minor esse debet suo denominatore, hincque $\zeta < \frac{\theta (\eta - 1) - s \eta}{s (\theta - 1)}$. His ergo conditionibus obseruatis, ponamus breuitatis gratia iterum, vt ante, $R = i \mathfrak{M}$, ita, vt esse debeat

$$i > \frac{(\theta - 1) \mathfrak{E}}{(\eta - 1) \mathfrak{E} \theta - s (\theta - 1) (\zeta - \eta)}$$

habe-

habebimus inde

$$k = \frac{i(\eta-1)\mathcal{C}\theta - (\theta-1)\mathcal{C} - 3i(\theta-1)(\zeta-\eta)}{3i(\theta-1)\zeta\mathcal{C}} \text{ et } P = \frac{\eta-1}{k};$$

pro quo valore duos casus considerari conuenit. Si $P > 1$; tum debet esse $A > 0$. ideoque $B < 0$. quod quidem sponte euenit, cum prodeat $\mathfrak{B} < 0$. Hoc ergo euenit, quando $k < \eta - 1$; ex quo concluditur

$$i < \frac{(\theta-1)\mathcal{C}}{(\eta-1)\mathcal{C}\theta - 3(\theta-1)(\eta-1)\zeta\mathcal{C} + \zeta - \eta}$$

qui limes manifesto maior est superiore. Sin autem littera i adeo hunc limitem superet, tunc fiet $P < 1$. ideoque A negatiue sumi debet et quia \mathfrak{B} prodit positium, B eatenus tantum erit positium, vti requiritur, quatenus fit $\mathfrak{B} < 1$. Fit vero semper $\mathfrak{B} < 1$, nisi fuerit $\zeta < 1$. atque si etiam fuerit $\zeta < 1 - P$ casus erit impossibilis. Deinde cum fit

$$PkR = (\eta - 1)i\mathfrak{M},$$

neglectis terminis prae \mathfrak{M} valde paruis ob $\mathfrak{M}M = 3$ proxime, erit

$$\mathfrak{D} = -3i(\eta - 1), \text{ hinc } D = -\frac{3i(\eta-1)}{3i(\eta-1)+1}.$$

Porro cum fit

$$PkRk' = \frac{\mathfrak{M}}{T} = \mathfrak{M}(\theta - 1),$$

fiet eodem modo

$$\mathcal{E} = 3\mathfrak{D} - 4 \text{ et } E = \frac{3\theta-4}{3\theta-5} \text{ seu } E = -\frac{(\theta-4)}{3\theta-5}.$$

Hinc ergo distantiae focales ita se habebunt:

$$p = \mathfrak{A}a; q = -\frac{\mathfrak{A}\mathfrak{B}}{P}a; r = -\frac{\mathfrak{A}\mathcal{B}\mathcal{C}}{k-1}a;$$

$$s = + \frac{ABC \cdot z i (\eta - 1)}{(\eta - 1) i \cdot \mathfrak{M}} \cdot a = - \frac{z \cdot ABC}{\mathfrak{M}} \cdot a;$$

$$t = - \frac{z i (\eta - 1) (\theta - 1)}{(z i (\eta - 1) + 1) (\theta - 1)} \cdot \frac{ABC}{\mathfrak{M}} \cdot a;$$

$$u = - \frac{z (\eta - 1) (z \theta - 1)}{(z i (\eta - 1) + 1) (z \theta - 1)} \cdot \frac{ABC}{\mathfrak{M}} \cdot a;$$

vbi notetur esse quoque $q = \frac{1}{2} (1 - \frac{1}{p}) A a$, ita, vt sit q ad primum interuallum, vt 1: ζ . Interualla autem erunt

$$1^{um} = A a (1 - \frac{1}{p}) = \zeta q;$$

$$2^{um} = - A B \cdot a (\frac{1}{p} + \frac{1}{\eta - 1});$$

$$3^{tum} = - A B C \cdot a (\frac{1}{\eta - 1} - \frac{1}{z (\eta - 1) \mathfrak{M}}) \\ = - \frac{ABC}{\eta - 1} \cdot a (1 - \frac{1}{z \mathfrak{M}});$$

$$4^{tum} = \frac{ABCD \cdot a}{\mathfrak{M}} (\frac{1}{z (\eta - 1)} + \frac{1}{\theta - 1}) \\ = - \frac{z (i (\eta - 1) + \theta - 1)}{(z i (\eta - 1) + 1) (\theta - 1)} \cdot \frac{ABC}{\mathfrak{M}} \cdot a;$$

$$5^{tum} = - \frac{z i (\eta - 1) (z \theta - 1) (\theta - 1)}{(z i (\eta - 1) + 1) (\theta - 1) (z \theta - 1)} \cdot \frac{ABC}{\mathfrak{M}} \cdot a.$$

Distantia vero oculi erit

$$O = \frac{u}{\mathfrak{M} M} = \frac{1}{2} u. \text{ proxime.}$$

Spatii vero conspicui semidiameter erit

$$z = \frac{1}{4} a M = \frac{z \cdot a}{4 \cdot \mathfrak{M}}.$$

Tum vero semidiameter aperturae lentis primae est $= x$ siue per formulam notam siue per experientiam definiendus. Lentis vero

$$2^{dae} = \frac{1}{4} q q + \frac{\infty}{p} = \frac{z}{4} \cdot \frac{z}{\mathfrak{M}} q + \frac{\infty}{p};$$

$$3^{tiae} = \frac{1}{4} r r + \frac{\infty}{p k} = \frac{z}{4} \cdot \frac{z - \eta}{\mathfrak{M}} r + \frac{\infty}{\eta - 1};$$

reli quarum vero lentium, quae debent esse vtrunque
aeque

aeque conuexae, semidiametri aperturarum erunt respectiue $\frac{1}{2} s$; $\frac{1}{2} t$ et $\frac{1}{2} u$. Denique autem mensura claritatis fiet $= 20. \frac{\infty}{\theta}$.

COROLL

286. Si statuatur $\zeta = 0$, pro secunda lente erit $q = \infty$, qui casus eodem redit, ac si haec lens plane abesset; tum autem erit $k = \infty$ et $P = 0$, $\mathfrak{B} = \infty$ et $B = -1$, unde etsi primum interuallum fit $= -\infty$, ob secundam lentem deficientem interuallum primae et tertiae lentis fiet nihilominus finitum $= \frac{\eta}{\eta-1}$. A a . Deinde vero capi debet

$$i > \frac{(\theta-1)\mathfrak{C}}{(\eta-1)\mathfrak{C}\theta + \eta(\theta-1)}.$$

Pro \mathfrak{C} vero sufficit, vt capiatur intra limites 1 et 0, quandoquidem C debet esse numerus positius ob $D < 0$. Reliquae vero determinationes manent, vt ante, si modo notetur, esse $B = -1$.

COROLL. 2.

287. Quia hoc casu lens secunda tollitur, hoc modo solutio habebitur problematis, quo microscopium ex quinque lentibus constructum quaeritur, quae ita sint dispositae, vt prima imago realis in primum interuallum, posterior vero in tertium incidat; cuius ergo problematis solutio etiam suppeditat campum triplicatum.

Coroll. 3.

288. Quia in genere ob rationes ante allegatas littera C semper designare debet numerum satis magnum, ne scilicet lentes posteriores fiant nimis exiguae, satis prope erit $C = 1$ atque adeo in praxi tuto statuere licebit $C = 1$. Tum igitur sumi debet

$$i > \frac{\theta - 1}{(\eta - 1)\theta - 2(\theta - 1)(\zeta - \eta)} \text{ deinde vero}$$

$$k = \frac{2(\eta - 1)\theta - 2(\theta - 1)(\zeta - \eta) - \theta + 1}{2(\theta - 1)\zeta}$$

CASVS I.

quo $i = \infty$ et $\theta = 3$.

289. Hoc ergo casu esse debet $\zeta < \frac{3\eta - 1}{2}$ tum vero erit

$$k = \frac{3\eta - 2\zeta - 1}{2\zeta} \text{ et } P = \frac{2(\eta - 1)\zeta}{3\eta - 2\zeta - 1}.$$

Vt igitur fiat $P > 1$, debet esse $\zeta > \frac{3\eta - 1}{2\eta}$. Quare fiet $P > 1$, si capiatur ζ intra limites $\frac{3\eta - 1}{2}$ et $\frac{3\eta - 1}{2\eta}$; quo ergo casu A sumi debet positue et quia reperitur $B < 0$, sponte fit $B < 0$. Sin autem fit $\zeta < \frac{3\eta - 1}{2\eta}$, tunc erit $P < 1$ hincque $A < 0$, ita, vt debeat esse $B > 0$, hincque $B < 1$, erit vero $B < 1$, si $1 - P < \zeta$, siue $P > 1 - \zeta$ quod quidem semper euenit, nisi sit $\zeta < 1$. Superest ergo examinare casum $\zeta < 1$. et quia tum esse debet $P > 1 - \zeta$ oritur inde haec relatio $\zeta(5\eta - 1) - 2\zeta^2 > 3\eta - 1$. vnde patet, esse debere

$$\zeta = \frac{5\eta-1}{4} - \frac{1}{4} \sqrt{25\eta^2 - 34\eta + 9} - a$$

denotante a numerum quempiam positivum, siue

$$\zeta > \frac{5\eta-1}{4} - \frac{1}{4} \sqrt{25\eta^2 - 34\eta + 9}$$

qui ergo limes pro ζ pendet ab η , ita, ut sumto $\eta = 2$ debeat esse $\zeta > \frac{9-\sqrt{41}}{4}$ seu $\zeta > \frac{5}{4}$ si autem

fuerit $\eta = 4$, debet esse $\zeta > \frac{19-\sqrt{272}}{4}$ seu proxime

$\zeta > \frac{5}{4}$. At si sit $\eta = 6$, prodit $\zeta > \frac{29-\sqrt{785}}{4}$ seu

proxime $\zeta > \frac{5}{4}$ ut ante, sicque patet ζ nunquam infra $\frac{5}{4}$ accipi posse. Nunc igitur pronuntiare poterimus, limites, intra quos ζ capi debeat, esse $\frac{5}{4}$ et $\frac{5\eta-1}{4}$.

His notatis distantiae focales erunt

$$p = 2a; q = -\frac{A^2 B}{P} a = \frac{1}{2} \left(1 - \frac{1}{P}\right) A a.$$

$$r = -\frac{ABC}{\eta-1} a; s = -3 \frac{ABC}{\eta} a.$$

$$t = -\frac{5}{2} \frac{ABC}{\eta} a; u = -\frac{5}{4} \frac{ABC}{\eta} a.$$

et lentium intervalla

$$1^{um} = A a \left(1 - \frac{1}{P}\right);$$

$$2^{um} = -A B \left(\frac{1}{P} + \frac{1}{\eta-1}\right) a;$$

$$3^{um} = -\frac{ABC}{\eta-1} a;$$

$$4^{um} = -\frac{1}{2} \frac{ABC}{\eta} a;$$

$$5^{um} = -\frac{5}{4} \frac{ABC}{\eta} a.$$

Reliqua momenta se habent vti in problemate, quippe quae ab u non pendent.

Scholion.

290. Mirum hic videbitur, quod hoc casu tam P maius unitate, quam minus unitate fieri possit, cum in solutione problematis ostenderit, tum solum P fieri > 1 , quando littera i contineatur intra limites

$\frac{(\theta-1)\mathcal{E}}{(\eta-1)\mathcal{E}\theta-s(\theta-1)(\zeta-\eta)}$ et $\frac{(\theta-1)\mathcal{E}}{(\eta-1)\mathcal{E}\theta-s(\theta-1)(\zeta-\eta+(\eta-1)\mathcal{E})}$ quando vero i etiam posteriorem limitem superauerit; tum semper fore $P < 1$. Quare cum hic adeo sumserimus $i = \infty$, hinc utique sequi videtur semper esse debere $P < 1$, quod tamen, ut vidimus, secus euenit. Ad quod dubium soluendum natura posterioris limitis accuratius perpendi debet, si enim is ipse iam fieret infinitus, tum certe mirari desinemus, si etiam sumto $i = \infty$ reperiatur $P > 1$. Sin autem in hoc limite posteriore denominator non solum euanescat, sed adeo negatiuus euadat; tum ipse limes non tam negatiuus, quam infinito maior spectari debebit, ita, ut positio $i = \infty$ adhuc inter illos limites contineri sit censenda. Nunc igitur manifestum est, limitem posteriorem fieri $= \infty$, si sumto $\mathcal{E} = 1$ fuerit

$$\zeta = \frac{(\eta-1)\theta-s\eta}{s\eta(\theta-1)},$$

sumtoque $\mathcal{D} = 3$, uti fecimus, si fuerit $\zeta = \frac{s\eta-1}{2\eta}$. Sin autem sit $\zeta > \frac{s\eta-1}{2\eta}$ (semper autem esse debet $\zeta < \frac{s\eta-1}{2\eta}$); tum ille limes fit quasi infinito maior, hincque $i = \infty$ ipso minor; vnde necessario fieri debet

bebit $P > 1$. Sin autem fit $\zeta < \frac{2\eta-1}{2\eta}$; tum ille limes adhuc erit finitus ideoque valor $i = \infty$ illo erit sine dubio maior; vnde etiam his tantum casibus fiet $P < 1$. Hoc notato istum casum aliquot exemplis illustremus.

Exempl. I.

291. Sumamus $\eta = 2$ et cum pro ζ prior limes fit $\zeta < \frac{5}{4}$; posterior vero limes $\frac{5}{4}$; sumamus $\zeta = 2$, vt cadat intra hos limites; vnde fiet $k = \frac{1}{4}$ et $P = 4$, hinc $\mathfrak{B} = -\frac{3}{2}$ et $B = -\frac{3}{2}$ vnde distantiae focales erunt

$$p = 2a; q = -\frac{1}{4}\mathfrak{B}.Aa = \frac{3}{8}Aa;$$

$$r = \frac{3}{2}Aa; s = \frac{3}{2} \cdot \frac{C}{\mathfrak{R}}.Aa;$$

$$t = \frac{3}{2} \cdot \frac{C}{\mathfrak{R}}.Aa, u = \frac{3}{4} \cdot \frac{C}{\mathfrak{R}}.Aa.$$

et lentium interualla

$$1^{um} = \frac{3}{4}Aa; 2^{dum} = \frac{3}{4}Aa;$$

$$3^{tium} = \frac{3}{2}C.Aa; 4^{tium} = \frac{3}{10} \cdot \frac{C}{\mathfrak{R}}.Aa.$$

$$\text{et } 5^{tium} = \frac{3}{8} \cdot \frac{C}{\mathfrak{R}}.Aa.$$

Distantia oculi $O = \frac{1}{3}u$ proxime; $z = \frac{3a}{4\mathfrak{R}}$;

semidiameter aperturae lentis primae $= x$;

secundae $= \frac{3}{2\mathfrak{R}}q + \frac{\infty}{4}$; tertiae $= x$.

ac denique mensura claritatis $= 20 \cdot \frac{\infty}{\mathfrak{R}}$, vti semper.

Exemplum 2.

292. Manente $\eta = 2$, aequetur ζ ipsi alteri limiti, scilicet $\zeta = \frac{5}{4}$ fietque $k = 1$ et $P = 1$; hinc ergo prodit $\mathfrak{B} = 0$ et $B = 0$. Quare nec tam prima lens, quam interualla euanescent, sumi debet $A = \infty$ ideoque $\mathfrak{A} = 1$, ita, ut sit $A \mathfrak{B}$ sine $AB = -\frac{q}{2}$ et cum sit $\mathfrak{B} = \frac{1}{2}(1 - P)$, erit reuera $1 - P = \frac{5}{4}\mathfrak{B}$ hincque $Aa(1 - \frac{1}{P}) = \frac{5}{4}q$. Quare distantiae focales erunt

$$p = a; q = q; r = Cq;$$

$$s = 3 \cdot \frac{C}{\mathfrak{M}} \cdot q; t = \frac{5}{2} \cdot \frac{C}{\mathfrak{M}} \cdot q; u = \frac{5}{4} \cdot \frac{C}{\mathfrak{M}} \cdot q;$$

vbi secunda q arbitrio nostro relinquitur. Tum vero lentium interualla

$$1^{mum} = \frac{5}{4}q; 2^{dum} = 2q;$$

$$3^{tium} = Cq; 4^{tum} = \frac{1}{2} \cdot \frac{C}{\mathfrak{M}} \cdot q; 5^{tum} = \frac{5}{4} \cdot \frac{C}{\mathfrak{M}} \cdot q.$$

Valores O et x erunt, ut ante; at semidiameter aperturæ lentis

$$2^{dae} = \frac{25}{16\mathfrak{M}} q + x \text{ et } 3^{tia} = \frac{9}{16\mathfrak{M}} r + x.$$

Exempl. 3.

293. Manente $\eta = 2$, sumatur $\zeta < \frac{5}{4}$ et cum esse debeat $\zeta > \frac{5}{8}$ vti offendimus, sumatur $\zeta = \frac{3}{4}$ eritque $k = \frac{7}{8}$ et $P = \frac{7}{8}$; vnde fit $\mathfrak{B} = \frac{16}{27}$, hincque $B = \frac{16}{9}$, qui valor cum sit positivus, littera A negativè capi debet, vti etiam primum interuallum postulat ob

$$P < 1.$$

$P < 1$. Sit igitur $A = -a$ ideoque $\mathcal{A} = \frac{a}{a-1}$ erunt-
que distantiae focales

$$p = \frac{aa}{2a-1}; q = \frac{16}{9} aa;$$

$$r = \frac{16}{3} C aa; s = \frac{48}{5} \cdot \frac{C}{21} aa;$$

$$t = 8 \cdot \frac{C}{21} aa; u = 4 \cdot \frac{C}{21} aa.$$

Intervalla lentium

$$1^{mum} = \frac{4}{3} aa; 2^{dum} = \frac{32}{5} aa;$$

$$3^{tium} = \frac{16}{3} C aa; 4^{tum} = \frac{8}{5} \cdot \frac{C}{21} aa;$$

$$5^{tum} = 2 \cdot \frac{C}{21} aa.$$

Reliqua se habebunt, ut ante ac si qua differentia in
aperturis deprehenditur, ea in praxi attendi non me-
retur; interim tamen semidiameter secundae lentis

$$= \frac{9}{16} \frac{q}{21} + \frac{7}{3} x \text{ et } 3^{tiae} = \frac{3}{16} \frac{r}{21} + x.$$

Exempl. 4.

294. Statuatur nunc $\eta = 4$ et cum esse de-
beat $\zeta < \frac{11}{5}$, posterior vero limes sit $\frac{11}{5}$, quo scilicet
casus $P > 1$ et $P < 1$ distinguntur. Sumatur $\zeta = 3$
et erit $k = \frac{5}{3}$ et $P = \frac{18}{5}$. Vnde fit $\mathcal{B} = -\frac{13}{15}$ et
 $B = -\frac{13}{27}$. Distantiae ergo focales lentium erunt

$$p = \mathcal{A} a; q = \frac{13}{24} A a;$$

$$r = \frac{13}{24} C A a; s = \frac{39}{27} \cdot \frac{C}{21} A a;$$

$$t = \frac{65}{36} \cdot \frac{C}{21} A a; u = \frac{65}{112} \cdot \frac{C}{21} A a = \frac{1}{2} t.$$

Inter-

Interualla vero lentium

$$1^{mum} = \frac{13}{18} A a; \quad 2^{dum} = \frac{143}{504} A a;$$

$$3^{tium} = \frac{13}{24} \cdot C \cdot A a; \quad 4^{tum} = \frac{17}{28} \cdot \frac{C}{\mathfrak{M}} \cdot A a;$$

$$5^{tum} = \frac{65}{224} \cdot \frac{C}{\mathfrak{M}} \cdot A a.$$

Denique semidiameter aperturae lentis

$$2^{dae} = \frac{9}{4 \mathfrak{M}} \cdot q + \frac{5}{18} x. \quad \text{et} \quad 3^{tiae} = \frac{7}{4 \mathfrak{M}} r + \frac{1}{3} x.$$

Exempl. 5.

295. Manente $\eta = 4$, fit $\zeta = \frac{11}{8}$ ac erit $k = 3$
 et $P = 1$ vnde $\mathfrak{B} = \frac{1-P}{\zeta} = \frac{0}{\frac{11}{8}} (1-P) = 0$ et $B = 0$.
 Vnde assumi debet $A = \infty$, ita, vt fiat $\mathfrak{A} = 1$; tum
 igitur introducto q in calculum fiet $A \mathfrak{B} = AB = -\frac{q}{a}$;
 vnde fit $A a \left(\frac{1-P}{P} \right) = \frac{11}{8} q$ sicque distantiae focales
 erunt

$$p = a; \quad q = q;$$

$$r = \frac{1}{2} C q; \quad s = 3 \cdot \frac{C}{\mathfrak{M}} q;$$

$$t = \frac{5}{2} \cdot \frac{C}{\mathfrak{M}} q \quad \text{et} \quad u = \frac{5}{4} \cdot \frac{C}{\mathfrak{M}} q.$$

et lentium interualla

$$1^{mum} = \frac{11}{8} q; \quad 2^{dum} = \frac{4}{3} q;$$

$$3^{tium} = \frac{1}{2} C q; \quad 4^{tum} = \frac{1}{2} \cdot \frac{C}{\mathfrak{M}} q$$

$$\text{et} \quad 5^{tum} = \frac{5}{4} \cdot \frac{C}{\mathfrak{M}} q.$$

Semidiameter aperturae lentis

$$2^{dae} = \frac{55}{42} \cdot \frac{q}{\mathfrak{M}} + x; \quad 3^{tiae} = \frac{67}{42} \cdot \frac{r}{\mathfrak{M}} + \frac{1}{2} x.$$

Exempl.

Exempl. 6.

296. Manente adhuc $\eta = 4$, fit $\zeta = \frac{2}{3}$, ac reperitur $k = \frac{29}{2}$ et $P = \frac{12}{29}$ vnde fit $\mathfrak{B} = \frac{51}{28}$ et $B = \frac{51}{7}$ ex quo tanto valore iam perspicuum est, huiusmodi microscopiis in praxi locum concedi non posse.

C A S V S II.

quo $\eta = 4$ et $\mathfrak{D} = 3$.

297. Quoniam debet esse $\eta > 1$ eiusque valor nimis parvus quibusdam incommodis est obnoxius; nimis magnus vero campo nocet, mediocri semper valore uti conueniet, cuiusmodi est $\eta = 4$; tum vero valor $\mathfrak{D} = 3$ seu $T = \frac{1}{2}$ idoneum interuallum inter vltimas lentes praebet; littera autem \mathfrak{C} tam parum ab vnitatem deficere debet, ut in nostris formulis liceat sumere $\mathfrak{C} = 1$. His praemissis pro ζ limes erit $\zeta < \frac{11}{2}$. Pro \mathfrak{C} vero habebimus

$$\mathfrak{C} < 1 \text{ et } \mathfrak{C} > \frac{2}{3}(\zeta - 4);$$

vnde patet, etiam si fit $\zeta = \frac{11}{2}$, tamen fore $\mathfrak{C} = 1$, ita, ut certe sumi possit $\mathfrak{C} = 1$. Porro pro littera i fiet $i > \frac{2}{33-6\zeta}$. Deinde obtinebimus

$$k = \frac{(33-6\zeta)i-2}{6\zeta i} \text{ vnde fit } P = \frac{18\zeta i}{(33-6\zeta)i-2}$$

$$\text{Hinc oritur } \mathfrak{B} = \frac{(33-24\zeta)i-2}{((33-6\zeta)i-2)\zeta}$$

Hic duos casus distingui oportet, prior est, quo fit $P > 1$ hincque A valorem posituum habere debet, quod

Tom. III.

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euenit

evenit, si \mathfrak{B} fiat ≤ 0 ideoque $i < \frac{2}{33-24\zeta}$, siue quando i continetur intra limites $\frac{2}{33-6\zeta}$ et $\frac{2}{33-24\zeta}$ atque hoc casu etiam B fit negativum, ita, ut sit $AB < 0$. Alter casus, quo $P < 1$; locum habet, si fuerit $i > \frac{2}{33-24\zeta}$, quo casu A valorem habebit negativum; quo igitur B nanciscatur valorem positivum, debet esse $\mathfrak{B} < 1$. hincque $i < \frac{2(1-\zeta)}{6\zeta^2-57\zeta+33}$. Cum igitur sit $i > \frac{2}{33-24\zeta}$, id evenit, si fuerit

$$\frac{1-\zeta}{6\zeta^2-57\zeta+33} > \frac{2}{33-24\zeta}$$

unde sequeretur $\zeta > 0$. His igitur notatis distantiae focales erunt

$$p = \mathfrak{A}a; q = -\frac{AB}{P} \cdot a;$$

$$r = -\frac{2}{3} ABCa; s = -3 \cdot \frac{ABC}{\mathfrak{M}} \cdot a;$$

$$t = -\frac{45i}{4i+2} \cdot \frac{ABC}{\mathfrak{M}} \cdot a. \text{ et}$$

$$u = -\frac{45i}{36i+4} \cdot \frac{ABC}{\mathfrak{M}} \cdot a.$$

siue etiam erit $q = +\frac{2}{3}(1-\frac{1}{P})Aa$ et intervalla lentium sunt

$$1^{mum} = Aa(1-\frac{2}{P})$$

$$2^{dum} = -ABa(\frac{2}{P}+\frac{1}{3})$$

$$3^{tium} = -\frac{ABC}{3} \cdot a(1-\frac{1}{\mathfrak{M}})$$

$$4^{tum} = -\frac{3(3i+2)}{2(9i+1)} \cdot \frac{ABC}{\mathfrak{M}} \cdot a$$

$$5^{tum} = -\frac{45i}{3(9i+1)} \cdot \frac{ABC}{\mathfrak{M}} \cdot a.$$

Deinde

Deinde cum fit

$$M = \frac{3}{2n+3} \text{ fiet } z = \frac{1}{4} \cdot \frac{3a}{2n+3}$$

et distantia oculi

$$O = \frac{1}{3} z \left(1 + \frac{z}{2n} \right).$$

Semidiameter porro aperturæ lentis secundæ erit

$$= \frac{3}{4} \frac{z}{2n} q + \frac{z}{P}, \text{ et } 3^{tiae} = \frac{3}{4} \frac{(z^2 - 4)r}{2n} + \frac{1}{3} z.$$

Denique definito x erit mensura claritatis $= \frac{20 \cdot x}{2n}$.

Exempl. I.

298. Sit $\zeta = 0$ et cum esse debeat $i > \frac{2}{37}$, præterea vero pro $P > 1$ fit $i < \frac{2}{37}$ statuamus $i = \frac{2}{37}$ eritque $P = 0$. siue P non determinatur, modo non fit minus unitate; adeoque $\mathfrak{B} = -\frac{(P-1)}{2}$, ita, ut \mathfrak{B} semper fit ∞ , nisi capiatur $P = 1$. Primo igitur non fit $P = 1$. erit $\mathfrak{B} = \infty$ et $B = -1$, ita, ut A sit maius nihilo; hinc igitur distantie focales erunt $p = 2a$; $q = \infty$; seu quod idem est secunda lens tollitur;

$$r = \frac{1}{3} A C \cdot a; \quad s = \frac{3AC}{2n} \cdot a;$$

$$i = \frac{15}{17} \cdot \frac{AC}{2n} \cdot a \text{ et } u = \frac{15}{34} \cdot \frac{AC}{2n} \cdot a.$$

et interualla lentium

$$1^{tum} + 2^{dum} = \frac{4}{3} A a;$$

$$3^{tium} = \frac{AC}{3} \left(1 - \frac{33}{2 \cdot 2n} \right); \quad 4^{tum} = \frac{75}{34} \cdot \frac{AC}{2n} \cdot a;$$

$$\text{et } 5^{tum} = \frac{15}{136} \cdot \frac{AC}{2n} \cdot a.$$

Reliqua manent, nisi quod sit semidiameter aperturae lentis tertiae $= 3. r + \frac{1}{2} x$.

Sin autem caperetur $P = 1$, utcumque calculus instituat, primum interuallum semper euanesceret; verum superfluum est, ad hunc casum attendere, cum in prioribus formulis littera P plane ex calculo euauerit, ita, ut illae formulae subsistant, quicumque valor ipsi P tribuatur atque adeo non solum si ponatur $P = 1$, sed etiam si P unitate minus sumeretur, quod etsi nostrae hypothesei repugnat, tamen ob lentem secundam prorsus deficientem haec anomalia admitti debet.

Exempl. 2.

299. Maneat $\zeta = 0$, sed capiatur $i > \frac{2}{33}$, si fieri queat, quo casu fiet $P < 1$, quia autem hoc ipso casu iterum esse debet $i < \frac{2}{33}$, hic casus ad praecedentem redigitur; quem quidem iam notauimus aequae ad valores $P < 1$, quam ad $P > 1$ patere. Interim tamen cum secunda lens plane deficiat, posterior limes $i < \frac{2}{33}$ sponte cessat, ita, ut nunc liceat assumere $i > \frac{2}{33}$, uti iam obseruauimus in Coroll. 1. problemati subnexo. Tum igitur erit

$$p = 2a; q = \infty; \text{ seu lens secunda deest;}$$

$$r = \frac{AC}{3} \cdot a; s = \frac{2AC}{33} \cdot a;$$

$$t = \frac{45 \cdot i}{18i + 2} \cdot \frac{AC}{33} \cdot a; u = \frac{45 \cdot i}{36i + 4} \cdot \frac{AC}{33} \cdot a.$$

Inter-

Interualla vero lentium

$$1^{mum} + 2^{dum} = \frac{4}{3} \cdot A a;$$

$$3^{tium} = \frac{1}{3} A C a \left(1 - \frac{1}{201} \right);$$

$$4^{tum} = \frac{s(3i+2)}{2(9i+1)} \cdot \frac{AC}{201} \cdot a;$$

$$5^{tum} = \frac{45i}{2(9i+1)} \cdot \frac{AC}{201} \cdot a.$$

Quod autem ad litteram i attinet, quoniam k non amplius in calculum ingreditur, ex aequatione, vnde k definiuimus, iam i definiatur eritque $i = \frac{2}{33-62} = \frac{2}{33}$. Proprie autem hae formulae continent solutionem problematis, quo quinque tantum lentes postulatur, ita disponendae, vt ambae imagines reales in primum et tertium interuallum incidant.

Exempl. 3.

300. Sit $\zeta = \frac{1}{2}$, sumique debeat $i > \frac{1}{17}$ fietque $P > 1$, si fuerit $i < \frac{2}{27}$; sin autem sit $i > \frac{2}{27}$ simul fiet $P < 1$. Hic vero sumamus $i = \frac{1}{12}$; fietque $P = \frac{3}{2}$, hinc $B = -1$, et $B = -\frac{1}{2}$, vnde distantiae focales erunt

$$p = 2a; q = \frac{2}{3} A a;$$

$$r = \frac{1}{3} A C a; s = \frac{3}{2} \cdot \frac{AC}{201} \cdot a;$$

$$t = \frac{45}{24} \cdot \frac{AC}{201} \cdot a; u = \frac{15}{36} \cdot \frac{AC}{201} \cdot a.$$

et lentium interualla

$$1^{mum} = \frac{1}{3} A a; 2^{dum} = \frac{1}{2} A a;$$

D d d 3

3^{tium}

$$3^{tium} = \frac{1}{2} A C. a \left(1 - \frac{12}{301} \right);$$

$$4^{tum} = \frac{27}{21} \cdot \frac{AC}{301} \cdot a; \quad 5^{tum} = \frac{15}{172} \cdot \frac{AC}{301} \cdot a.$$

Reliqua momenta sunt, vt ante, nisi quod sit semidiameter aperturæ

$$2^{dae} \text{ lentis} = \frac{5}{8} \cdot \frac{q}{301} + \frac{2}{3} x;$$

$$\text{et } 3^{tiae} = \frac{21}{8} \cdot \frac{r}{301} + \frac{1}{3} x.$$

Has formulas commode ad telescopia adplicare licet, quia posito $A a = p$ longitudo tantum fit $= \frac{5}{8} p + \frac{1}{3} C. p$, ita, vt ea non multum superet p , etiam si pro C numerus satis magnus capiatur.

Exempl. 4.

301. Maneat $\zeta = \frac{1}{2}$, sed sumatur $i > \frac{2}{31}$ et quia hinc fit $P = \frac{9i}{301-2}$ ideoque $B = \frac{21i-2}{151-1}$, vt fiat $B < 1$. debet esse $21. i - 2 < 15 i - 1$ siue $i < \frac{1}{6}$. Capiatur ergo $i = \frac{1}{7}$ fietque $P = \frac{9}{204}$ et $B = \frac{5}{7}$, hinc $B = \frac{5}{7}$; ergo A debet esse negativum. Statuatur ergo $A = -a$ et distantiae focales erunt

$$p = \frac{a}{1-a} a; \quad q = \frac{10}{9} a a.$$

$$r = \frac{5}{8} C. a a. \quad s = \frac{15}{8} \cdot \frac{C}{301} \cdot a a.$$

$$t = \frac{225}{68} \cdot \frac{C}{301} a. a. \quad \text{et } u = \frac{225}{196} \cdot \frac{C}{301} \cdot a a.$$

et lentium interualla

$$1^{tum} = \frac{5}{9} \cdot a a. \quad 2^{dum} = \frac{65}{18} \cdot a a.$$

$$3^{tum} = \frac{5}{2} \cdot C \cdot a a \left(1 - \frac{8}{27}\right).$$

$$4^{tum} = \frac{57}{34} \cdot \frac{C}{27} \cdot a a.$$

$$5^{tum} = \frac{225}{272} \cdot \frac{C}{27} \cdot a a.$$

Tum vero semidiameter aperturae lentis secundae et tertiae

$$= \frac{3}{8} \cdot \frac{q}{27} + \frac{11}{9} x \text{ et } \frac{21}{8} r + \frac{1}{2} x.$$

Has autem formulas ad telescopia adplicare non licet, quia A erat negativum.

Exempl. 5.

302. Sit $\zeta = 1$. et cum sumi debeat $i > \frac{2}{27}$ atque ut prodeat $P > 1$, $i < \frac{2}{9}$ capiatur $i = \frac{1}{10}$ fietque $P = \frac{11}{7}$ et $\mathfrak{B} = -\frac{11}{7}$ et $B = -\frac{11}{18}$; unde distantiae focales erunt

$$p = 21 a; q = \frac{11}{18} A a;$$

$$r = \frac{11}{34} \cdot C \cdot A a; s = \frac{11}{6} \cdot \frac{AC}{27} \cdot a;$$

$$t = \frac{55}{76} \cdot \frac{C}{27} \cdot A a; u = \frac{55}{152} \cdot \frac{C}{27} \cdot A a.$$

et lentium intervalla

$$1^{num} = \frac{11}{18} A a; 2^{dum} = \frac{145}{224} \cdot A a;$$

$$3^{tum} = \frac{11}{34} \cdot A C \cdot a \left(1 - \frac{30}{27}\right);$$

$$4^{tum} = \frac{253}{328} \cdot \frac{AC}{27} \cdot a. 5^{tum} = \frac{55}{304} \cdot \frac{AC}{27} \cdot a.$$

et semidiameter aperturae lentis

$$2^{dae} = \frac{3}{4} \cdot \frac{q}{27} + \frac{7}{18} x. 3^{tae} = \frac{9}{4} \cdot \frac{r}{27} + \frac{1}{2} x.$$

quas

quas formulas etiam commode ad telescopia transferre licet.

Exempl. 6.

303. Maneat $\zeta = 1$. sed sumatur $i = \frac{2}{3}$ fietque $P = 1$. et $\mathfrak{B} = 0$. hinc $B = 0$ hinc A capi debet $= \infty$, ideoque $\mathfrak{A} = 1$. tum autem esse debet

$$A \mathfrak{B} = AB = -\frac{q}{a} \text{ vnde fit}$$

$$A a \left(1 - \frac{1}{P}\right) = A a (P - 1) = -A \mathfrak{B} a = q.$$

hinc distantiae focales erunt

$$p = a; \quad q = q;$$

$$r = \frac{1}{3} C \cdot q; \quad s = 3 \cdot \frac{C}{\mathfrak{M}} q;$$

$$t = \frac{5}{3} \cdot \frac{C}{\mathfrak{M}} q; \quad u = \frac{5}{3} \cdot \frac{C}{\mathfrak{M}} q.$$

et lentium interualla

$$1^{um} = q; \quad 2^{dum} = \frac{4}{3} \cdot q;$$

$$3^{tium} = \frac{1}{3} C \left(1 - \frac{9}{2\mathfrak{M}}\right) q; \quad 4^{tum} = \frac{4}{3} \cdot \frac{C}{\mathfrak{M}} q;$$

$$5^{tum} = \frac{5}{12} \cdot \frac{C}{\mathfrak{M}} q.$$

et semidiametri aperturæ lentis

$$2^{dae} = \frac{3}{4} \cdot \frac{q}{\mathfrak{M}} + x;$$

$$3^{tiae} = \frac{9}{4} \cdot \frac{r}{\mathfrak{M}} + \frac{1}{3} x.$$

Exempl. 7.

304. Maneat $\zeta = 1$. sed capiatur $i > \frac{2}{3}$ et cum fit $P = \frac{18i}{27i-2}$, hincque $\mathfrak{B} = \frac{9i-2}{27i-2}$, quæ fractio iam sponte

spontè unitate est minor, sumamus ergo $i = 1$. erit $P = \frac{12}{25}$ et $B = \frac{7}{25}$, $B = \frac{7}{18}$ vnde A. debet esse negativum, statuatur ergo $= -a$; eruntque distantiae focales

$$p = \frac{a}{a-1} \cdot a; \quad q = \frac{7}{18} \cdot a a;$$

$$r = \frac{7}{54} \cdot C \cdot a a; \quad s = \frac{7}{6} \cdot \frac{C}{D} \cdot a a;$$

$$t = \frac{55}{40} \cdot \frac{C}{D} \cdot a a; \quad u = \frac{25}{20} \cdot \frac{C}{D} \cdot a a.$$

et interualla

$$1^{um} = \frac{12}{18} a a;$$

$$2^{dum} = \frac{217}{824} a a;$$

$$3^{tum} = \frac{7}{54} \cdot C \cdot a a \left(1 - \frac{1}{D}\right);$$

$$4^{ium} = \frac{7}{54} \cdot \frac{C}{D} \cdot a a;$$

$$5^{tum} = \frac{7}{32} \cdot \frac{C}{D} \cdot a a.$$

et semidiametri aperturæ lentis

$$2^{dae} = \frac{5}{4} \cdot \frac{q}{D} + \frac{25}{18} x.$$

$$3^{tiae} = \frac{9}{4} \cdot \frac{r}{D} + \frac{1}{5} x.$$

Exempl. 8.

305. Sit nunc $\zeta = 4 = \eta$ et cum esse debeat $i > \frac{2}{5}$; vt autem fiat $P > 1$, $i < -\frac{2}{18}$, qui limes vt infinito maior spectari debet, si enim sumissemus ζ minus, scilicet $\zeta = \frac{11}{8}$; tum prodisset $i < \frac{2}{5}$, quod indicio fuisset, i quantumuis magnum accipi posse, semperque fore $P < 1$. Quod autem de valore $\zeta = \frac{11}{8}$

Tom. III.

E e e

valet,

valet, multo magis de maioribus valet. Sit ergo $i = \frac{1}{3}$ eritque $P = 24$ et $\mathfrak{B} = -\frac{23}{4}$ et $B = -\frac{23}{17}$; vnde colliguntur distantiae focales:

$$p = 24a; q = \frac{23}{96} \cdot Aa;$$

$$r = \frac{23}{81} A \mathfrak{C} a; s = \frac{23}{9} \cdot \frac{AC}{\mathfrak{M}} \cdot a;$$

$$t = \frac{115}{72} \cdot \frac{AC}{\mathfrak{M}} \cdot a; u = \frac{115}{144} \cdot \frac{AC}{\mathfrak{M}} \cdot a.$$

et interualla lentium

$$1^{mum} = \frac{23}{24} \cdot Aa;$$

$$2^{dum} = \frac{23}{72} \cdot Aa;$$

$$3^{tium} = \frac{23}{81} \cdot CAa \left(1 - \frac{3}{\mathfrak{M}}\right);$$

$$4^{tum} = \frac{23}{24} \cdot \frac{AC}{\mathfrak{M}} \cdot a;$$

$$5^{tum} = \frac{115}{288} \cdot \frac{AC}{\mathfrak{M}} \cdot a.$$

et semidiameter aperturae lentis

$$2^{dae} = \frac{27}{\mathfrak{M}} + \frac{5}{24} x. \text{ et } 3^{tiae} = \frac{5}{3} x.$$

Exempl. 9.

306. Maneat $\zeta = 4$ et sit $i = 1$. fiet $P = \frac{72}{7}$; ergo $\mathfrak{B} = -\frac{65}{28}$; hinc $B = -\frac{65}{93}$; vnde colliguntur distantiae focales

$$p = 24a; q = \frac{65}{287} \cdot Aa;$$

$$r = \frac{65}{287} A \mathfrak{C} a; s = \frac{65}{28} \cdot \frac{AC}{\mathfrak{M}} \cdot a;$$

$$t = \frac{195}{287} \cdot \frac{AC}{\mathfrak{M}} \cdot a; u = \frac{195}{574} \cdot \frac{AC}{\mathfrak{M}} \cdot a = \frac{1}{2} t.$$

et

et interualla lentium

$$1^{mum} = \frac{65}{72} \cdot A a;$$

$$2^{dum} = \frac{65}{216} \cdot A a;$$

$$3^{tium} = \frac{65}{279} \cdot A C \cdot a \left(1 - \frac{1}{21} \right);$$

$$4^{tium} = \frac{65}{124} \cdot \frac{A C}{21} \cdot a;$$

$$5^{tium} = \frac{195}{495} \cdot \frac{A C}{21} \cdot a.$$

et femidiameter aperturæ lentis

$$2^{dae} = \frac{54}{21} + \frac{7}{72} x \text{ et } 3^{tiae} = \frac{1}{3} x.$$

Scholion.

307. Horum exemplorum ea inprimis sunt notatu digna, in quibus fiebat $P = 1$. quia tum litera A abibat in infinitum eratque $\mathcal{A} = 1$. In casu igitur secundo, quem hactenus sumus contemplati, statuamus in genere $P = 1$. ac sumi debeat $i = \frac{2}{33 - 24\zeta}$. Tum ergo erit $\mathcal{B} = 0$, simulque $B = 0$. vnde sumto $\mathcal{A} = 1$ et $A = \infty$ fiet $q = -A \mathcal{B} a$, hinc vicissim $A \mathcal{B} = A B = -\frac{q}{a}$, ita, vt nunc distantia focalis secundæ lentis q arbitrio nostro penitus relinquatur; tum autem ad primum interuallum inueniendum ob $\mathcal{B} = \frac{1-P}{\zeta}$, erit $A a \cdot \frac{1-P}{\zeta} = A a \cdot P - 1 = -A \mathcal{B} \cdot \zeta a = \zeta q$, atque hinc in genere distantie focales ita se habebunt:

$$p = a; q = q; r = \frac{1}{3} \zeta q;$$

$$s = \frac{2C}{21} q; t = + \frac{15}{17-8\zeta} \cdot \frac{C}{21} q;$$

$$u = \frac{15}{34-16\zeta} \cdot \frac{C}{21} \cdot q \text{ seu } u = \frac{1}{2} t.$$

E e e 2

Inter-

Interualla vero lentium erunt

$$1^{mum} = \zeta' q; \quad 2^{dum} = \frac{4}{3} q; \quad 3^{tium} = \frac{1}{3} q \left(1 - \frac{22-24\zeta}{22} \right) = \frac{1}{3} q \left(1 - \frac{5(11-3\zeta)}{22} \right);$$

$$4^{tum} = \frac{12(1-2\zeta)}{17-8\zeta} \cdot \frac{C}{22} q; \quad 5^{tum} = \frac{15}{4(17-8\zeta)} \cdot \frac{C}{22} \cdot q.$$

vbi ergo manifestum est, necessario sumi debere $\zeta < \frac{3}{2}$.
Tum vero erit semidiameter aperturæ lentis

$$2^{dae} = \frac{3}{4} \cdot \frac{\zeta}{22} \cdot q + x; \quad \text{et} \quad 3^{tiae} = \frac{3}{4} \cdot \frac{\zeta-1}{22} \cdot r + \frac{1}{3} x.$$

Ceterum hic manifestum est, istas formulas ad telescopia accommodari neūtiquam posse. Cum igitur omnes casus, qui quidem in praxi locum habere possunt, adeo pro sex lentibus euoluerimus, isque tantum campum conciliauerimus, quo maior desiderari vix queat; huic capiti finem imponimus ad sequens idque vltimum progressuri, in quo ostendimus, quemadmodum loco lentis obiectiuæ duas pluresue lentes siue ex eodem siue ex diuerso vitri genere factas substituendo omnis plane confusio tolli possit, vt hoc modo microscopia omnibus numeris absoluta nanciscamur.