

§. 172. Hæ autem singulæ inæqualitates ad numerum minorum fecundorum reductæ dabunt :

		log. coeff.
☉ = Præc.	+ 701 ¹¹ , 1 fin s	2,845780
—	75, 8 fin (2 η -s)	1,879971
—	56, 7 fin (2 η +s)	1,753736
+	3, 8 fin 2s	0,585551
+	1, 4 fin (2 η -2s)	0,132862
+	0, 5 fin (2 η +2s)	9,689472
+	140, 9 fin (s-s)	2,148809
—	116, 7 fin (s+s)	2,067000
+	153, 7 fin (2 η -s+s)	2,186530
—	2, 7 fin (2 η +s-s)	0,426300
+	45, 2 fin (2 η -s-s)	1,655200
+	1, 2 fin (2 η +s+s)	0,064600

Hic scilicet et inæqualitates, quas in capite præcedente inuenimus, et illas in hoc capite eruras simul sum complexus, vt coniunctim confectui exponerentur,

CAPUT

CAPUT XI.

INVESTIGATIO INÆQUALITATUM LUNAE
A PARALLAXI SOLIS PENDENTIIUM.

§. 173.

Jam in formulis nostris primariis ad eos quoque terminos progrediamur, qui littera ν sunt affecti, et quoniam est $1:\nu$ vt distantia Solis media ad distantiam Lunae mediam a Terra, erit $1:\nu$ vt parallaxis Lunae media ad parallaxin solis: ex quo inæqualitates Lunae, quae hinc oriuntur, a parallaxi solis pendere dicuntur. Quoniam vero valor ipsius ν est valde paruus, quippe $\frac{1}{285}$ propemodum, alios terminos non contemphabimur, nisi qui per ν ac per νk et νe sunt multiplicati, propterea quod magis compofiti fiant minimi.

§. 174. Ex terminis ergo iam inuentis hic reuenamus eos, qui sunt alicuius momenti, et cum his nos determinandos coniungamus; sic ergo:

$$\begin{aligned} \mathcal{R}k &= \mathcal{M} \operatorname{cof} 2\eta + \mathcal{E} k \operatorname{cof} r + \mathcal{D} k \operatorname{cof} (2\eta - r) + \mathcal{F} e \operatorname{cof} r + \mathcal{Q} e \operatorname{cof} (2\eta - s) \\ &\quad + \mathcal{G} k \operatorname{cof} (2\eta + s) + \mathcal{H} e \operatorname{cof} (2\eta + s) \\ &+ \mathcal{S} \nu \operatorname{cof} \nu + \mathcal{D} \nu k \operatorname{cof} (q - r) + \mathcal{R} \nu e \operatorname{cof} (q - s) \\ &+ \mathcal{B} \nu \operatorname{cof} 3\eta + \mathcal{G} \nu k \operatorname{cof} (q + r) + \mathcal{Q} \nu e \operatorname{cof} (q + s) \\ \nu &= A \operatorname{cof} 2\eta \dots + D k \operatorname{cof} (2\eta - r) + P e \operatorname{cof} r + Q e \operatorname{cof} (2\eta - s) \\ &\quad + E k \operatorname{cof} (2\eta + r) + R e \operatorname{cof} (2\eta + s) \\ &+ F \nu \operatorname{cof} q + H \nu k \operatorname{cof} (q - r) + K \nu e \operatorname{cof} (q - s) \\ &+ G \nu \operatorname{cof} 3\eta + J \nu k \operatorname{cof} (q + r) + L \nu e \operatorname{cof} (q + s) \end{aligned}$$

Non

JT

Non difficulter enim praeuidere licet, terminos, qui angulos r et s cum angulo 3η habeant coniunctos, fore tam exiguos, ut sine errore praetermitti queant.

§. 175. Quodsi iam retentis litterarum §. 153. inductarum valoribus, praeterca ponamus:

$$f = \frac{2kF + \mathcal{G}}{nn} ; g' = \frac{2kG + \mathcal{G}}{nn} ; h' = \frac{2kH + \mathcal{G}}{nn}$$

$$i' = \frac{2kJ + \mathcal{G}}{nn} ; k' = \frac{2kK + \mathcal{G}}{nn} ; l' = \frac{2kL + \mathcal{G}}{nn}$$

habebimus:

$$\frac{d\theta}{dt} = \frac{a-d'cf2\eta - e'kcf - d'kcf(2\eta - r) - p'e'cf s - q'e'cf(2\eta - s)}{nn} - \frac{e'kcf(2\eta + r)}{nn} - \frac{p'e'cf(2\eta + s)}{nn}$$

$$- f'v \cos \eta - b'vk \cos(\eta - r) - kvv \cos(\eta - s)$$

$$- g'v \cos 3\eta - h'vk \cos(\eta + r) - l'vv \cos(\eta + s)$$

atque

$$\frac{d\eta}{dt} = \frac{a-d'cf2\eta - e'kcf - d'kcf(2\eta - r) + (\frac{2}{n} - p')e'cf s - q'e'cf(2\eta - s)}{nn}$$

$$- \frac{e'kcf(2\eta + r)}{nn} - \frac{p'e'cf(2\eta + s)}{nn}$$

$$- f'v \cos \eta - b'vk \cos(\eta - r) - kvv \cos(\eta - s)$$

$$- g'v \cos \eta - h'vk \cos(\eta + r) - l'vv \cos(\eta + s)$$

§. 176. Jam vero pro his terminis ab v pendentibus sequentes colligemus aequationes.

$$R \equiv v \sin \eta \left(\frac{3}{2} + \frac{3F - 3G}{2nn} \right)$$

$$v \sin 3\eta \left(\frac{3}{2} + \frac{3F}{2nn} \right)$$

vk

$$v k \sin(\eta - r) \left(\frac{11}{2} + \frac{3J}{2nn} + \frac{3F}{nn} - \frac{3G}{nn} \right)$$

$$v k \sin(\eta + r) \left(\frac{11}{2} + \frac{3H}{2nn} - \frac{3G}{nn} + \frac{3F}{nn} \right)$$

$$v e \sin(\eta - s) \left(-\frac{1}{2} + \frac{3L}{2nn} - \frac{9F}{4nn} + \frac{9G}{4nn} \right)$$

$$v e \sin(\eta + s) \left(-\frac{1}{2} + \frac{3K}{2nn} + \frac{9G}{4nn} - \frac{9F}{4nn} \right)$$

$$\frac{d\theta}{dt} =$$

$$v \cos \eta \left\{ \frac{1}{2} - \frac{6F}{4nn} + \frac{3G}{4nn} - 2k \frac{\mathcal{G}}{nn} + \frac{9\mathcal{G}}{nn} + \frac{9\mathcal{G}}{2nn} \right\}$$

$$v \cos 3\eta \left\{ \frac{1}{2} - \frac{6G}{4nn} + \frac{3F}{4nn} - 2k \frac{\mathcal{G}}{nn} + \frac{9\mathcal{G}}{nn} \right\}$$

$$v k \cos(\eta - r) \left\{ \frac{11}{2} - \frac{6H}{4nn} + \frac{3bF}{4nn} + \frac{3J}{4nn} + \frac{3F}{2nn} + \frac{3G}{2nn} - 2k \frac{\mathcal{G}}{nn} \right.$$

$$\left. + \frac{9\mathcal{G}}{nn} + \frac{9\mathcal{G}}{2nn} + \frac{9\mathcal{G}}{2nn} + \frac{9\mathcal{G}}{2nn} \right\}$$

$$v k \cos(\eta + r) \left\{ \frac{11}{2} - \frac{6J}{4nn} + \frac{3bF}{4nn} + \frac{3H}{4nn} + \frac{3G}{2nn} + \frac{3F}{2nn} - 2k \frac{\mathcal{G}}{nn} \right.$$

$$\left. + \frac{9\mathcal{G}}{nn} + \frac{9\mathcal{G}}{2nn} + \frac{9\mathcal{G}}{2nn} + \frac{9\mathcal{G}}{2nn} \right\}$$

V

ve cos

v e col (q-s)

$$\left\{ \begin{aligned} & -\frac{3}{2} \frac{6K}{m} + \frac{3L}{4m} - \frac{3F}{4m} - \frac{9G}{8m} - \frac{9G}{8m} - 2 \frac{K}{m} \\ & + \frac{9E}{m} + \frac{9H}{m} + \frac{9I}{m} + \frac{9J}{m} \end{aligned} \right.$$

v e col (q+s)

$$\left\{ \begin{aligned} & -\frac{3}{2} \frac{6L}{m} + \frac{3K}{4m} - \frac{3F}{4m} - \frac{9G}{8m} - \frac{9G}{8m} - 2 \frac{K}{m} \\ & + \frac{9E}{m} + \frac{9H}{m} + \frac{9I}{m} + \frac{9J}{m} \end{aligned} \right.$$

sequentes partes geom. exponamus:

v col q

$$\left\{ \begin{aligned} & + \frac{39E}{m} + \frac{39G}{m} + \frac{3A\mathcal{G}}{m} + \frac{3AF}{m} + \frac{3AG}{m} \\ & + \frac{9A}{8m} + \frac{15A}{8m} \end{aligned} \right.$$

v col 3q

$$\left\{ \begin{aligned} & + \frac{39E}{m} + \frac{39G}{m} + \frac{3AF}{m} + \frac{9A}{8m} \end{aligned} \right.$$

v k col (q-s)

$$\left\{ \begin{aligned} & + \frac{39J}{m} + \frac{3EF}{m} + \frac{3DF}{m} + \frac{3EG}{m} + \frac{3AG}{m} + \frac{3DS}{m} \\ & + \frac{3E\mathcal{G}}{m} + \frac{3AJ}{m} + \frac{3DE}{m} + \frac{3EG}{m} - \frac{3AF}{m} - \frac{3AG}{m} \\ & + \frac{9D}{8m} + \frac{15E}{8m} \end{aligned} \right.$$

v k col (q+s)

$$\left\{ \begin{aligned} & + \frac{39H}{m} + \frac{3EF}{m} + \frac{3D\mathcal{G}}{m} + \frac{3EF}{m} + \frac{3A\mathcal{G}}{m} \\ & + \frac{3D\mathcal{G}}{m} + \frac{3E\mathcal{G}}{m} + \frac{3AH}{m} + \frac{3DG}{m} + \frac{3EF}{m} \\ & + \frac{3AF}{2m} + \frac{3AG}{2m} + \frac{9E}{8m} + \frac{15D}{8m} \end{aligned} \right.$$

v e col

v e col (q-s)

$$\left\{ \begin{aligned} & + \frac{39L}{m} + \frac{39F}{m} + \frac{3\mathcal{G}F}{m} + \frac{3NG}{m} + \frac{3AR}{m} \\ & + \frac{3P\mathcal{G}}{m} + \frac{3Q\mathcal{G}}{m} + \frac{3R\mathcal{G}}{m} + \frac{3AL}{m} + \frac{3PF}{m} + \frac{3QF}{m} \\ & + \frac{3RG}{m} + \frac{9P}{8m} + \frac{9Q}{8m} + \frac{15R}{8m} \end{aligned} \right.$$

v e col (q+s)

$$\left\{ \begin{aligned} & + \frac{39K}{m} + \frac{39F}{m} + \frac{3\mathcal{G}G}{m} + \frac{3RF}{m} + \frac{3AR}{m} + \frac{3PF}{m} \\ & + \frac{3Q\mathcal{G}}{m} + \frac{3R\mathcal{G}}{m} + \frac{3AK}{m} + \frac{3PF}{m} + \frac{3QF}{m} + \frac{3Q\mathcal{G}}{m} \\ & + \frac{9P}{8m} + \frac{9R}{8m} + \frac{15Q}{8m} \end{aligned} \right.$$

§. 177. Verum differendo nanciscemur

R

$$9 \sin q \quad [9p^2 - 9q^2 - a\mathcal{G} - \frac{1}{2}\mathcal{G}d^2 + \frac{1}{2}\mathcal{G}d^2]$$

v sin 3q

$$9 \sin (q-s) \quad [9p^2 + 9q^2 - \mathcal{G}d^2 + \frac{1}{2}\mathcal{G}d^2 - (\alpha-1)\mathcal{G}d^2 - \frac{1}{2}\mathcal{G}d^2]$$

v k sin (q+s)

$$+ 9p^2 - 9q^2 + \mathcal{G}d^2 + \frac{1}{2}\mathcal{G}d^2 - (\alpha+1)\mathcal{G}d^2$$

v e sin (q-s)

$$+ 9p^2 + 9q^2 - 9R^2 - \frac{1}{2}(\frac{2}{m}-p^2)\mathcal{G}d^2 - \frac{1}{2}\mathcal{G}d^2 + \frac{1}{2}\mathcal{G}d^2$$

v e sin (q+s)

$$+ 9p^2 - 9q^2 + 9R^2 - \frac{1}{2}(\frac{2}{m}-p^2)\mathcal{G}d^2 - \frac{1}{2}\mathcal{G}d^2 + \frac{1}{2}\mathcal{G}d^2$$

V 2

Deinde

Deinde posito

$$F' = \alpha F - A(F' - g') + \frac{1}{2}F' - \frac{1}{2}G'$$

$$G' = 3\alpha G - Af' - \frac{1}{2}F'$$

$$H' = (\alpha - 1)H - A'f' - D'g' + E'g' - \frac{1}{2}F' + \frac{1}{2}F'D' - \frac{1}{2}G' + \frac{1}{2}J'$$

$$J' = (\alpha + 1)J - A'f' + D'g' - E'g' - \frac{1}{2}F' + \frac{1}{2}F'D' - \frac{1}{2}G' + \frac{1}{2}H'$$

$$K' = (\alpha - 1)K - A'f' - Q'g' + R'g' + \frac{1}{2}F'(\frac{2}{m} - p') + \frac{1}{2}F'D' - \frac{1}{2}G' + \frac{1}{2}L'$$

$$L' = (\alpha + 1)L - A'f' + Q'g' - R'g' + \frac{1}{2}F'(\frac{2}{m} - p') + \frac{1}{2}F'D' - \frac{1}{2}G' + \frac{1}{2}K'$$

erit

$$\frac{d^2v}{dt^2} = -A'f \sin 2q - D'k \sin(2q - r) - P' \sin s - Q' \sin(2q - s)$$

$$- E'k \sin(2q + r) - R' \sin(2q + s)$$

$$- F'v \sin q - H'vk \sin(q - r) - K'v \sin(q - s)$$

$$- G'v \sin 3q - J'vk \sin(q + r) - L'v \sin(q + s)$$

§. 178. Hinc iam denovo differentiando conseque-

mur:

$$\frac{d^2v}{dt^2} = \text{Præc.}$$

$$+ v \cos q \quad [+A'f' + A'g' - \alpha F' + \frac{1}{2}F'D' + \frac{1}{2}G'D']$$

$$+ v \cos 3q \quad [+A'f' + \frac{1}{2}F'D' - 3\alpha G']$$

$$+ vk \cos(q - r) \quad [+A'f' + D'g' + E'g' + \frac{1}{2}F'D' + \frac{1}{2}F'D' + \frac{1}{2}G'D' - (\alpha - 1)H' + \frac{1}{2}J'D']$$

$$+ vk \cos(q + r) \quad [+A'f' + E'g' + D'g' + \frac{1}{2}F'D' + \frac{1}{2}F'D' + \frac{1}{2}G'D' + \frac{1}{2}H'D' - (\alpha + 1)J']$$

$$+ v \cos(q - s) \quad [+A'f' + Q'g' + R'g' - \frac{1}{2}(\frac{2}{m} - p')F' + \frac{1}{2}F'D' + \frac{1}{2}G'D' - (\alpha - \frac{1}{m})K' + \frac{1}{2}L'D']$$

$$+ v \cos(q + s) \quad [+A'f' + Q'g' + R'g' - \frac{1}{2}(\frac{2}{m} - p')F' + \frac{1}{2}F'D' + \frac{1}{2}G'D' - (\alpha - \frac{1}{m})K' + \frac{1}{2}L'D']$$

$+\frac{1}{2}J'D'$
 $+\frac{1}{2}H'D'$
 $+\frac{1}{2}L'D'$
 $+\frac{1}{2}K'D'$
 $(2q - s)$
 $(2q + s)$
 $(q - s)$
 $(q + s)$
 æque-
 $G'D'$
 $J'D'$
 $G'D'$
 J'
 $G'D'$
 $L'D'$
 $+$

$$- + v \cos(q + s) \left\{ \begin{array}{l} +A'f' + R'g' + Q'g' - \frac{1}{2}(\frac{2}{m} - p')F' + \frac{1}{2}F'D' + \frac{1}{2}G'D' \\ + \frac{1}{2}K'D' - (\alpha + \frac{1}{m})L' \end{array} \right.$$

qui valores cum antecedentibus comparari debent, vk inde valores coefficientium elicantur.

§. 179. Sumamus primo duos valores ab initio positos, quoniam hi a sequentibus non pendent, æque habebimus,

$$\bullet = \alpha \mathfrak{F} - \mathfrak{H}f' + \mathfrak{H}g' + \frac{1}{2}\mathfrak{F}' - \frac{1}{2}\mathfrak{G}' + \frac{3F - 3G}{2ms}$$

$$\bullet = 3\alpha \mathfrak{G} - \mathfrak{H}f' - \frac{1}{2}\mathfrak{F}' + \frac{3F}{2ms}$$

$$\bullet = \alpha F' - A'f' - A'g' - \frac{1}{2}F'D' - \frac{1}{2}G'D' + \frac{1}{2}\mathfrak{F}' - 2\mathfrak{F}\mathfrak{G} + \frac{3(F+G)}{4ms} + \frac{(3A+3A)(\mathfrak{F}+\mathfrak{G})}{ms} + \frac{(3A+3A)(F+G) + 3A}{ms}$$

$$\bullet = 3\alpha G' - A'f' - \frac{1}{2}F'D' + \frac{1}{2}\mathfrak{F}' - 6G - 2\mathfrak{F}\mathfrak{G} + \frac{3F}{4ms} + \frac{(3A+3A)\mathfrak{G}}{ms} + \frac{(3\mathfrak{H}+3A)F + 9A}{8ms}$$

$$\bullet = F' = \alpha F - A(f' - g') + \frac{1}{2}F'D' - \frac{1}{2}G'D'$$

$$G' = 3\alpha G - Af' - \frac{1}{2}F'D'$$

$$\text{at est } f' = \frac{2x F + \mathfrak{F}}{ms} \text{ et } g' = \frac{2x G + \mathfrak{G}}{ms}$$

Hic igitur primum litterarum, quae sunt cognitae, valores in numeris substituamus, critique

$$\begin{aligned}
 F' &= 0,93905 F + 0,00753 \mathcal{S} + 0,01443 G - 0,00753 \mathcal{Q} \\
 G' &= 2,80121 G + 0,02505 F + 0,00753 \mathcal{S} \\
 0 &= 0,92847 \mathcal{S} + 0,01776 F - 0,01776 G - 0,02501 \mathcal{Q} \\
 &\quad + 0,37500 \\
 0 &= 2,80121 \mathcal{Q} + 0,01447 \mathcal{S} + 0,01776 F + 1,87500 \\
 &\text{hincque} \\
 \mathcal{S} &= -0,01930 F + 0,01913 G - 0,42145 \\
 \mathcal{Q} &= -0,00623 F - 0,00010 G - 0,66717
 \end{aligned}$$

§. 180. Inde porro colligemus

$$\begin{aligned}
 F' &= 0,93895 F + 0,01457 G + 0,00185 \\
 G' &= 0,02491 F + 2,80135 G - 0,00317 \\
 f &= 0,01138 F + 0,00011 G - 0,00240 \\
 g' &= -0,00004 F + 0,01148 G - 0,00380 \\
 &\text{quibus valoribus substituitis pervenimus ad has aequa-} \\
 &\text{tiones:} \\
 0,09337 F &= + 0,05421 G + 1,96867 \\
 6,83135 G &= - 0,08830 F - 3,20944
 \end{aligned}$$

vnde fit:

$$\begin{aligned}
 F &= 20,65700 & \dots & \dots & 1 F &= 1,315067 \\
 G &= -0,73681 & \dots & \dots & 1-G &= 9,867356 \\
 \mathcal{S} &= -0,83418 & \dots & \dots & 1-\mathcal{S} &= 9,921260 \\
 \mathcal{Q} &= -0,79580 & \dots & \dots & 1-\mathcal{Q} &= 9,900804 \\
 F' &= +19,38712 & \dots & \dots & 1-F' &= 1,287512 \\
 G' &= -1,55266 & \dots & \dots & 1-G' &= 0,191075 \\
 f &= +0,23259 & \dots & \dots & 1-f &= 9,366501 \\
 g' &= -0,01309 & \dots & \dots & 1-g' &= 8,116940
 \end{aligned}$$

§. 181.

§. 181. His valoribus, qui ad inaequalitates abfolutas pertinent, expeditis, progrediamur ad eos, qui ab excentricitate orbitae lunaris pendent, ac his aequationibus continentur:

$$\begin{aligned}
 (\alpha-1) \mathcal{S} &- \mathcal{Q} \mathcal{S}' + \frac{1}{2} \mathcal{Q} \mathcal{S}' - \mathcal{D} f' + \mathcal{Q} g' - \frac{1}{2} \mathcal{S} (\mathcal{S}' - \mathcal{D}') - \frac{1}{2} \mathcal{Q} \mathcal{S}' \\
 &+ \frac{1}{2} \mathcal{S}' + \frac{3 \mathcal{J}}{2 m} + \frac{3(F-G)}{m} = 0 \\
 (\alpha+1) \mathcal{S} &- \mathcal{Q} \mathcal{H}' + \frac{1}{2} \mathcal{S} \mathcal{H}' + \mathcal{D} g' - \mathcal{Q} f' - \frac{1}{2} \mathcal{S} (\mathcal{S}' - \mathcal{D}') - \frac{1}{2} \mathcal{Q} \mathcal{S}' \\
 &+ \frac{1}{2} \mathcal{S}' + \frac{3 \mathcal{H}}{2 m} + \frac{3(F-G)}{m} = 0
 \end{aligned}$$

$$\begin{aligned}
 H' &= (\alpha-1) H - A \mathcal{H}' + \frac{1}{2} \mathcal{J} \mathcal{H}' - D f' + E g' - \frac{1}{2} F (\mathcal{S}' - \mathcal{D}') - \frac{1}{2} G \mathcal{S}' \\
 J' &= (\alpha+1) J - A \mathcal{H}' + \frac{1}{2} H \mathcal{H}' + D g' - E f' - \frac{1}{2} F (\mathcal{S}' - \mathcal{D}') - \frac{1}{2} G \mathcal{S}' \\
 (\alpha-1) H' &- A \mathcal{H}' - \frac{1}{2} J' \mathcal{H}' - D f' - E g' - \frac{1}{2} F' (\mathcal{S}' + \mathcal{D}') - \frac{1}{2} G \mathcal{S}'
 \end{aligned}$$

$$\begin{aligned}
 &+ \frac{1}{2} \mathcal{S}' - E \mathcal{H} - 2 \mathcal{K} \mathcal{S}' + \frac{1}{2} b F + \frac{3 \mathcal{J}}{4 m} + \frac{3(F+G)}{2 m} + \frac{3 \mathcal{Q}}{m} + \frac{3 \mathcal{Q}}{m} \\
 &+ \frac{\mathcal{Q} + 3 A}{m} \mathcal{S}' + \frac{3 \mathcal{Q} + 3 A}{m} \mathcal{J}' + \frac{\mathcal{D} + 3 D}{m} \mathcal{S}' + \frac{3 \mathcal{D} + 3 D}{m} F \\
 &+ \frac{\mathcal{Q} + 3 E}{m} \mathcal{S}' + \frac{3 \mathcal{Q} + 3 E}{m} G - \frac{3 A(F+G)}{2 m} + \frac{9 D + 9 E}{8 m} = 0
 \end{aligned}$$

$$\begin{aligned}
 (\alpha+1) J' &- A \mathcal{H}' - \frac{1}{2} H \mathcal{H}' - E f' - D g' - \frac{1}{2} F' (\mathcal{S}' + \mathcal{D}') - \frac{1}{2} G \mathcal{S}' \\
 &+ \frac{1}{2} \mathcal{S}' - E \mathcal{J} - 2 \mathcal{K} \mathcal{J}' + \frac{1}{2} b F + \frac{3 \mathcal{H}}{4 m} + \frac{3(F+G)}{2 m} + \frac{3 \mathcal{Q}}{m} + \frac{3 \mathcal{Q}}{m} \\
 &+ \frac{\mathcal{Q} + 3 A}{m} \mathcal{S}' + \frac{3 \mathcal{Q} + 3 A}{m} H + \frac{\mathcal{D} + 3 D}{m} \mathcal{S}' + \frac{3 \mathcal{D} + 3 D}{m} G \\
 &+ \frac{\mathcal{Q} + 3 E}{m} \mathcal{S}' + \frac{3 \mathcal{Q} + 3 E}{m} F - \frac{3 A(F+G)}{2 m} + \frac{9 E + 9 D}{8 m} = 0
 \end{aligned}$$

§. 182.

aequa-

181.

$$\begin{aligned}
 K &= 0,85830K + 0,00527L + 0,01447M + 1,48970 \\
 &+ 0,00750\mathfrak{E} - 0,00750\mathfrak{M} \\
 L &= 1,00916L + 0,00527K + 0,01447N + 1,55183 \\
 &+ 0,00750\mathfrak{F} - 0,00750\mathfrak{M} \\
 M &= 2,72528M + 0,02501K - 1,17408 \\
 &+ 0,00750\mathfrak{F} \\
 N &= 2,87666N + 0,02501L - 0,95901 \\
 &+ 0,00750\mathfrak{E}
 \end{aligned}$$

$\frac{ddv}{d^2s}$ = Præc.

$$\begin{aligned}
 +vecl(q+s) &= \left[\begin{array}{l} -0,85830K - 0,02843L - 0,02843M - 0,32674 \\ -0,00987L - 0,01409\mathfrak{E} - 0,01409\mathfrak{M} \\ -0,02961M \end{array} \right] \\
 +vecl(q+s) &= \left[\begin{array}{l} -1,00918L - 0,02843K - 0,02843N - 0,56620 \\ -0,00987K - 0,01409\mathfrak{F} - 0,01409\mathfrak{M} \\ -0,02981N \end{array} \right] \\
 +vecl(3q+s) &= \left[\begin{array}{l} -2,72578M - 0,02843K + 0,74683 \\ -0,00987K - 0,01409\mathfrak{F} \end{array} \right] \\
 +vecl(3q+s) &= \left[\begin{array}{l} -2,87666N - 0,02853L + 0,67486 \\ -0,00987L - 0,01409\mathfrak{E} \end{array} \right]
 \end{aligned}$$

§. 192. Valores autem litterarum commate nota-
tarum hic substituentur dabunt

$$\frac{ddv}{d^2s} = \left[\begin{array}{l} -0,73747K - 0,04264L - 0,12156M - 0,00014N \\ -0,00029\mathfrak{F} - 0,02053\mathfrak{E} - 0,00705\mathfrak{M} + 0,00008\mathfrak{M} \\ -1,58590 \end{array} \right]$$

18970
5183

$$\begin{aligned}
 +vecl(q+s) &= \left[\begin{array}{l} -1,01923L - 0,04222K - 0,12821N - 0,00014M \\ -0,00029\mathfrak{E} - 0,02166\mathfrak{F} - 0,00642\mathfrak{M} + 0,00008\mathfrak{M} \\ -2,11858 \end{array} \right] \\
 +vecl(3q+s) &= \left[\begin{array}{l} -7,443058M - 0,09507K - 0,000005L + 3,93257 \\ +0,000008\mathfrak{M} - 0,03453\mathfrak{F} - 0,000007\mathfrak{E} \end{array} \right] \\
 +vecl(3q+s) &= \left[\begin{array}{l} -8,27529N - 0,11938L - 0,000005K + 3,41830 \\ +0,00008\mathfrak{M} - 0,03566\mathfrak{E} - 0,000007\mathfrak{F} \end{array} \right]
 \end{aligned}$$

574
520

§. 193. His expressionibus ita evolutis atque ad
calculum numericum præparatis, quaeramus easdem
expressiones ex formulis supra traditis pro R et $\frac{ddv}{d^2s}$
quae continentur in §. 52 et 54. Inde autem omitten-
dis terminis, quos iam traſtaurimus, consequemur.

$$\begin{aligned}
 R &= Pr. + v \sin(q+s) \left(-\frac{1}{2} + \frac{3L}{2ms} - \frac{3M}{2ms} - \frac{9F}{4ms} + \frac{9G}{4ms} \right) \\
 &+ v \sin(q+s) \left(-\frac{1}{2} + \frac{3K}{2ms} - \frac{3N}{2ms} + \frac{9G}{4ms} - \frac{9F}{4ms} \right) \\
 &+ v \sin(3q+s) \left(-\frac{1}{2} + \frac{3K}{2ms} - \frac{3N}{2ms} + \frac{9G}{4ms} - \frac{9F}{4ms} \right) \\
 &+ v \sin(3q+s) \left(-\frac{1}{2} + \frac{3L}{2ms} - \frac{9F}{4ms} \right)
 \end{aligned}$$

10ca-
197
ve

ddv

$$\frac{ddv}{dv} = \text{Præc.}$$

$$+ve \text{ cof}(q-s) \left[\begin{aligned} & -\frac{2}{4} + \frac{9P}{8ms} + \frac{9Q}{8ms} + \frac{5R}{8ms} - \frac{6K}{4ms} + \frac{3L}{4ms} + \frac{3M}{4ms} \\ & - \frac{3F}{4ms} - \frac{9G}{8ms} - \frac{2K}{8ms} + \frac{(2l+3A)}{ms} (R+9D) \\ & + \frac{(32l+3A)}{ms} (L+M) + \frac{(9l+3P)}{ms} G + \frac{(39l+3P)}{ms} F \\ & + \frac{(2l+3Q)}{ms} G + \frac{(32l+3Q)}{ms} F + \frac{(3+3R)}{ms} G \\ & + \frac{(3R+3K)}{ms} G \end{aligned} \right]$$

$$+ve \text{ cof}(q+s) \left[\begin{aligned} & -\frac{2}{4} + \frac{9P}{8ms} + \frac{9R}{8ms} + \frac{5Q}{8ms} - \frac{6L}{4ms} + \frac{3K}{4ms} + \frac{3N}{4ms} \\ & - \frac{3F}{4ms} - \frac{9G}{8ms} - \frac{2K}{8ms} + \frac{(2l+3A)}{ms} (R+9D) \\ & + \frac{(32l+3A)}{ms} (K+N) + \frac{(9l+3P)}{ms} G + \frac{(39l+3P)}{ms} F \\ & + \frac{(2l+3Q)}{ms} G + \frac{(32l+3Q)}{ms} F + \frac{(3R+3K)}{ms} G \\ & + \frac{(3R+3K)}{ms} F \end{aligned} \right]$$

$$+ve \text{ cof}(3q-s) \left[\begin{aligned} & -\frac{2}{4} + \frac{5P}{8ms} + \frac{9Q}{8ms} - \frac{6M}{4ms} + \frac{3K}{4ms} - \frac{3G}{4ms} \\ & - \frac{9F}{8ms} - \frac{2K}{8ms} + \frac{(2l+3A)}{ms} G + \frac{(39l+3P)}{ms} G \\ & + \frac{(39l+3P)}{ms} G + \frac{(2l+3Q)}{ms} G + \frac{(32l+3Q)}{ms} F \end{aligned} \right]$$

+ve

$$+ve \text{ cof}(3q+s) \left[\begin{aligned} & -\frac{2}{4} + \frac{5P}{8ms} + \frac{9R}{8ms} - \frac{6N}{4ms} + \frac{3L}{4ms} + \frac{3G}{4ms} + \frac{9F}{8ms} \\ & + \frac{(2l+3A)}{ms} G + \frac{(32l+3A)}{ms} L + \frac{(9l+3P)}{ms} G + \frac{(39l+3P)}{ms} G \\ & + \frac{(3R+3K)}{ms} G + \frac{(3R+3K)}{ms} F \end{aligned} \right]$$

§. 194. Introducantur hic quoque valores inaeogniti, ac prodibit

$$R = \text{Pr. } +ve \text{ fin}(q-s) [-1,02394 + 0,00854L - 0,00854M] \\ +ve \text{ fin}(q+s) [-1,02394 + 0,00854K - 0,00854N] \\ -ve \text{ fin}(3q-s) [-4,01450 + 0,00854K] \\ -ve \text{ fin}(3q+s) [-4,01450 + 0,00854L]$$

$$\frac{ddv}{dv} = \text{Præc.}$$

$$+ve \text{ cof}(q-s) \left[\begin{aligned} & -1,01591K - 0,03207L - 0,03207M - 1,16867I \\ & -2,01798F - 0,02711R - 0,027119D \end{aligned} \right] \\ +ve \text{ cof}(q+s) \left[\begin{aligned} & -1,01591L - 0,03207K - 0,03207N - 1,03904 \\ & -2,01798R - 0,02711R - 0,027119D \end{aligned} \right] \\ +ve \text{ cof}(3q-s) \left[\begin{aligned} & -1,01591M - 0,03207K - 2,49545 \\ & -2,017989D - 0,02711R \end{aligned} \right] \\ +ve \text{ cof}(3q+s) \left[\begin{aligned} & -1,01591N - 0,03207L - 2,78450 \\ & -2,017989D - 0,02711R \end{aligned} \right]$$

Y

§. 194.

§. 195. Hinc ergo odo sequentes aequationes resultabunt

L 0,85830 S = + 0,00526 Z - 0,02500 M + 1,39986 - 0,01785 L + 0,01785 M

II. 1,00918 Z = + 0,00526 S - 0,02500 M + 1,36509 - 0,01785 K + 0,01785 N

III. 2,72578 M = - 0,01448 S + 4,50673 - 0,01785 K

IV. 2,87666 M = - 0,01448 Z + 4,48566 - 0,01785 L

V. 1,027844 K - 0,01057 L - 0,08949 M - 0,00014 N + 0,00081 = 12,01769 S + 0,00658 Z - 0,01946 M + 0,00008 M

VI. 0,00332 L - 0,01015 K - 0,09614 N - 0,00014 M - 0,17954 = 12,01769 S + 0,00545 Z + 0,02069 M + 0,00008 M

VII. -6,41487 M - 0,06300 K - 0,00005 L + 6,42802 = + 2,01806 M - 0,00742 S - 0,00007 Z

VIII. -7,25938 N - 0,08131 L - 0,00005 K + 6,20280 = + 2,01806 M - 0,00855 Z - 0,00007 S

§. 196. Ex aequationibus III et IV faciam efficitur: tur hi valores

M = - 0,00531 S - 0,0055 Z + 1,65336
N = - 0,00503 Z - 0,00621 L + 1,55933

qui

equationes

1,39986

1,36509

0,0081 =

7954 =

102 =

80 =

efficitur:

qui

qui in I et II substituit praebent:

0,85817 S = 1,35853 + 0,00526 Z - 0,01785 (L-M) + 0,00016 K
1,00905 Z = 1,32611 + 0,00526 S - 0,01785 (K-N) + 0,00015 L

unde obtineatur:

S = + 1,59116 - 0,02080 (L-M) + 0,00008 K + 0,00010 N
Z = + 1,32251 - 0,01769 (K-N) + 0,00005 L + 0,00011 M
M = + 1,64491 - 0,00655 K + 0,00010 (L-M)
N = + 1,55268 - 0,00621 L + 0,00009 (K-N)

§. 197. His valoribus substituitis caeterae aequationes adhibent in formas sequentes:

0,27862 K - 0,05252 L - 0,04754 M + 0,00017 N + 3,35313 =
- 0,00346 L - 0,04584 K - 0,06045 N + 0,00020 M + 2,53002 =
- 6,41502 M - 0,07622 K + 0,00090 L + 9,73587 =
- 7,25971 N - 0,09384 L + 0,00028 K + 9,32499 =

ex quarum binis postremis faciam obtineatur:

M = - 0,01188 K + 0,00005 L + 1,51765
N = - 0,01293 L + 0,00004 K + 1,28448

unde colligitur:

+ 0,27918 K - 0,05252 L + 3,28120 =
- 0,00268 L - 0,04584 K + 2,45267 =

ac demique

K = + 40,44710 . . . / K = 1,606887
L = + 368,40200 . . . / L = 2,566322
M = + 1,05555 . . . / M = 0,023478
N = - 3,47730 . . . / N = 0,541242

§. 197.

§. 198. Literarum germanicarum valores hinc erunt:

R	=	5,04677	· · · · ·	I-R	=	0,703013
Q	=	0,56302	· · · · ·	I-Q	=	9,750524
M	=	1,41672	· · · · ·	I-M	=	0,151283
N	=	0,73119	· · · · ·	I-N	=	9,864030

ac literarum hinc derivatarum:

K	=	0,43578	· · · · ·	I-K	=	9,639267
L	=	4,23401	· · · · ·	I-L	=	0,626752
M'	=	0,02018	· · · · ·	I-M'	=	8,304921
N'	=	0,04409	· · · · ·	I-N'	=	8,644340

§. 199. Nunc igitur intelligimus inaequalitates ab angulis 3^r-s et 3ⁿ+s pendentes tam esse parvas, ut sine vilo errore reici queant, etiam si valores K et L aliquantum immutauerint. Diffancia ergo lunae curata a terra $x = \frac{(1-k)ax}{1-k \cos r}$ ita ab his inaequalitatibus parallactis pendebit, ut sit

# = Praec.	+	0,11756	v	col 7	Log. coeff.	9,070249
		0,00419	v	col 3 ⁿ		7,622540
		0,1234	w	col (q-r)		9,091265
		0,0784	w	col (q+r)		8,894183
	+	0,2302	w	col (q-s)		9,362071
	+	2,0965	w	col (q+s)		0,321506

Morus

Morus autem momentaneus ita hinc afficietur, ut sit

$\frac{d\theta}{dt}$ = Praec.	—	0,23259	v	col 7	9,366591
		0,01309	v	col 3 ⁿ	8,116940
	+	0,0939	w	col (q-r)	8,972943
	+	0,1653	w	col (q+r)	9,218352
	—	0,4358	w	col (q-s)	9,639267
	—	4,2340	w	col (q+s)	0,626752

§. 200. Quod si iam ipsam longitudinem lunae, quatenus ab his inaequalitatibus parallactis pendet, ponamus:

$$\begin{aligned} \phi &= \text{Praec.} + S^r \sin q + S^s \sin (q-r) + S^w \sin (q-s) \\ &+ S^v \sin 3q + S^x \sin (q+r) + S^y \sin (q+s) \end{aligned}$$

sequentes obtinebimus aequationes pro horum coefficientium determinatione:

$$\begin{aligned} -0,23259 &= a S^r - S^w \sin q - \frac{1}{2} S^x \sin 2q - \frac{1}{2} S^y \sin 2s \\ + 0,01309 &= 3 S^v - S^x \sin 2q - \frac{1}{2} S^y \sin 2s \\ + 0,0939 &= (a-1) S^r - S^w \sin q - \frac{1}{2} S^x \sin 2q - \frac{1}{2} S^y \sin 2s \\ + 0,1653 &= (a+1) S^r - S^w \sin q - \frac{1}{2} S^x \sin 2q - \frac{1}{2} S^y \sin 2s \\ - 0,4358 &= -\frac{1}{2} S^x \sin 2q - \frac{1}{2} S^y \sin 2s \\ - 4,2340 &= -\frac{1}{2} S^x \sin 2q - \frac{1}{2} S^y \sin 2s \end{aligned}$$

	Log. coeff.	coeff. integri	quo $\frac{d\theta}{dr}$
# = 1	0,00074991	0,0007499	
+	0,0000532	0,000053	
+	0,01915577	0,010430	
+	0,003293	0,000179	
+	0,003321	0,000181	
+	0,000049	0,000005	
+	0,00511	0,000015	
+	0,08022	0,000238	
+	0,00237	0,000007	
+	0,07892	0,000234	
+	0,000001	0,000000	
+	0,006400	0,000206	
+	0,014801	0,000249	
+	0,011415	0,000192	
+	0,00364	0,000001	
+	0,01482	0,000004	
+	0,00584	0,000001	
+	0,37957	0,000347	
+	0,30653	0,000281	
+	0,36337	0,000332	
+	0,00093	0,000009	
+	0,07752	0,000071	
+	0,00305	0,000003	
+	0,011756	0,000408	
+	0,00419	0,000015	
+	0,1234	0,000024	
+	0,0784	0,000015	
+	0,2302	0,000013	
+	0,0965	0,000122	

Hic ad laevis adiunxi valores coefficientium integrorum in numeris absolutis expressis, ponendo $k = 0,05445$, $r = 0,01680$ et $v = 1r$; quos proinde, si hi valores aliter per observationes determinentur, facile erit emendare.

\$. 204.

	Log. coeff.	coeff. integri	quo $\frac{d\theta}{dr}$
# = 1	1,009176	1,009176	
+	0,0195144	0,019514	
+	0,0000322	0,000032	
+	0,001231	0,000067	
+	0,366103	0,019934	
+	0,012832	0,000099	
+	0,002829	0,000154	
+	0,000171	0,000009	
+	0,01182	0,000035	
+	0,02057	0,000061	
+	0,01063	0,000032	
+	0,09883	0,000293	
+	0,00004	0,000000	
+	0,013760	0,000231	
+	0,037487	0,000530	
+	0,03062	0,000505	
+	0,00722	0,000002	
+	0,03470	0,000010	
+	0,01533	0,000005	
+	0,75204	0,000688	
+	0,62626	0,000573	
+	0,69420	0,000635	
+	0,02440	0,000022	
+	0,12160	0,000111	
+	0,02000	0,000018	
+	0,23259	0,0000808	
+	0,01309	0,000045	
+	0,0939	0,000018	
+	0,1653	0,000031	
+	0,4358	0,000025	
+	4,2340	0,000247	

\$. 204. Pro motu autem lunae momentaneo, ex quo eius motus horarius definiti poterit, habebimus:

Z

\$. 105.

§. 207. Si iam longitudo Lunae per solam excentricitatem secundum regulas Keplerianas determinata ponatur $\xi = C$ sita vt postea eius anomalia vera $= r$, futurum sit $\xi = C + 1,0085272 r$, erit longitudo vera per haecenus inuentas inaequalitates.

$\phi = \xi$	Log.coeff.	VAl.coeff.	in min. fec.
+ 0,0103597 fin 2 η	8,0153347	—	+ 21377/1
— 0,0000382 fin 4 η	5,582063	—	8
+ 0,0101466 fin r	8,006295	+ 114	—
— 0,4202266 fin (2 η - r)	9,623483	- 4720	—
+ 0,0049926 fin (2 η + r)	7,698261	+ 56	—
+ 0,0052866 fin (4 η - r)	7,723163	+ 59	—
— 0,0000866 fin (4 η + r)	5,935307	—	—
+ 0,0042066 fin 2 r	7,623250	+ 21	—
+ 0,5732866 fin (2 η -2 r)	9,758367	+ 351	—
+ 0,0031866 fin (2 η +2 r)	7,402427	—	—
— 0,1508366 fin (4 η -2 r)	9,178488	+ 14	—
— 0,0000266 fin (4 η +2 r)	5,301030	—	—
+ 0,2013856 fin r	9,304026	+ 701	—
— 0,0218896 fin (2 η - s)	8,340237	—	—
— 0,0169686 fin (2 η + s)	8,214002	—	—
+ 0,0661566 fin 2 s	8,820508	+ 57	—
+ 0,0233266 fin (2 η -2 s)	8,367825	+ 4	—
+ 0,0084066 fin (2 η +2 s)	7,924429	+ 1	—
+ 0,7476066 fin (r - s)	9,873165	+ 141	—
— 0,6185066 fin (r + s)	9,791317	—	—
+ 0,8143066 fin (2 η - r + s)	9,910800	+ 118	—
— 0,0142066 fin (2 η + r - s)	8,150690	—	—
+ 0,2396066 fin (2 η - r - s)	9,379550	+ 154	—
— 0,0061066 fin (2 η + r + s)	7,788910	—	—
— 0,244227 η	9,387868	+ 45	—
+ 0,00639 η	7,805991	—	—
+ 1,1959 η fin (4 η - r)	0,077694	+ 175	—
+ 0,0757 η fin (4 η + r)	8,879096	+ 4	—
— 0,3939 η fin (4 η - s)	9,597508	+ 59	—
— 4,1738 η fin (4 η + s)	0,620530	+ 4	—
		5	—
		49	—

CAPUT

II

tricia- onatur $\xi = C$ uenas	Log.coeff.	in min. fec.
— 21377/1	—	—
8	—	—
— 114	—	—
4720	—	—
56	—	—
59	—	—
1	—	—
21	—	—
351	—	—
14	—	—
92	—	—
0	—	—
701	—	—
76	—	—
57	—	—
4	—	—
1	—	—
141	—	—
118	—	—
154	—	—
3	—	—
45	—	—
1	—	—
175	—	—
4	—	—
59	—	—
4	—	—
5	—	—
49	—	—

CAPUT

CAPUT XII.

INVESTIGATIO INAEQUALITATUM MOTUM LINEAE NODORUM AFRICIENTIUM.

§. 206.

A nequam reliquas motus Lunae inaequalitates, quae ab inclinatione eius orbitae ad eclipticam pendit, desinit licet, cum variationes, quae in motu lineae nodorum Lunae, cum eas, quae in ipsa inclinatione eius orbitae ad eclipticam deprehenduntur, inuestigari oportet. Residua enim pars aequationis notatae principalis, qua omnes motus Lunae inaequalitates continentur, litteras π et ρ implicat, quarum illa longitudinem nodi ascendentis, haec vero ρ inclinationem ad eclipticam designat. Nisi igitur utriusque huius quantitatis incrementa vel decrementa ad differentiale $d\pi$ reduxerimus, residuas motus Lunae inaequalitates determinare non poterimus.

207. Aequatio autem supra (§5) pro motu lineae nodorum tradita, cum sit $\frac{2\pi\eta + \sqrt{R}d\eta}{n\pi} = a + \frac{1+2e}{n\pi} \frac{d\phi}{d\eta}$

$$\text{ideoque } \frac{2\pi\eta + \sqrt{R}d\eta}{n\pi} = a + \frac{1+2e}{n\pi} \frac{d\phi}{d\eta}$$

$$+ e^2 \cos e + g^2 \cos(2\eta - s)$$

$$+ d^2 \cos(2\eta + r)$$

$$+ d^2 \cos(2\eta + s)$$

§. 2.

Si

Si ponamus breviskatis grata $\frac{3(1+2kk+3r^2)}{kk\pi\pi} = i$, ut sit $i = 0,0168918$, inducet formam sequentem:

$$\frac{d\pi}{dr} = -1 \left(k + \omega \cos 2\pi - \left(\frac{2}{\pi} - r^2 \right) k \cos r + \frac{d}{dr} k \cos(2\pi - r) + \text{etc.} \right)$$

$$\left(1 + k \cos r + k k \cos 2r - 6 k \cos(r - r) \right) \left(\frac{3}{4} \frac{1}{\pi} \cos 2\pi - \frac{1}{4} \cos(2\theta - 2\pi) \right)$$

$$- \frac{k r^2}{4} \left(\frac{1}{2} \cos r + \frac{1}{2} \cos 3r - \frac{1}{2} \cos(\theta + \theta - 2\pi) - \frac{1}{2} \cos(3\theta - \theta - 2\pi) \right)$$

vbi quidem plurimi termini tam sunt parvi, ut facile negligi queant.

§. 208. Productum autem ex duobus prioribus factoribus, quoniam id in formula pro inclinatione recurrit, seorsim exhibebamus: fiet id autem reiectis terminis, qui prae reliquis admodum sunt parvi ut sequitur:

$-k i + 2 i \left(\frac{2}{\pi} - r^2 \right) k k - \frac{1}{2} i r^2 r^2$	$- 0,017043$
$\cos 2\pi \quad (-i r^2 - 2 i k k k k$	$+ 0,000161 \cos 2\pi$
$k \cos r \quad \left(i \left(\frac{2}{\pi} - r^2 \right) - 4 k i \right)$	$- 0,068110 k \cos r$
$k \cos(2\pi - r) \quad (-i r^2 - 2 i k k$	$- 0,005791 k \cos(2\pi - r)$
$k \cos(2\pi + r) \quad (-i r^2 - 2 i k k$	$+ 0,000828 k \cos(2\pi + r)$
$k k \cos 2r \quad (-5 k i + 2 i \left(\frac{2}{\pi} - r^2 \right))$	$- 0,085091 k k \cos 2r$
$e \cos r \quad (-i r^2 + 3 k i$	$+ 0,051362 e \cos r$
$e \cos(2\pi - r) \quad (i r^2 + \frac{1}{2} k k$	$- 0,000929 e \cos(2\pi - r)$
$e \cos(2\pi + r) \quad (-i r^2 + \frac{1}{2} k k$	$- 0,000802 e \cos(2\pi + r)$
$e e \cos 2r \quad (-\frac{1}{2} k i + \frac{1}{2} i r^2$	$- 0,025914 e e \cos 2r$

§. 209.

§. 209. His valoribus substituentis prodibit

$\frac{d\pi}{dr} = \frac{0,004261 + 0,000020}{1}$	
$\cos 2\pi \quad (+0,0000040 - 0,0094261$	
$k \cos r \quad (-0,017043 - 0,000620$	
$k \cos(2\pi - r) \quad (-0,001448 - 0,008514) - 0,010636 k k \cos(2\pi - 2r)$	
$k \cos(2\pi + r) \quad (+0,000207 - 0,008514) - 0,010636 k k \cos(2\pi + 2r)$	
$k k \cos 2r \quad (-0,021273$	
$e \cos r \quad (+0,012842 - 0,000217$	
$e \cos(2\pi - r) \quad (-0,000232 + 0,0006420) - 0,003239 e e \cos(2\pi - 2r)$	
$e \cos(2\pi + r) \quad (-0,000201 + 0,0006420) - 0,003239 e e \cos(2\pi + 2r)$	
$e e \cos 2r \quad (-0,0006479)$	
$\cos 2(\theta - \pi) \quad (+0,003261 - 0,000020) - 0,000041 \cos \theta$	
$\quad + 0,000019 \cos(3\theta - \theta - 2\pi)$	
$\cos 2(\theta - \pi) \quad (+0,004261 - 0,000020) - 0,000019 \cos 3\theta$	
$\quad + 0,000019 \cos(3\theta - \theta - 2\pi)$	
$k \cos(2\theta - 2\pi + r) \quad (+0,0008514 - 0,000103)$	
$\quad + 0,000022 \cos(\theta + \theta - 2\pi)$	
$k \cos(2\theta - 2\pi + r) \quad (+0,0008514 - 0,000103)$	
$\quad + 0,010636 k k \cos 2(\theta - \theta - \pi)$	
$k \cos(2\theta - 2\pi - r) \quad (+0,0008514 - 0,000103)$	
$\quad + 0,000022 \cos(\theta - \theta - \pi)$	
$k \cos(2\theta - 2\pi - r) \quad (+0,0008514 - 0,000103)$	
$\quad + 0,000022 \cos(\theta - \theta - \pi)$	
$e \cos(2\theta - 2\pi + r) \quad (-0,0006420 + 0,000100)$	
$\quad + 0,0003239 e e \cos 2(\theta - \theta - \pi)$	
$e \cos(2\theta - 2\pi + r) \quad (-0,0006420 + 0,000100)$	
$\quad + 0,0003239 e e \cos 2(\theta - \theta - \pi)$	

§. 110.

§. 210. Habebimus ergo

$$\begin{aligned} \frac{ds}{dt} &= -0,004241 & -0,000041 \cos \varphi \\ -0,004221 \cos 2\varphi & & -0,000019 \cos 3\varphi \\ -0,017663 k \cos r & & +0,000020 \cos 4\varphi \\ -0,009962 k \cos(2\varphi-r) & & +0,004241 \cos(2\varphi-2\pi) \\ -0,008307 k \cos(2\varphi+r) & & +0,004241 \cos(2\theta-2\pi) \\ -0,010636 k \cos(2\varphi-2r) & & +0,000022 \cos(\varphi+\theta-2\pi) \\ -0,010636 k \cos(2\varphi+2r) & & +0,000019 \cos(3\varphi-\theta-2\pi) \\ -0,021273 k \cos r & & +0,000019 \cos(3\theta-\varphi-2\pi) \\ +0,012623 e \cos r & & \\ +0,006188 e \cos(2\varphi-r) & & \\ +0,006219 e \cos(2\varphi+r) & & \\ -0,006479 ee \cos 2s & & \\ -0,003239 ee \cos(2\varphi-2s) & & \\ +0,009238 k \cos(2\varphi-2\pi-r) & & -0,006304 e \cos(2\varphi-2\pi-r) \\ +0,008411 k \cos(2\varphi-2\pi+r) & & -0,006320 e \cos(2\varphi-2\pi+r) \\ +0,008411 k \cos(2\theta-2\pi-r) & & -0,006320 e \cos(2\theta-2\pi-r) \\ +0,009238 k \cos(2\theta-2\pi+r) & & -0,006304 e \cos(2\theta-2\pi+r) \\ +0,010636 k \cos(2\varphi-2\pi-2r) & & +0,003239 ee \cos(2\theta-2\pi-2s) \end{aligned}$$

§. 211. Quoniam plurimi horum terminorum tam sunt parvi, ut in se speſſati tuto reſciſſi poſſent, tamen quidam per integrationem ad magnitudinem fatis notabilem excreſcere poſſunt. Huius autem indolis ſunt illi termini, qui eiſusmodi complectuntur angulos, quorum

rum differentialia ad dt admodum parvam tenent rationem, cuiusmodi ſunt anguli r , $2r$, $2\theta-2\pi$, $2\theta-2\pi-r$, $2\theta-2\pi+r$, $2\varphi-2\pi-2r$ et $2\theta-2\pi-2r$; quorum natura differentialium ex ſequentibus formulis colligi poſſeſt:

$$\begin{aligned} \frac{dr}{dt} &= a - a' \cos 2\varphi - c'k \cos r - d'k \cos(2\varphi-r) - e'k \cos(2\varphi+r) \\ &+ \left(\frac{2}{n} - p'\right) e \cos r - q' e \cos(2\varphi-r) - r' e \cos(2\varphi+r) \\ \frac{d\theta}{dt} &= a + \frac{1+2ee}{n} - a' \cos 2\varphi + \left(\frac{2}{n} - c'\right) k \cos r \\ &- d'k \cos(2\varphi-r) - p'e \cos r - q'e \cos(2\varphi-r) \\ &- c'k \cos(2\varphi+r) - r'e \cos(2\varphi+r) \\ \frac{ds}{dt} &= \frac{d\theta}{dt} = \frac{1+2ee}{n} + \frac{2}{n} k \cos r - \frac{2}{n} e \cos r \end{aligned}$$

§. 212. Quoniam primo inaequalitates motus nodorum, quae neque ab excentricitate orbitae lunaris neque ſolaris pendunt, ſiguntur:

$$\begin{aligned} \pi &= \text{Conſt.} - Or + \mathfrak{N} \sin 2\varphi + \mathfrak{B} \sin(2\varphi-2\pi) + \mathfrak{C} \sin(2\theta-2\pi) \\ &+ \text{reſtiſ reliquis terminis, quos praecedimus fore minimos, ac differendiando obtinebimus:} \\ \frac{d\pi}{dt} &= -O - \mathfrak{N}' a' - 0,004241 \mathfrak{B} - 0,004241 \mathfrak{C} \\ &+ \cos 2\varphi (2a\mathfrak{N}' - 0,004241 \mathfrak{B} - 0,004241 \mathfrak{C}) \\ &+ \cos(2\varphi-2\pi) (2(a+\frac{1}{n})\mathfrak{B} + 0,008482 \mathfrak{B} + 0,004221 \mathfrak{C}) \\ &+ \cos(2\theta-2\pi) (-\mathfrak{B}' a' + 0,004221 \mathfrak{B} + \frac{2\mathfrak{C}}{n} + 0,008482 \mathfrak{C}) \end{aligned}$$

vide

vnde oritur :

$$\begin{aligned}
O &= 0,0019744 \mathfrak{N} + 0,004241 \mathfrak{B} + 0,004241 \mathfrak{C} = 0,004241 \\
1,867476 \mathfrak{N} &= 0,004241 \mathfrak{B} - 0,004241 \mathfrak{C} = -0,004241 \\
2,026834 \mathfrak{B} &+ 0,004221 \mathfrak{C} = 0,004241 \\
0,023965 \mathfrak{B} &+ 0,159358 \mathfrak{C} = 0,004241
\end{aligned}$$

§. 213. Valores hinc igitur prodibunt sequentes :

$$\begin{aligned}
O &= + 0,004078 \quad \dots \quad / O = 7,610447 \\
\mathfrak{N} &= - 0,002196 \quad \dots \quad / -\mathfrak{N} = 7,341634 \\
\mathfrak{B} &= + 0,002037 \quad \dots \quad / \mathfrak{B} = 7,308991 \\
\mathfrak{C} &= + 0,026307 \quad \dots \quad / \mathfrak{C} = 8,420081
\end{aligned}$$

Vbi primum obleruo valorem ipsius O iam proxime accedere ad motum medium iunae nodorum, vi per observationes constat; inde enim esse deberet $O = 0,004053$ facile autem intelligitur, huic exiguum defectum per reliquas inaequalitates suppleri posse. Quocirca hinc erit

$\pi = \text{Conf.} = 0,004078 \pi$		Valores in min. sec.
$\quad = 0,002196 \sin 2\eta$	$\quad = 453''$	
$\quad + 0,002037 \sin (2\phi - 2\pi)$	$\quad + 420$	
$\quad + 0,026307 \sin (2\theta - 2\pi)$	$\quad + 5426$	

quae inaequalitates mirifice conveniunt cum observationibus. His addi potest terminus :

$$+ 0,000336 \sin (4\theta - 4\pi)$$

cuius in minutis secundis valor est $-1' 69''$, qui terminus autem posterius illo facile corrigi potest.

§. 214. Quaevis iam ferisim. inaequalitates, quae ab excentricitate orbitae lunaris pendunt, sitque

$$\pi =$$

$$\begin{aligned}
\pi &= \text{Conf.} - O + \mathfrak{N} \sin 2\eta + \mathfrak{B} \sin (2\phi - 2\pi) + \mathfrak{C} \sin (2\theta - 2\pi) \\
&+ \mathfrak{D} k \sin (2\eta - \pi) + \mathfrak{E} k \sin (2\eta + \pi) + \mathfrak{F} k \sin \pi \\
&+ \mathfrak{G} k k \sin (2\eta - 2\pi) + \mathfrak{H} k k \sin 2\pi + \mathfrak{I} k \sin (2\phi - 2\pi - \pi) \\
&+ \mathfrak{K} k \sin (2\phi - 2\pi + \pi) + \mathfrak{L} k \sin (2\theta - 2\pi - \pi) + \mathfrak{M} k \sin (2\theta - 2\pi + \pi) \\
&+ \mathfrak{N} k k \sin (2\phi - 2\pi - 2\pi)
\end{aligned}$$

eritque differentiendo

$$\begin{aligned}
\frac{d\pi}{dt} &= \text{Pr.} + k \cos (2\eta - \pi) (-\mathfrak{N}' - 0,009238 \mathfrak{B}' - 0,009238 \mathfrak{C}' + (2\pi - 1) \mathfrak{D}' \\
&+ k \cos (2\eta + \pi) (-\mathfrak{N}' - 0,008411 \mathfrak{B}' - 0,008411 \mathfrak{C}' + (2\pi + 1) \mathfrak{E}' \\
&+ k \cos \pi (-\mathfrak{N}' - \mathfrak{N}' - 0,017649 \mathfrak{B}' - 0,017649 \mathfrak{C}' + \mathfrak{F}' - \mathfrak{F}' \\
&+ k k \cos 2\pi \quad (-0,010636 \mathfrak{B}' + 2 \mathfrak{H}' - \mathfrak{D}' - \mathfrak{D}' - \mathfrak{G}' \\
&+ k k \cos (2\eta - 2\pi) (-0,010636 \mathfrak{C}' + 2 (\mathfrak{I}' - 1) \mathfrak{G}' - \mathfrak{D}'
\end{aligned}$$

$+ k \cos (2\theta - 2\pi - \pi)$	$- \mathfrak{B}' + 0,008307 \mathfrak{B}' + \frac{2}{\pi} \mathfrak{C}' + 0,017663 \mathfrak{C}'$
$+ k \cos (2\theta - 2\pi + \pi)$	$+ (\frac{2}{\pi} - 1) \mathfrak{L}' + 0,008482 \mathfrak{L}'$
$+ k \cos (2\phi - 2\pi - \pi)$	$- \mathfrak{B}' + 0,009962 \mathfrak{B}' + \frac{2}{\pi} \mathfrak{C}' + 0,017663 \mathfrak{C}'$
$+ k \cos (2\phi - 2\pi + \pi)$	$+ (\frac{2}{\pi} + 1) \mathfrak{M}' + 0,008482 \mathfrak{M}'$
$+ k \cos (2\phi - 2\pi - \pi)$	$+ 0,017663 \mathfrak{B}' + 0,009962 \mathfrak{C}' + 2 (\mathfrak{A}' + \frac{1}{\pi}) \mathfrak{B}'$
$+ k \cos (2\phi - 2\pi + \pi)$	$- \mathfrak{B}' + 0,008482 \mathfrak{B}'$
$+ k \cos (2\phi - 2\pi - \pi)$	$+ 0,017663 \mathfrak{B}' + 0,008307 \mathfrak{C}' + 2 (\mathfrak{A}' + \frac{1}{\pi}) \mathfrak{B}'$
$+ k \cos (2\phi - 2\pi + \pi)$	$+ \mathfrak{B}' + 0,008482 \mathfrak{B}'$
$+ k k \cos (2\phi - 2\pi - 2\pi)$	$+ 0,021273 \mathfrak{B}' + 0,010636 \mathfrak{C}' + 2 (\mathfrak{A}' + \frac{1}{\pi}) \mathfrak{B}'$
	$- 2 \mathfrak{N}' + 0,008482 \mathfrak{N}'$

A a

§. 215.

§. 215. Superfluum foret maiorem curam in his differentialibus adhibere, quia vero proxime tantum rem determinare sufficit; erit ergo:

$$0,867476 \mathcal{D} = 0,009962$$

$$2,867476 \mathcal{E} = 0,008307$$

Hincque

erit in min fec.

$$\mathcal{D} = 0,011480 \quad | \quad \mathcal{D} k = 129''$$

$$\mathcal{E} = 0,002900 \quad | \quad \mathcal{E} k = 33''$$

quae inaequalitates in loco nodi vix alicuius sunt momenti, vnde eas exactius determinare non est opus.

§. 216. Calculo autem evolutio erit

$\pi =$ Pr. - 0,011480k $\sin(2\eta - r)$	8,059940	-129'
- 0,002900k $\sin(2\eta + r)$	7,462400	- 33
- 0,017663k $\sin r$	8,247064	-198
+ 0,090497kk $\sin(2\eta - 2r)$	8,956634	+ 55
- 0,011978kk $\sin 2r$	8,078384	- 7
+ 0,008707k $\sin(2\phi - 2\pi - r)$	7,939851	+ 98
+ 0,002701k $\sin(2\phi - 2\pi + r)$	7,431516	+ 30
- 0,004680k $\sin(2\theta - 2\pi - r)$	7,670224	- 53
+ 0,004685k $\sin(2\theta - 2\pi + r)$	7,670680	+ 53
+ 0,384848kk $\sin(2\phi - 2\pi - 2r)$	9,585229	+235

§. 217. Simili modo inuestigemus inaequalitates motus nodorum, quae pendent ab excentricitate orbitae solaris hincque:

$\pi =$

$$\pi = \text{Cont.} - Or + \mathcal{M} \sin 2\eta + \mathcal{N} \sin(2\phi - 2\pi) + \mathcal{O} \sin(2\theta - 2\pi)$$

$$+ \mathcal{D} \sin s + \mathcal{E} \sin(2\eta - s) + \mathcal{F} \sin(2\eta + s)$$

$$+ \mathcal{G} \sin 2s + \mathcal{H} \sin(2\phi - 2\pi - s) + \mathcal{I} \sin(2\theta - 2\pi - s)$$

$$+ \mathcal{J} \sin(2\phi - 2\pi + s) + \mathcal{K} \sin(2\theta - 2\pi + s)$$

$$+ \mathcal{L} \sin(2\theta - 2\pi - 2s)$$

vnde differentiando pro terminis quaevis erit:

$$\frac{d\pi}{ds} = \text{Praec.}$$

$$+ e \cos s \left\{ -\mathcal{M}g' - \mathcal{M}' + 0,006304\mathcal{N} + 0,006320\mathcal{O} - \frac{1}{\pi} \mathcal{D} - \mathcal{E}g', \mathcal{F}g' \right.$$

$$\left. + 0,006320\mathcal{J} + 0,006304\mathcal{K} \right\} + 0,006304\mathcal{E} + (2a - \frac{1}{\pi}) \mathcal{E}$$

$$+ e \cos(2\eta - s) \left\{ \mathcal{M}(\frac{2}{\pi} - p') + 0,006320\mathcal{N} + 0,006320\mathcal{O} + (2a + \frac{1}{\pi}) \mathcal{F} \right.$$

$$+ e \cos 2s \left\{ -0,003239 \mathcal{E} - \frac{1}{\pi} \mathcal{D} - \mathcal{E}g' - \mathcal{F}g' + \frac{2}{\pi} \mathcal{G} \right.$$

$$\left. + \mathcal{H}p' - 0,012623 \mathcal{N} - 0,006188 \mathcal{E} \right.$$

$$\left. + (2a \pm \frac{1}{\pi} + 0,008482) \mathcal{H} \right.$$

$$+ e \cos(2\phi - 2\pi + s) \left\{ - \mathcal{N}p' - 0,012623 \mathcal{N} - 0,006219 \mathcal{E} \right.$$

$$\left. + (2a + \frac{3}{\pi} + 0,008482) \mathcal{J} \right.$$

A a 2

+

$$\begin{aligned}
 & + r \cos(2\theta - 2\pi - s) \left\{ - 837 - 0,00621938 - \frac{2}{n} \mathcal{C} - 0,012623 \mathcal{C} \right. \\
 & \quad \left. + \left(\frac{1}{n} + 0,008482 \right) \mathcal{R} \right. \\
 & + r \cos(2\theta - 2\pi + s) \left\{ - 837 - 0,00618883 - \frac{2}{n} \mathcal{C} - 0,012623 \mathcal{C} \right. \\
 & \quad \left. + \left(\frac{2}{n} + 0,008482 \right) \mathcal{R} \right. \\
 & + r \cos(2\theta - 2\pi - 2s) \left\{ + 0,003239 \mathcal{B} + \frac{1}{2n} \mathcal{C} + 0,006479 \mathcal{C} \right. \\
 & \quad \left. + 0,008482 \mathcal{M} - \frac{1}{n} \mathcal{R} - 0,012623 \mathcal{R} \right.
 \end{aligned}$$

§. 218. Hinc reperuntur sequentes valores

- $\mathcal{D} = 0,159070$; $\mathcal{D} = 9,201585$; $\mathcal{D} = 5514$
- $\mathcal{E} = 0,003562$; $\mathcal{E} = 7,551680$; $\mathcal{E} = 12\frac{1}{2}$
- $\mathcal{F} = 0,003301$; $\mathcal{F} = 7,518677$; $\mathcal{F} = 11\frac{1}{2}$
- $\mathcal{G} = 0,031650$; $\mathcal{G} = 8,587191$; $\mathcal{G} = 211$
- $\mathcal{H} = -0,003153$; $\mathcal{H} = 7,498692$; $\mathcal{H} = 111$
- $\mathcal{I} = -0,002932$; $\mathcal{I} = 7,467118$; $\mathcal{I} = 101$
- $\mathcal{K} = -0,025730$; $\mathcal{K} = 8,410784$; $\mathcal{K} = 90$
- $\mathcal{L} = -0,009076$; $\mathcal{L} = 7,957885$; $\mathcal{L} = 32$

At valor ipsius \mathcal{M} tam sit parvus, vt merito pro nihilo haberi possit.

§. 219.

§. 219. Colligamus ergo has inaequalitates in vnam summam, atque obtinebimus longitudinem veram nodi ascendentis

$\pi = \text{Conf.} - 0,004053 \quad r$

	Valor in minut. sec.
$- 0,002196 \sin 2r$	$- 453''$
$+ 0,002037 \sin(2\theta - 2\pi)$	$+ 420$
$+ 0,026307 \sin(2\theta - 2\pi)$	$+ 5426$
$+ 0,000370 \sin(4\theta - 4\pi)$	$+ 75$
$- 0,017666 \sin r$	$- 198$
$- 0,011486 \sin(2r - r)$	$- 129$
$- 0,002906 \sin(2r + r)$	$- 33$
$+ 0,090566 \sin(2r - 2r)$	$+ 55$
$- 0,012066 \sin 2r$	$- 7$
$+ 0,0008716 \sin(2\theta - 2\pi - r)$	$+ 98$
$+ 0,002706 \sin(2\theta - 2\pi + r)$	$+ 30$
$- 0,0004686 \sin(2\theta - 2\pi - r)$	$- 53$
$+ 0,0004686 \sin(2\theta - 2\pi + r)$	$+ 53$
$+ 0,384866 \sin(2\theta - 2\pi - 2r)$	$+ 235$
$+ 0,159076 \sin s$	$+ 551$
$- 0,025756 \sin(2\theta - 2\pi - s)$	$- 90$
$- 0,009076 \sin(2\theta - 2\pi + s)$	$- 32$

omissis scilicet his inaequalitatibus, quae non supra 30' exurgunt

A a 3

CAPUT

C A P U T XIII.

INVESTIGATIO INCLINATIONIS ORBITAE LUNARIS AD ECLIPTICAM.

§ 220.

Pro inclinatione orbitae lunaris ad eclipticam invenienda, forma §. 208. euoluta multiplicari debet per $-\frac{1}{2} \sin 2\eta + \frac{1}{2} \sin 2(\Phi - \pi) + \frac{1}{2} \sin 2(\theta - \pi)$, ac productum erit $= \frac{d \text{ tang } \ell}{d r}$: Hinc ergo habebitur:

$$\begin{aligned} \frac{d \text{ tang } \ell}{d r} &= +0,004261 \sin 2\eta && -0,000020 \sin 4\eta \\ &+0,008514k \sin(2\eta - r) \\ &+0,008514k \sin(2\eta + r) \\ &+0,000827k \sin r \\ &-0,004261 \sin(2\Phi - 2\pi) \text{ adice} \\ &-0,004261 \sin(2\theta - 2\pi) \text{ adcoeff.} \\ &-0,008514k \sin(2\Phi - 2\pi - r) - 0,000724 \\ &-0,008514k \sin(2\Phi - 2\pi + r) + 0,000103 \\ &-0,008514k \sin(2\theta - 2\pi - r) + 0,000103 \\ &-0,008514k \sin(2\theta - 2\pi + r) - 0,000724 \\ &-0,010636kk \sin(2\Phi - 2\pi - 2r) \\ &+0,006420 e \sin(2\theta - 2\pi - r) - 0,0000100 \\ &+0,006420 e \sin(2\theta - 2\pi + r) - 0,0000116 \end{aligned}$$

§. 221.

C A P U T XIII.

§. 221. Queramus primo terminos, qui a neutra excentricitate pendent, sique

$$\frac{d \text{ tang } \ell}{\text{tang } e} = \dots$$

$2\eta \cos 2\eta + \cos 4\eta + 2\mathcal{B} \cos(2\Phi - 2\pi) + \mathcal{C} \cos(2\theta - 2\pi) + \cos(4\theta - 4\pi)$ erique differentiando:

$$\frac{d \text{ tang } \ell}{d r} = \sin 2 \left[-2e\mathcal{B} + 0,004241 \mathcal{B} - 0,004241 \mathcal{C} \right.$$

$$\left. \sin 4\eta \left[-4ea + \frac{2}{n} a' \right] \right.$$

$$\left. \sin(2\Phi - 2\pi) \left\{ -2\left(\omega + \frac{1}{n}\right) \mathcal{B} - 0,008482 \mathcal{B} - 0,004227 \mathcal{C} \right\} \right.$$

$$\left. \sin(2\theta - 2\pi) \left\{ -\frac{2}{n} \mathcal{C} + 2a' - 0,004221 \mathcal{B} - 0,008482 \mathcal{C} \right\} \right.$$

$$\left. \sin(4\theta - 4\pi) \left\{ +0,004241 \mathcal{C} - \frac{4}{n} e \right\} \right.$$

§. 222. Ex his iam reperitur:

$$\begin{aligned} \mathcal{B} &= -0,002630 && \cdot && \cdot && l - \mathcal{B} = 7,419914 \\ \mathcal{C} &= +0,002037 && \cdot && \cdot && l \mathcal{B} = 7,308991 \\ \mathcal{C} &= +0,026307 && \cdot && \cdot && l \mathcal{C} = 8,420081 \\ a &= +0,000019 && \cdot && \cdot && l a = 5,278753 \\ e &= +0,000370 && \cdot && \cdot && l e = 6,567931 \end{aligned}$$

ita vt hinc fit:

$$\begin{aligned} \frac{d \text{ tang } \ell}{\text{tang } e} &= -0,002630 \cos 2\eta \\ &+0,000019 \cos 4\eta \\ &+0,002037 \cos(2\Phi - 2\pi) \\ &+0,026307 \cos(2\theta - 2\pi) \\ &+0,000370 \cos(4\theta - 4\pi) \end{aligned}$$

§. 223.

m inue-
i: debet
: $(\theta - \pi)$,
ebitur:
220 sin 4η
E
24
03
03
24
00
16
§. 221.

§. 223. Quateramus iam seorsim terminos ab excentricitate Lunae pendentes: sicque

$$\frac{1}{\text{tang}^2} = \mathcal{M} \cos 2\eta + \mathcal{N} \cos(2\phi - 2\pi) + \mathcal{O} \cos(2\theta - 2\pi) + \mathcal{D} \cos(2\eta - \eta) + \mathcal{E} \cos(2\eta + \eta) + \mathcal{F} \cos \phi + \mathcal{G} \cos(2\phi - 2\pi - \eta) + \mathcal{H} \cos \cos(2\theta - 2\pi - \eta) + \mathcal{I} \cos \cos(2\theta - 2\pi + \eta) + \mathcal{K} \cos(2\phi - 2\pi + \eta) + \mathcal{L} \cos \cos(2\phi - 2\pi - 2\eta)$$

unde differentialibus sumendis habebitur: $\frac{d}{d\eta} \frac{1}{\text{tang}^2} =$

$$\begin{aligned} \mathcal{M} \sin(2\eta - \eta) & \left\{ + \mathcal{M}' + 0,00923823 - 0,00923826 + 0,00424115 \right. \\ & \left. + (2\pi - 1) \mathcal{D} - 0,00424115 \right. \\ \mathcal{N} \sin(2\eta + \eta) & \left\{ + \mathcal{M}' + 0,00084112 - 0,00084116 + (2\pi + 1) \mathcal{E} \right. \\ & \left. + 0,00424115 - 0,00424115 \right. \\ \mathcal{O} \sin \eta & \left\{ + \mathcal{M}' - \mathcal{M}' + 0,000082723 - 0,00082726 - \mathcal{F} - 0,00424115 \right. \\ & \left. + 0,00424115 - \mathcal{D}' + \mathcal{O}' - 0,00424115 + 0,00424115 \right. \end{aligned}$$

$$\mathcal{M} \sin(2\phi - 2\pi - \eta) \left\{ -0,01766323 - 0,0099626 - 2\left(\pi + \frac{1}{\pi}\right) \mathcal{G} + \mathcal{O} - 0,00848215 - 0,00422115 \right.$$

$$\mathcal{M} \sin(2\phi - 2\pi + \eta) \left\{ -0,01766323 - 0,0083076 - 2\left(\pi + \frac{1}{\pi}\right) \mathcal{H} - \mathcal{I} - 0,00848215 - 0,00422115 \right.$$

$$\mathcal{M} \sin(2\theta - 2\pi - \eta) \left\{ + \mathcal{M}' - 0,00830723 - \frac{2}{\pi} \mathcal{E} - 0,01766323 + \mathcal{O}' - 0,00422115 - \frac{2}{\pi} \mathcal{G} + \mathcal{H} - 0,00848215 \right.$$

k sin

$$\mathcal{M} \sin(2\theta - 2\pi + \eta) \left\{ + \mathcal{M}' - 0,0099626 - \frac{2}{\pi} \mathcal{E} - 0,01766323 + \mathcal{H}' - 0,00422115 - \frac{2}{\pi} \mathcal{H} - \mathcal{H} - 0,00848215 \right.$$

$$\mathcal{M} \sin(2\phi - 2\pi - 2\eta) \left\{ -0,02127323 - 0,0106366 - 2\left(\pi + \frac{1}{\pi}\right) \mathcal{I} + 2\mathcal{L} - 0,00848215 \right.$$

hincque reperitur:

$\mathcal{D} =$	0,010487	...	$\mathcal{D} =$	8,020638
$\mathcal{E} =$	0,003166	...	$\mathcal{E} =$	7,500439
$\mathcal{F} =$	0,001600	...	$\mathcal{F} =$	7,204170
$\mathcal{G} =$	0,008719	...	$\mathcal{G} =$	7,940484
$\mathcal{H} =$	0,002699	...	$\mathcal{H} =$	7,431136
$\mathcal{I} =$	0,004460	...	$\mathcal{I} =$	7,649305
$\mathcal{L} =$	0,004717	...	$\mathcal{L} =$	7,623628
$\mathcal{I} =$	0,384890	...	$\mathcal{I} =$	9,583335

§. 224. Nunc denique pro inaequalitibus ab excentricitate orbis solaris pendens ponatur.

$$\frac{1}{\text{tang}^2} = \mathcal{M} \cos 2\eta + \mathcal{N} \cos(2\phi - 2\pi) + \mathcal{O} \cos(2\theta - 2\pi - \eta) + \mathcal{E} \cos(2\theta - 2\pi) + \mathcal{F} \cos \cos(2\theta - 2\pi + \eta)$$

ac differentiendo prodibit: $\frac{d}{d\eta} \frac{1}{\text{tang}^2} =$

$$\mathcal{M} \sin(2\eta - \eta) \left\{ + \mathcal{M}' + 0,00621923 + \frac{2}{\pi} \mathcal{E} + 0,01262323 - \frac{1}{\pi} \mathcal{M}' - 0,00848215 \right.$$

Bb

k sin

$$\left. \begin{aligned} & \sin(2\theta - 2\pi + s) \\ & + 98^{\frac{1}{2}} + 0,06188 \mathcal{B} + \frac{2}{\pi} \mathcal{C} + 0,012623 \mathcal{C} \\ & - \frac{3}{\pi} \mathcal{D} - 0,008482 \mathcal{D} \end{aligned} \right\}$$

unde reperitur

$$\mathcal{D} = -0,024034 \dots L\mathcal{D} = 8,380835$$

$$\mathcal{D} = -0,008519 \dots L\mathcal{D} = 7,930332$$

§. 225. Si ergo ϵ denotet inclinationem medianam orbitae lunaris ad eclipticam, et φ inclinationem veram, erit

$\frac{\text{tang } \varphi}{\text{tang } \epsilon} =$	$0,002630 \text{ cof } 2\varphi$	7,419915
	$+ 0,000019 \text{ cof } 4\varphi$	5,278753
	$+ 0,002037 \text{ cof}(2\theta - 2\pi)$	7,308991
	$+ 0,026307 \text{ cof}(2\theta - 2\pi)$	8,420081
	$+ 0,000370 \text{ cof}(4\theta - 4\pi)$	6,567931
	$+ 0,010496 \text{ cof}(2\varphi + \varphi)$	8,020638
	$+ 0,003176 \text{ cof}(2\varphi + \varphi)$	7,500439
	$- 0,001606 \text{ cof } \varphi$	7,2004120
	$+ 0,008726 \text{ cof}(2\theta - 2\pi - \varphi)$	7,940484
	$+ 0,002706 \text{ cof}(2\theta - 2\pi + \varphi)$	7,431136
	$- 0,004466 \text{ cof}(2\theta - 2\pi - \varphi)$	7,649305
	$+ 0,004726 \text{ cof}(2\theta - 2\pi + \varphi)$	7,673628
	$+ 0,384966 \text{ cof}(2\theta - 2\pi - 2\varphi)$	9,585335
	$- 0,02403 \epsilon \text{ cof}(2\theta - 2\pi - \varphi)$	8,380835
	$- 0,00852 \epsilon \text{ cof}(2\theta - 2\pi + \varphi)$	7,930332

log. coeff.

§. 226.

§. 226. Quodsi iam ponatur $\frac{\text{tang } \varphi}{\text{tang } \epsilon} = S$, erit ad

numeros ipsos procedendo $\frac{\text{tang } \varphi}{\text{tang } \epsilon} = 1 + S + \frac{1}{2} S S$

Hinc igitur negligendo terminos minimos, consequemur:

$\frac{\text{tang } \varphi}{\text{tang } \epsilon} =$	$1 + 0,002604 \text{ cof } 2\varphi$
	$+ 0,000020 \text{ cof } 4\varphi$
	$+ 0,002003 \text{ cof}(2\theta - 2\pi)$
	$+ 0,026307 \text{ cof}(2\theta - 2\pi)$
	$+ 0,000490 \text{ cof}(4\theta - 4\pi)$
	$- 0,001606 \text{ cof } \varphi$
	$+ 0,010496 \text{ cof}(2\varphi - \varphi)$
	$+ 0,003176 \text{ cof}(2\varphi + \varphi)$
	$+ 0,008856 \text{ cof}(2\theta - 2\pi - \varphi)$
	$+ 0,004726 \text{ cof}(2\theta - 2\pi + \varphi)$
	$- 0,004466 \text{ cof}(2\theta - 2\pi - \varphi)$
	$+ 0,004726 \text{ cof}(2\theta - 2\pi + \varphi)$
	$+ 0,384966 \text{ cof}(2\theta - 2\pi - 2\varphi)$
	$- 0,02403 \epsilon \text{ cof}(2\theta - 2\pi - \varphi)$
	$- 0,00852 \epsilon \text{ cof}(2\theta - 2\pi + \varphi)$

Bb 2

§. 227.

§. 227. Cum in aequatione nostra principali, quae motum Lunae continet, infra terminus $\frac{\text{tang } \rho^2}{\text{tang } \epsilon^2}$, huius- quoque valorem euolui conueniet: erit ergo

$$\frac{\text{tang } \rho^2}{\text{tang } \epsilon^2} = 1$$

$$\begin{aligned} & - 0,005154 \text{ cof } 2\eta \quad + 0,000040 \text{ cof } 4\eta \\ & + 0,003938 \text{ cof } (2\phi - 2\pi) \\ & + 0,052614 \text{ cof } (2\theta - 2\pi) \quad + 0,001320 \text{ cof } (4\theta - 4\pi) \\ & - 0,002906 \text{ cof } \nu \\ & + 0,020986 \text{ cof } (2\eta - \nu) \\ & + 0,006346 \text{ cof } (2\eta + \nu) \\ & + 0,017966 \text{ cof } (2\phi - 2\pi - \nu) \\ & + 0,005356 \text{ cof } (2\phi - 2\pi + \nu) \\ & - 0,008966 \text{ cof } (2\theta - 2\pi - \epsilon) \\ & + 0,009408 \text{ cof } (2\theta - 2\pi + \epsilon) \\ & + 0,769866 \text{ cof } (2\phi - 2\pi - 2\nu) \\ & + 0,048806 \text{ cof } (2\theta - 2\pi - \nu) \\ & + 0,017046 \text{ cof } (2\theta - 2\pi + \nu) \end{aligned}$$

Hicque ergo valor in superiori illa aequatione substitui poterit.

§. 228.

quae huius-

§. 228. Celeb. autem Clairaut concludit inclinatio- nem mediam ϵ ex observationibus exquisitissimis $5^\circ 8' 9''$, ex qua igitur ad quoduis tempus inclinationem veram elicere licebit. Sit enim $\rho = \epsilon + \omega$, erit tang $\rho = \frac{\text{tg } \epsilon + \omega}{1 - \omega \text{ tg } \epsilon}$ = tang $\epsilon + \frac{\omega}{\text{cof } \epsilon^2}$ = V tang ϵ , ponendo V pro expres- sione ipsius $\frac{\text{tang } \rho}{\text{tang } \epsilon}$. Hinc erit $\omega = (V-1) \text{ fin } \epsilon \text{ cof } \epsilon = \frac{1}{2} (V-1) \text{ fin } 2\epsilon = 0,08915 (V-1)$: vnde reperitur in minutis secundis

$$\rho = \epsilon +$$

+	48''	cof 2 η
+	36	cof (2 $\phi - 2\pi$)
+	484	cof (2 $\theta - 2\pi$)
+	9	cof (4 $\theta - 4\pi$)
-	2	cof ν
+	11	cof (2 $\eta - \nu$)
+	3	cof (2 $\eta + \nu$)
+	9	cof (2 $\phi - 2\pi - \nu$)
+	3	cof (2 $\phi - 2\pi + \nu$)
-	5	cof (2 $\theta - 2\pi - \epsilon$)
+	5	cof (2 $\theta - 2\pi + \epsilon$)
+	23	cof (2 $\phi - 2\pi - 2\nu$)
-	7	cof (2 $\theta - 2\pi - \nu$)
-	3	cof (2 $\theta - 2\pi + \nu$)

Bb 3

§. 229.

§. 229. Hic notandum est, etiam si valor inclinationis mediae & aliquantillum immutetur, aequationes has tamen inde vix alterari, ita ut eae semper eadem sint manfure. Perfpicuum quoque est in calculo astro-nomico sufficere tres inaequalitates primores, et reliquas omnes sine errore sensibili praetermitti posse; nisi forte aequatio 23 col ($2\phi - 2\pi - 2r$) retinenda censetur, quae inter reliquas est maxima. Exprimunt autem angulus $2\phi - 2\pi - 2r$ duplani distantiam apogei Lunae ab eius nodo, a quo angulo quoque locum nodi non medicrter affici vidimus, cum correctio hinc oriunda pro loco nodi vsque ad 235'' adfurgere possit.

CAPUT

CAPUT XIV.

INVESTIGATIO INAEQUALITATUM MOTUS
LUNAE AB EIU8 INCLINATIONE AD ELLIPTICAM ORIUNDARUM.

§. 230.

Ponamus more adhuc vitato :

$$\begin{aligned}
 \sqrt{Rd} = & \mathcal{M} \cos 2\eta + \mathcal{N} \cos r + \mathcal{O} k \cos(2\eta - r) + \mathcal{P} \sin \epsilon \cos r + \mathcal{Q} e \cos(2\eta - r) \\
 & + \mathcal{R} k \cos(2\eta + r) + \mathcal{S} e \cos(2\eta + r) \\
 & + \mathcal{T} f \cos 2\eta + \mathcal{U} f \cos(2\phi - 2\pi) + \mathcal{V} f k \cos r + \mathcal{W} f \cos(2\eta - r) \\
 & + \mathcal{X} f \cos(2\phi - 2\pi) + \mathcal{Y} f \cos(2\theta - 2\pi) \\
 & + \mathcal{Z} f k \cos(2\phi - 2\pi - r) + \mathcal{A} f k \cos(2\theta - 2\pi - r) \\
 & + \mathcal{B} f k \cos(2\phi - 2\pi + r) + \mathcal{C} f k \cos(2\theta - 2\pi + r) \\
 & + \mathcal{D} f k \cos(2\phi - 2\pi - 2r) + \mathcal{E} f \cos(2\phi - 2\pi - r) \\
 & + \mathcal{F} f k \cos(2\phi - 2\pi - 2r) + \mathcal{G} f k \cos(2\theta - 2\pi - r) \\
 & + \mathcal{H} f \cos(2\theta - 2\pi) \\
 & + \mathcal{I} f k \cos(2\phi - 2\pi - r) + \mathcal{J} f k \cos(2\theta - 2\pi - r) \\
 & + \mathcal{K} e \cos(2\eta + r) \\
 & + \mathcal{L} f k \cos(2\eta + r) \\
 & + \mathcal{M} f k \cos(2\phi - 2\pi - r) + \mathcal{N} f k \cos(2\theta - 2\pi - r) \\
 & + \mathcal{O} f k \cos(2\phi - 2\pi + r) + \mathcal{P} f k \cos(2\theta - 2\pi + r) \\
 & + \mathcal{Q} f k \cos(2\phi - 2\pi - 2r) + \mathcal{R} f e \cos(2\theta - 2\pi - r) \\
 & + \mathcal{S} f e \cos(2\theta - 2\pi + r) \\
 & + \mathcal{T} f k \cos(2\phi - 2\pi + r) + \mathcal{U} f e \cos(2\theta - 2\pi - r) \\
 & + \mathcal{V} f e \cos(2\theta - 2\pi + r)
 \end{aligned}$$

§. 231.

PUT

§. 231. His valoribus substitutis in formula §. 52.

orientur

$$R = f \sin 2\gamma \left(\dots \dots \dots f_k \sin(2\theta - 2\pi - \gamma) \left(-\frac{3M}{2m} - \frac{3G}{m} \right) \right)$$

$$f \sin(2\phi - 2\pi) \left(+\frac{3H}{2m} \dots \dots f_k \sin(2\theta - 2\pi + \gamma) \left(-\frac{3N}{2m} - \frac{3G}{m} \right) \right)$$

$$f \sin(2\theta - 2\pi) \left(-\frac{3G}{2m} \dots \dots f_k \sin(2\theta - 2\pi - \gamma) \left(+\frac{9G}{4m} \right) \right)$$

$$f_k \sin \gamma \left(+\frac{3K}{2m} - \frac{3L}{2m} \dots \dots f_k \sin(2\theta - 2\pi + \gamma) \left(+\frac{9G}{4m} \right) \right)$$

$$f_k \sin(2\gamma - \gamma) \left(+\frac{3I}{2m} \right)$$

$$f_k \sin(2\gamma - \gamma) \left(+\frac{3I}{2m} \right)$$

$$f_k \sin(2\phi - 2\pi - \gamma) \left(+\frac{3S}{2m} + \frac{3H}{m} \right)$$

$$f_k \sin(2\phi - 2\pi + \gamma) \left(+\frac{3T}{2m} + \frac{3H}{m} \right)$$

$$f_k \sin(2\phi - 2\pi - 2\gamma) \left(+\frac{3S}{m} \right)$$

§. 232. Altera vero aequatio fundamentalis inducet formam sequentem :

$$\frac{d^2 u}{dt^2} = P \cos \dots + f \cos 2\gamma [-EF - 2\pi G - 0,005156 - 0,026307 + f \cos(2\phi - 2\pi) [-EG - 2\pi H + 0,003938 - 1 + f \cos(3\theta - 2\pi) [-GH - 2\pi I + 0,052614 + 0,002578 +$$

+

in §. 52.

$$\frac{3M}{2m} - \frac{3G}{m}$$

$$\frac{3N}{2m} - \frac{3G}{m}$$

$$\frac{9G}{4m}$$

$$\frac{9G}{4m}$$

$$+ f_k \cos \gamma [-EJ - 2\pi G - 0,00290 - 0,00893 - 0,000098 - 0,00278 - 0,00098 - 0,00258 - 0,01315 + f_k \cos(2\gamma - \gamma) [-EK + \frac{1}{2} b F - 2\pi R + 0,02098 - 0,00470 - 0,00258 - 0,01315 + f_k \cos(2\gamma + \gamma) [-EL + \frac{1}{2} b F - 2\pi Q + 0,00634 + 0,00448 - 0,00258 - 0,01315 + f_k \cos(2\phi - 2\pi - \gamma) [-EM + \frac{1}{2} b G - 2\pi M + 0,01796 + 0,00145 + 0,00197 - \frac{1}{2} + f_k \cos(2\phi - 2\pi + \gamma) [-EN + \frac{1}{2} b G - 2\pi N + 0,00556 + 0,00145 + 0,00197 - \frac{1}{2} + f_k \cos(2\phi - 2\pi - \gamma) [-EO + \frac{1}{2} b M + \frac{1}{2} G - 2\pi O + 0,7698 + 0,0089 + 0,0010 + 0,0007 - \frac{1}{2} + f_k \cos(2\theta - 2\pi - \gamma) [-ES + \frac{1}{2} b H - 2\pi S - 0,00896 - 0,00317 + 0,02631 + 0,00129 + f_k \cos(2\theta - 2\pi + \gamma) [-ET + \frac{1}{2} b H - 2\pi T + 0,00940 - 0,01049 + 0,02631 + 0,00129 + f_k \cos(2\theta - 2\pi - \gamma) [-EU - 2\pi U - 0,04806 + f_k \cos(2\theta - 2\pi + \gamma) [-EV - 2\pi V - 0,01704$$

induct

026307

578

+

§. 233. Quoniam manifestum est, coefficientes F, G, H etc. admodum fore paruos; cum maxime hiis generis inaequalitas alliquot minuta prima non excedat, hi hinc e- coefficientes per π divisi cum eridentur parvi, ut sine errore reici queant. Hoc autem factu quoque hincrae generantiae S, Q, R etc. pro nihilo esse habenda, ex quo sola posterior aequatio differentio-differentialis reformanda superent; in qua ob eandem rationem, terminos ex divisione coefficientium per π oriundos omnino, cum in eam operoso calculo eliminantur.

Cc

ciac

ciat correctiones inde reluctantes proximè saltem determinasse; præsertim cum hæc prætermisso vix ad aliquot minuta secunda sit ascensura.

§. 234. Ob eandem rationem licebit in valoribus

differentialium $\frac{d\Phi}{dt}$ et $\frac{d\eta}{dt}$ particulas ab inclinatione pen-

dentes negligere, unde erit: $\frac{d\psi}{dt} =$

$$f \sin 2\eta [-2aF = -F\eta]$$

$$f \sin (2\phi - 2\pi) [-2(a + \frac{1}{n})G - 0,008482 G = -G\eta]$$

$$f \sin (2\theta - 2\pi) [+G'd - \frac{2}{n}H - 0,008482 H = -H\eta]$$

$$f \sin \nu [F'd - F'd - J = -J\eta]$$

$$f \sin (2\eta - \nu) [+F'd - (2\eta - 1)K = -K\eta]$$

$$f \sin (2\eta + \nu) [+F'd - (2\eta + 1)L = -L\eta]$$

$$f \sin (2\phi - 2\pi - \nu) [-2(a + \frac{1}{n})M + M - 0,008482 M = -M\eta]$$

$$f \sin (2\phi - 2\pi + \nu) [-2(a + \frac{1}{n})N - N - 0,008482 N = -N\eta]$$

$$f \sin (2\phi - 2\pi - \nu) [-2(a + \frac{1}{n})O + 2O - 0,008482 O = -O\eta]$$

$$f \sin (2\theta - 2\pi - \nu) [+G'd - \frac{2}{n}H - \frac{2}{n}S + S - 0,008482 S = -S\eta]$$

$$f \sin (2\theta - 2\pi + \nu) [+G'd - \frac{2}{n}H - \frac{2}{n}T - T - 0,008482 T = -T\eta]$$

$$f \sin (2\theta - 2\pi - \nu) [+G'd + \frac{2}{n}H - \frac{1}{n}U - 0,008482 U = -U\eta]$$

$$f \sin (2\theta - 2\pi + \nu) [+G'd + \frac{2}{n}H - \frac{3}{n}V - 0,008482 V = -V\eta]$$

$$f \sin (2\theta - 2\pi - \nu) [+G'd + \frac{2}{n}H - \frac{3}{n}V - 0,008482 V = -V\eta]$$

§. 235.

deter-
d ali-
ribus
pen-

§. 235. Si nunc famli modo denue differendamus,

proditur: $\frac{d\psi}{dt} =$

$$f \cos 2\eta [-2aF\eta]$$

$$f \cos (2\phi - 2\pi) [-2(a + \frac{1}{n})G' - 0,008482 G']$$

$$f \cos (2\theta - 2\pi) [+G' - \frac{2}{n}H' - 0,008482 H']$$

$$f \cos \nu [F'd + F'd - J']$$

$$f \cos (2\eta - \nu) [F'd - (2\eta - 1)K']$$

$$f \cos (2\eta + \nu) [F'd - (2\eta + 1)L']$$

$$f \cos (2\phi - 2\pi - \nu) [-2(a + \frac{1}{n})M' + M' - 0,008482 M']$$

$$f \cos (2\phi - 2\pi + \nu) [-2(a + \frac{1}{n})N' - N' - 0,008482 N']$$

$$f \cos (2\phi - 2\pi - \nu) [-2(a + \frac{1}{n})O' + 2O' - 0,008482 O']$$

$$f \cos (2\theta - 2\pi - \nu) [G'd - \frac{2}{n}H' - \frac{2}{n}S' + S' - 0,008482 S']$$

$$f \cos (2\theta - 2\pi + \nu) [G'd - \frac{2}{n}H' - \frac{2}{n}T' - T' - 0,008482 T']$$

$$f \cos (2\theta - 2\pi - \nu) [G'd + \frac{2}{n}H' - \frac{1}{n}U' - 0,008482 U']$$

$$f \cos (2\theta - 2\pi + \nu) [G'd + \frac{2}{n}H' - \frac{3}{n}V' - 0,008482 V']$$

§. 236. Hinc autem sequentes eliduntur valores

$$F = 0,01273 \quad / F = 8,104833$$

$$G = 0,32213 \quad / G = 9,508032$$

$$H = 0,06976 \quad / H = 8,843500$$

$$J = -1,87800 \quad / J = 0,273710$$

C c 2

K =

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K = + 0,05615	L-K = 8,749352
L = - 0,00077	L-L = 6,888904
M = - 0,29638	L-M = 9,471854
N = + 0,00012	L-N = 6,089109
O = + 0,32287	L-O = 9,509034
S = + 0,39091	L-S = 9,592073
T = + 0,69579	L-T = 9,842475
U = - 0,07922	L-U = 8,898830
V = - 0,05141	L-V = 8,711093

§. 237. Pro distantia ergo lunae a sole curvata x =

$$\frac{(1-k)}{1-k \cos^2 \phi} \text{ erit}$$

x = Praec.

	Log. coeff.	coefficienti.	Valores
+ 0,000072f	5,860017	+ 0,000079	
+ 0,001833f	7,263216	+ 0,002005	
+ 0,000397f	6,598774	+ 0,000434	
- 0,010099f	8,028894	- 0,000634	
+ 0,000322f	6,504536	+ 0,000019	
- 0,000009f	4,644988	- 0,000080	
- 0,001699f	7,227038	- 0,000100	
+ 0,000009f	3,834293	+ 0,000000	
+ 0,001848f	7,264218	+ 0,000006	
+ 0,002237f	7,347257	+ 0,000132	
+ 0,000396f	7,597659	+ 0,000235	
- 0,000459f	6,654014	- 0,000000	
- 0,000296f	6,466277	- 0,000000	

vbi notandum est esse f = 1,093756, et / f = 0,038921.

§. 238.

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§. 238. Deinde pro motu momentaneo habebitur

$$\frac{d\phi}{dt} = \text{Praec.}$$

	Log. coeff.	coefficienti.	Valores
- 0,000146f	6,164934	- 0,000160	
- 0,003700f	7,568133	- 0,002946	
- 0,000801f	6,903691	- 0,000876	
+ 0,021577f	8,333811	+ 0,001286	
- 0,000659f	6,809453	- 0,000038	
+ 0,000009f	4,949005	+ 0,000001	
+ 0,003409f	7,531955	+ 0,000203	
- 0,000009f	4,139210	- 0,000000	
- 0,003719f	7,569135	- 0,000012	
- 0,004499f	7,652174	- 0,000267	
- 0,007999f	7,902576	- 0,000476	
+ 0,000919f	6,958931	+ 0,000000	
+ 0,000599f	6,771194	+ 0,000000	

§. 239. Pro correctione longitudinis verae hinc

oriunda penatur,

$$\phi = \text{Praec.}$$

$$\begin{aligned} &+ \mathcal{X}'/f \sin 2\phi & + \mathcal{Y}'/f \sin \phi & + \mathcal{W}'/f \sin(2\phi - 2\pi - \phi) \\ &+ \mathcal{X}''/f \sin(2\phi - 2\pi) + \mathcal{Y}''/f \sin(2\phi - \phi) + \mathcal{W}''/f \sin(2\phi - 2\pi + \phi) \\ &+ \mathcal{X}'''/f \sin(2\phi - 2\pi) + \mathcal{Y}'''/f \sin(2\phi + \phi) + \mathcal{W}'''/f \sin(2\phi - 2\pi - 2\phi) \\ &+ \mathcal{X}''''/f \sin(2\phi - 2\pi + \phi) + \mathcal{Y}''''/f \sin(2\phi - 2\pi + \phi) \end{aligned}$$

Cc 3

erit.

critique :

$$\begin{aligned}
 2\alpha \mathcal{G}' &= -0,000146 \\
 0,026834 \mathcal{G}' &= -0,003700 \\
 0,159358 \mathcal{G}' &= -0,000801 \\
 \mathcal{G}' - \mathcal{G}'/d' &= +0,02157 \\
 0,867476 \mathcal{R}' - \mathcal{G}'/d' &= -0,00065 \\
 2,867476 \mathcal{R}' - \mathcal{G}'/d' &= -0,00000 \\
 1,026834 \mathcal{M}' &= +0,00340 \\
 3,026834 \mathcal{M}' &= -0,00000 \\
 0,026834 \mathcal{Q}' &= -0,00371
 \end{aligned}$$

$$-0,840642 \mathcal{G}' - \mathcal{G}'/d' + \frac{2}{n} \mathcal{G}' = -0,00449$$

$$+1,159358 \mathcal{G}' - \mathcal{G}'/d' + \frac{2}{n} \mathcal{G}' = -0,00799$$

$$+0,083920 \mathcal{M}' - \mathcal{G}'/d' - \frac{2}{n} \mathcal{G}' = +0,00091$$

$$+0,234796 \mathcal{R}' - \mathcal{G}'/d' - \frac{2}{n} \mathcal{G}' = +0,00059$$

§. 240. Expeditis igitur his formulis orientur :

	Logcoeff.	coeff. tot. in fac.
$\Phi = \text{Pr.} - 0,0000078f \sin 2\eta$	5,893680	18"
$- 0,001825f \sin (2\Phi - 2\pi)$	7,261316	422
$- 0,004800f \sin (\frac{1}{2}\theta - 2\pi)$	7,681286	1083
$+ 0,02154f \sin r$	8,333246	+ 264
$- 0,00074f \sin (2\eta - r)$	6,867925	9
$+ 0,00332f \sin (2\Phi - 2\pi - r)$	7,520465	+ 41
$- 0,13818f \sin (2\Phi - 2\pi - 2r)$	9,140450	92
$+ 0,00446f \sin (2\theta - 2\pi - r)$	7,649421	+ 55
$- 0,00685f \sin (2\theta - 2\pi + r)$	7,855624	84
$+ 0,00310f \sin (2\theta - 2\pi - r)$	7,491107	+ 11
$- 0,00030f \sin (2\theta - 2\pi + r)$	6,474418	1

§. 242.

§. 241. Haec omnia factis conveniunt cum notis inaequalitatis motus lunae, nisi quod inaequalitas ab angulo $2\theta - 2\pi$ pendens plane adverteri videatur, cum nullum eius vestigium in tabulis astronomicis occurrat; quod quidem eo magis est mirandum, cum correctio inde oriunda ad 18', 3" exurgat. Labens equidem agnosco, in hoc calculo non omnem curam esse adhibitam, ut hanc aequationem tanquam omnibus numeris absolutam spectare liceat, quoniam ad plurimos terminos, quos formulae nostrae suppeditant, non respexi. Interim tamen calculum repetenti mox patebit, non admodum enormiter esse aberratum, praesertim cum aequatio ab angulo $2\Phi - 2\pi$ pendens, quae pari passu procedit, veritati perquam consentanea prodierit, cum ea reductio lunae ad eclipticam contineatur. Ac si quidem haec inaequalitas ad femissem vsque diminuitur, tamen tanta remanet, ut merito dubitare debeamus, eius effectum ab Astronomis non esse animadvertum; cum eius omniffo vix per aliam aequationem compendiari queat. Hancockem, siue omniffo terminorum neglectorum sit in causa, siue etiam in calculo numerico error fuerit admiffus, quod facile evenire potuit, istam investigationem in capite sequenti accuratius suscipiamus.

CAPUT XV.

ACCURATIOR INVESTIGATIO INAEQUALI-
TATUM LUNAE AB INCLINATIONE EJUS
ORBITAE PENDENTIIUM.

§. 242.

Quoniam praecipuum dubium circa inaequalitatem ab angulo $2\theta - 2\pi$ pendentem veratur, nostram investigationem ab his inaequalitatibus, quae simul ab alterutra excentricitate pendent, abstrahamus. Po-

namus ergo:

$$\int R dr = \mathcal{K} \cos 2\eta + \mathcal{L} \cos 4\eta + \mathcal{M} \cos(2\Phi - 2\pi) + \mathcal{N} \cos(2\theta - 2\pi) + \mathcal{O} \cos(4\theta - 4\pi) + \mathcal{P} \cos(2\Phi - 2\pi) + \mathcal{Q} \cos(4\theta - 4\pi)$$

$$\text{et } v = A \cos 2\eta + F \cos 4\eta + H \cos(\alpha\theta - 2\pi) + J \cos(4\theta - 4\pi)$$

$$\text{Postroque } \frac{2\mathcal{K}F + \mathcal{G}}{n\mathcal{M}} = f; \quad \frac{2\mathcal{K}G + \mathcal{O}}{n\mathcal{M}} = g; \quad \frac{2\mathcal{K}H + \mathcal{D}}{n\mathcal{M}} = b';$$

$$\frac{2\mathcal{K}J + \mathcal{Q}}{n\mathcal{M}} = j \text{ erit:}$$

$$\frac{d\Phi}{dr} = u + \frac{1}{n} \mathcal{A} \cos 2\eta - f \cos 2\eta - g \cos(2\Phi - 2\pi) - b' \cos(2\theta - 2\pi) - j \cos(4\theta - 4\pi)$$

$$\frac{d\eta}{dr} = a - \mathcal{A} \cos 2\eta - f' \cos 2\eta - g' \cos(2\Phi - 2\pi) - b' \cos(2\theta - 2\pi) - j' \cos(4\theta - 4\pi)$$

$$\frac{d\theta}{dr} = \frac{1}{n} \text{ et } \frac{d\pi}{dr} = -0,004241 \text{ --- } 0,004221 \cos 2\eta + 0,004241 \cos(2\Phi - 2\pi) + 0,004241 \cos(2\theta - 2\pi)$$

$$\text{§. 243.}$$

VALI-

hicam
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simul
Po-

 $\theta - 4\pi$ $\theta - 4\pi$ b' $\theta - 4\pi$ $\theta - 4\pi$ $\cos 2\eta$

243.

CAPUT XV.

§. 243. His valoribus in formulis principalibus substituis habebimus has aequationes:

$$R = -\frac{3\mathcal{G}}{2n\mathcal{M}} f \sin(2\theta - 2\pi) + \frac{3}{2n\mathcal{M}} H f \sin(2\Phi - 2\pi)$$

$$\frac{ddv}{dr^2} = f \cos 2\eta \left\{ -\mathcal{E}F - 2\mathcal{K}\mathcal{G} \right.$$

$$\left. \begin{aligned} & -\mathcal{E}G + \frac{3H}{4n\mathcal{M}} - 2\mathcal{K}\mathcal{O} + \frac{\mathcal{M}\mathcal{D}}{n\mathcal{M}} + \frac{3A\mathcal{D}}{n\mathcal{M}} \\ & + \frac{3\mathcal{M}H}{n\mathcal{M}} + \frac{3AH}{n\mathcal{M}} \end{aligned} \right\}$$

$$\left. \begin{aligned} & f \cos(2\theta - 2\pi) \left\{ -\mathcal{E}H + \frac{3G}{4n\mathcal{M}} - 2\mathcal{K}\mathcal{F} + \frac{\mathcal{M}\mathcal{O}}{n\mathcal{M}} + \frac{3A\mathcal{O}}{n\mathcal{M}} \right. \\ & \left. + \frac{3\mathcal{M}G}{n\mathcal{M}} + \frac{3AG}{n\mathcal{M}} \right\} \\ & f \cos(4\theta - 4\pi) [-\mathcal{E}J - 2\mathcal{K}\mathcal{Q}] \end{aligned} \right\}$$

$$f \cos 2\eta \left(-0,005156 - 0,026307 + 0,001969 \frac{A}{n\mathcal{M}} \right)$$

$$f \cos(2\Phi - 2\pi) \left(+0,003938 - 1 - 0,052614 \frac{A}{n\mathcal{M}} - 0,002578 \frac{A}{n\mathcal{M}} \right)$$

$$f \cos(2\theta - 2\pi) \left(+0,052614 + 0,002578 - 0,001969 \frac{2A}{n\mathcal{M}} + \frac{A}{n\mathcal{M}} \right)$$

$$f \cos(4\theta - 4\pi) \left(+0,001320 + 0,026307 \frac{A}{n\mathcal{M}} \right)$$

§. 244. Vel in numeris erit
 $R = 0,008540H f \sin(2\Phi - 2\pi) - 0,008540G f \sin(2\theta - 2\pi)$
 $Dd \quad ddv =$

$$\frac{dhw}{dy^2} =$$

$$f \cos 2 \eta [-1, 01591 F - 2x \textcircled{+} - 0, 031478$$

$$f \cos(2\phi - 2\pi) \{-1, 01591 G - 2x \textcircled{+} + 0, 000636 H - 0, 845648$$

$$f \cos(2\theta - 2\pi) \{-1, 01591 H - 2x \textcircled{+} + 0, 000636 G + 0, 047722$$

$$f \cos(4\theta - 4\pi) [-1, 01591 J - 2x \textcircled{+} + 0, 001123$$

Nunc autem ex formulis assumtis erit

$$R = f \sin 2 \eta [-2 a \textcircled{+} + 0, 004241 \textcircled{+} - 0, 004241 \textcircled{+}]$$

$$f \sin(2\phi - 2\pi) [+ \textcircled{+} B' - 2, 026834 \textcircled{+} + 0, 004221 \textcircled{+}]$$

$$f \sin(2\theta - 2\pi) [- \textcircled{+} g' - 0, 023965 \textcircled{+} - 0, 159358 \textcircled{+}]$$

$$f \sin(4\theta - 4\pi) [+ 0, 004241 \textcircled{+} - 0, 318716 \textcircled{+}]$$

at est

$$\textcircled{+} B' = -0, 000931 H - 0, 00461 \textcircled{+}; \textcircled{+} g' = -0, 000931 G - 0, 00461 \textcircled{+}$$

$$\text{unde fit: } 1, 867476 \textcircled{+} = 0, 004241 \textcircled{+} (\textcircled{+} - \textcircled{+})$$

§. 245. Tum simili modo differentiando valorem ipsius ν ponatur

$$F' = 1, 867476 F - 0, 004241 G + 0, 004241 H$$

$$G' = 2, 026834 G + 0, 01091 H + 0, 00750 \textcircled{+}$$

$$H' = 0, 159358 H + 0, 03909 G - 0, 00750 \textcircled{+}$$

$$J' = 0, 318716 F - 0, 004241 H$$

VI

$$\text{vt fit } \frac{d\nu}{dy} =$$

$$A' \sin 2\eta - F' / \sin 2\eta - G' / \sin(2\phi - 2\pi) - J' / \sin(4\theta - 4\pi)$$

$$- H' / \sin(2\theta - 2\pi)$$

eritque

$$\frac{d\nu}{dy^2} = f \cos 2 \eta [-2x F' + 0, 004241 G' + 0, 004241 H']$$

$$f \cos(2\phi - 2\pi) [+ A' B' - 2, 026834 G' - 0, 004221 H']$$

$$f \cos(2\theta - 2\pi) [+ A' g' - 0, 159358 H' - 0, 023965 G']$$

$$f \cos(4\theta - 4\pi) [+ 0, 004241 H' - 0, 318716 J']$$

$$\text{feu } \frac{d\nu}{dy^2} =$$

$$f \cos 2\eta \{-3, 48745 F + 0, 01668 G - 0, 00721 H - 0, 000003 \textcircled{+} + 0, 000007 \textcircled{+}$$

$$f \cos(2\phi - 2\pi) \{-4, 10820 G - 0, 05121 H - 0, 02929 \textcircled{+} + 0, 000005 \textcircled{+}$$

$$f \cos(2\theta - 2\pi) \{-0, 02565 H - 0, 08323 G - 0, 01290 \textcircled{+} - 0, 00018 \textcircled{+}$$

$$f \cos(4\theta - 4\pi) \{-0, 10158 J + 0, 00016 G + 0, 00202 H - 0, 000003 \textcircled{+}$$

§. 246. Hinc pro $\textcircled{+}$ et $\textcircled{+}$ substituimus valores

$$\textcircled{+} = 0, 00227 (\textcircled{+} - \textcircled{+}) \text{ et } \textcircled{+} = 0, 01931 \textcircled{+}$$

habebimus has aequationes

$$+ 2, 47154 F - 0, 01668 G + 0, 000721 H = 0, 031478$$

$$- 0, 00456 \textcircled{+} + 0, 00456 \textcircled{+}$$

$$+ 3, 09229 G + 0, 05185 H + 0, 00218 \textcircled{+} = 0, 995648$$

$$- 2, 01799 \textcircled{+}$$

$$- 0, 99026 H + 0, 08387 G - 0, 01421 \textcircled{+} = -0, 047722$$

$$- 2, 01780 \textcircled{+}$$

$$- 0, 91433 J - 0, 00016 G - 0, 00202 H = -0, 001123$$

$$+ 0, 00003 \textcircled{+} - 0, 02685 \textcircled{+}$$

$$D d 2 \quad \text{§. 247.}$$

VI

§. 247. Deinde reperitur

$$\begin{aligned} \textcircled{G} &= -0,00081 \text{ H} - 0,00002 \text{ G} \\ \textcircled{F} &= +0,11201 \text{ G} + 0,00107 \text{ H} \end{aligned}$$

qui valores substitutioni praebent:

$$\begin{aligned} +2,47154 \text{ F} &- 0,01618 \text{ G} + 0,00725 \text{ H} = 0,031478 \\ +3,09256 \text{ G} &+ 0,06064 \text{ H} = 0,995648 \\ +0,99229 \text{ H} &+ 0,14239 \text{ G} = 0,047722 \\ -0,91430 \text{ J} &- 0,00297 \text{ G} - 0,00205 \text{ H} = -0,001123 \end{aligned}$$

vide tandem reperitur

$$\begin{aligned} \text{F} &= +0,014830 \dots \text{ I F} = 8,171166 \\ \text{G} &= +0,321910 \dots \text{ I G} = 9,507734 \\ \text{H} &= +0,001901 \dots \text{ I H} = 7,278927 \\ \text{J} &= +0,000180 \dots \text{ I J} = 6,256381 \end{aligned}$$

aeque

$$\begin{aligned} \textcircled{F} &= -0,000080 \dots \text{ I } \textcircled{F} = 5,903090 \\ \textcircled{G} &= -0,000023 \dots \text{ I } \textcircled{G} = 5,361728 \\ \textcircled{H} &= +0,036056 \dots \text{ I } \textcircled{H} = 8,556977 \\ \textcircled{J} &= +0,000480 \dots \text{ I } \textcircled{J} = 6,681241 \end{aligned}$$

§. 248. Hinc ergo pro distantia $x = \frac{(1-kb) \sin}{1-kt \cos}$

reperitur	log.coeff.	val.coeff.
\textcircled{F}	5,926350	+0,000092
\textcircled{G}	7,262918	+0,002004
\textcircled{H}	5,034111	+0,000012
\textcircled{J}	4,011565	+0,000001

At

At pro motu momentaneo erit

log.coeff. val.coeff.

\textcircled{F}	6,227887	-0,000185
\textcircled{G}	7,567849	-0,004043
\textcircled{H}	6,356026	-0,000248
\textcircled{J}	4,698970	-0,000005

vide quidem iam patet inaequalitatem ab angulo $2\theta-2\pi$ pendente multo fore minorem, quam supra innueneramus, in quo non parvum veritatis criterium cernitur.

§. 249. Pro ipsa iam longitudine lunae ponamus:

$$\begin{aligned} \phi &= \text{Pr.} + \mathcal{X}' \sin 2\pi + \mathcal{Y}' \sin 2\pi + \mathcal{Z}' \sin(2\phi-2\pi) \\ &\quad + \mathcal{J}' \sin(2\theta-2\pi) + \mathcal{K}' \sin(4\theta-4\pi) \end{aligned}$$

aeque obtinebimus has aequationes:

$$\begin{aligned} 2 \text{ a } \mathcal{X}' &- 0,004241 (\mathcal{X}' + \mathcal{Y}') = -0,000169 \\ 2,026834 \mathcal{X}' &+ 0,004221 \mathcal{Y}' - 0,000227 \mathcal{Z}' = -0,003697 \\ 0,159358 \mathcal{Y}' &+ 0,023965 \mathcal{Z}' - 0,003697 \mathcal{J}' = -0,000227 \\ 0,318716 \mathcal{J}' &- 0,004241 \mathcal{K}' = -0,000005 \end{aligned}$$

vide erit

ϕ	log. coeff.	val.coeff.
\textcircled{F}	5,984018	-22 ^u
\textcircled{G}	7,260310	-41 ^u
\textcircled{H}	6,959131	-20 ^u
\textcircled{J}	5,450835	6

Pater ergo reuera aequationem ab angulo $2\theta-2\pi$ ortam multo esse minorem, quam capite praecedente innueneramus; atque nunc quidem non vltra 205^u seu 3^u, 25^u ascendere. Nullum igitur est dubium, quin haec aequatio tabulas lunares ad multo maiorem perfectionem fietur.

Dd 3

§. 250.

$$\begin{aligned}
 & -6T-2k\mathcal{E} + \frac{1}{2}bH + 0,00854G + \frac{6}{m}G + \frac{36}{2m}H \\
 & + 0,55422\mathcal{G} + 0,51332G - \frac{3AG}{2m} + 0,00940 \\
 & - 0,01049 + 0,026307 + 0,00129 - 0,00556\frac{A}{m} \\
 & - 0,00145\frac{A}{m} - 0,003938\frac{A}{m} + \frac{A}{2m} \\
 & - 0,0394\frac{D}{m} + \frac{D}{m}
 \end{aligned}$$

§. 252. Hæc autem formulæ intricatæ reducuntur ad sequentes

$$\begin{aligned}
 \frac{d\theta}{dr} &= Pr + f k \cos r (-EJ - 2x\mathcal{G} - 0,01182) \\
 & f k \cos(2\phi - 2\pi - r) (-6M - 2k\mathcal{M}) + 0,03207St + 0,53619 \\
 & - 0,02711\mathcal{G} \\
 & f k \cos(2\phi - 2\pi - 2r) (-6O - 2k\mathcal{D}) + 1,52115M + 0,76615 \\
 & f k \cos(2\theta - 2\pi - r) (-6S - 2k\mathcal{C}) - 0,03207M + 0,00286 \\
 & - 0,02711\mathcal{M} \\
 & f k \cos(2\theta - 2\pi + r) (-6T - 2k\mathcal{E}) + 0,38147
 \end{aligned}$$

§. 253. Nunc eosdem valores ex formulis affirmatis eruanus, ac posito more ad huc recepto $\frac{2xJ + \mathcal{G}}{m} = u$,

$$\begin{aligned}
 \frac{2xM + \mathcal{M}}{m} &= m' \text{ etc.} & \text{erit} \\
 \frac{d\theta}{dr} &= u + \frac{1}{2} - d' \cos 2r - 2k \cos(2r - r) - g' f \cos(2\phi - 2\pi) \\
 & - h' k \cos(2\eta + r) - b' f \cos(2\theta - 2\pi) \\
 & - m' f k \cos(2\phi - 2\pi - r) - s' f k \cos(2\theta - 2\pi - r) \\
 & - n' f k \cos(2\phi - 2\pi - 2r) - t' f k \cos(2\theta - 2\pi + r)
 \end{aligned}$$

$$\begin{aligned}
 \frac{d\eta}{dr} &= + \frac{36}{m}H \\
 & 0940 \\
 & \frac{556}{m}A \\
 & + \frac{A}{2m}
 \end{aligned}$$

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$$\begin{aligned}
 \frac{d\eta}{dr} &= u - d' \cos 2r - e' k \cos r - h' k \cos(2\eta - r) - g' k \cos(2\phi - 2\pi) \\
 & - b' k \cos(2\eta + r) - b' k \cos(2\theta - 2\pi) \\
 & - m' f k \cos r - n' f k \cos(2\phi - 2\pi - r) - s' f k \cos(2\theta - 2\pi - r) \\
 & - t' f k \cos(2\phi - 2\pi - 2r) - t' f k \cos(2\theta - 2\pi + r)
 \end{aligned}$$

$$\frac{d\theta}{dr} = \frac{1}{u} + \frac{2}{m} k \cos r + \frac{3}{2m} k k \cos 2r \text{ atque}$$

$$\begin{aligned}
 \frac{d\pi}{dr} &= + 0,00424r + 0,004241 \cos(2\phi - 2\pi) \\
 & - 0,004221 \cos 2r + 0,004241 \cos(2\theta - 2\pi) \\
 & - 0,01766k \cos r - 0,00966k \cos(2\eta - r) - 0,010636k \cos(2\eta - 2r) \\
 & - 0,00831k \cos(2\eta + r) - 0,021273k \cos 2r \\
 & + 0,00924k \cos(2\phi - 2\pi - r) + 0,00841k \cos(2\theta - 2\pi - r) \\
 & + 0,00841k \cos(2\phi - 2\pi + r) + 0,00924k \cos(2\theta - 2\pi + r) \\
 & + 0,01064k \cos(2\phi - 2\pi - 2r)
 \end{aligned}$$

§. 254. His iam valoribus introducendis differentemur formulas nostras assumas pro $\int R dr$ et v ; atque obtinebimus primo:

$$\begin{aligned}
 R = \text{Præc.} \\
 & + 0,00924\mathcal{G} + 0,00841\mathcal{D} - 0,004241\mathcal{M} \\
 & - 0,00841\mathcal{G} - 0,00924\mathcal{D} - 0,004241\mathcal{E} \\
 & + 0,01766\mathcal{G} - 0,009966\mathcal{D} \\
 & - 1,026834\mathcal{M} - 0,004221\mathcal{E} \\
 & + \mathcal{D}' + 32M - 0,02127\mathcal{G} - 0,01064\mathcal{D} \\
 & - 0,026834\mathcal{D} - 0,01766\mathcal{M} - 0,009966\mathcal{E}
 \end{aligned}$$

Et

$$\begin{aligned}
 & + f \sin(2\theta - 2\pi - r) \left\{ \begin{aligned} & - M' + D' + G' - \frac{2}{n} G' + 0,840642 S + 0,000831 G + 0,004221 M \\ & - G' - 0,000831 G - 0,017666 H - 0,004221 M \end{aligned} \right\} \\
 & + f \sin(2\theta - 2\pi + r) \left\{ \begin{aligned} & - D' + E' + G' - \frac{2}{n} G' - 1,159358 S \\ & + 0,009966 G - 0,017666 H \end{aligned} \right\}
 \end{aligned}$$

§. 255. Hinc igitur consequimur illas aequationes

$$\begin{aligned}
 0,000831 (G - G') - G - 0,004241 (M + G - G') & = 0 \\
 M' - 0,017666 G - 0,009966 H - 0,004221 G - 1,026834 M & = 0,01708 H + 0,00854 S \\
 D' + 32H - 0,002127 G - 0,01064 H - 0,026834 D - 0,017666 M & = -0,009966 G - 0,01708 S
 \end{aligned}$$

$$\begin{aligned}
 -M' + D' - G' + G' - 0,00831 G - \frac{2}{n} G' - 0,017666 H + M' & \\
 - 0,004221 M + 0,840642 S & = -0,01708 G - 0,00854 M \\
 -D' + E' + G' - 0,009966 G - \frac{2}{n} G' - 0,017666 H - 1,159358 S & = -0,01708 G
 \end{aligned}$$

§. 256. Ponatur nunc videntur:

$$\begin{aligned}
 J' & = G - 0,000831 (G - H) + 0,004241 (M + S - T) \\
 M' & = 1,026834 M - A' + 0,017666 G + 0,009966 H + 0,004221 S \\
 O' & = 0,026834 O - D' + 14H + 0,02127 G + 0,01064 H \\
 & \quad + 0,017666 M + 0,009966 S \\
 S' & = -0,840642 S + A' - D' + E' - G' + 0,00831 G + \frac{2}{n} H \\
 & \quad + 0,017666 H - M' + 0,004221 M \\
 T' & = 1,159358 T + D' - E' - G' + 0,009966 G + \frac{2}{n} H \\
 & \quad + 0,017666 H
 \end{aligned}$$

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$$\frac{d'u}{d'r} = \text{Præc.}$$

$$\begin{aligned}
 & + f \cos r \left[+ 0,017666 (G' + H') - J' + 0,004241 (M' + S' + T') \right. \\
 & \quad \left. + f \cos(2\theta - 2\pi - r) \left\{ \begin{aligned} & M' - 0,017666 G' - 0,009966 H' \\ & - 0,004221 S' - 1,026834 M' \end{aligned} \right\} \right.
 \end{aligned}$$

$$\begin{aligned}
 & + f \sin \cos(2\theta - 2\pi - r) \left\{ \begin{aligned} & D' - 2,66H - 0,0127G' - 0,01064 H' \\ & - 0,017666 M' - 0,009966 S' \\ & - 0,026834 O' \end{aligned} \right\} \\
 & + f \sin \cos(2\theta - 2\pi + r) \left\{ \begin{aligned} & A' + D' + E' + G' + 0,840642 S' \\ & - 0,000831 G' - \frac{2}{n} H' - 0,017666 H' \\ & + M' - 0,004221 M' \end{aligned} \right\}
 \end{aligned}$$

$$\begin{aligned}
 & + f \sin \cos(2\theta - 2\pi + r) \left\{ \begin{aligned} & D' + E' + G' - 0,009966 G - \frac{2}{n} H' \\ & - 0,017666 H' - 1,159358 T' \end{aligned} \right\}
 \end{aligned}$$

§. 257. Prioris ordinis aequationes huc reducuntur:

$$\begin{aligned}
 G & = -0,004241 (M + G + S) \\
 1,026834 M & = -0,01785 S - 0,00883 G - 0,00039 \\
 0,026834 O & = -0,05835 S - 0,03041 G \\
 & \quad - 0,017666 M + 0,00690 \\
 -0,840642 S & = +0,01785 M - 0,01935 M + 0,00263 \\
 1,159358 S & = +0,01247
 \end{aligned}$$

Hinc fit

$$\begin{aligned}
 G & = + 0,00007 S + 0,00008 M + 0,00006 S \\
 M & = - 0,01738 S + 0,00018 M - 0,00035 \\
 O & = - 2,16266 S + 0,02394 M + 0,26092 \\
 S & = - 0,00040 S - 0,02123 M - 0,00313 \\
 S & = - 0,01076
 \end{aligned}$$

E e 2 §. 258.

§. 258. Porro reperitur

$$\begin{aligned}
 J' &= J + 0,004241(M+S-T) - 0,00026 \\
 M' &= 1,026834M + 0,01935S - 0,00016M + 0,00368 \\
 O' &= 0,026834O - 0,37652S + 0,01805M + 0,01063 \\
 S' &= 0,840642S + 0,00013S + 0,00883M - 0,00167 \\
 T' &= 1,159358T + 0,01023
 \end{aligned}$$

ac functionibus habebitur $\frac{d\theta}{dt} = \text{Præc.}$

$$+fk \operatorname{col} r [-J' + 0,004241(M+S+T) + 0,01175]$$

$$+fk \operatorname{col}(2\Phi - 2\pi - r) \left\{ \begin{array}{l} -1,026834M' - 0,00422S' - 0,02842S' \\ + 0,00030M - 0,01160 \end{array} \right\}$$

$$+fk \operatorname{col}(2\Phi - 2\pi - 2r) \left\{ \begin{array}{l} -0,026834O' - 0,01766M' - 0,00354M' \\ -0,01511 - 0,00996S' + 0,333746S \end{array} \right\}$$

$$+fk \operatorname{col}(2\theta - 2\pi - r) \left\{ \begin{array}{l} +0,840642S' - 0,02396M' - 0,02843M' \\ -0,01472 + 0,00025S \end{array} \right\}$$

$$+fk \operatorname{col}(2\theta - 2\pi + r) [-1,159358T' + 0,33836]$$

§. 259. Hinc tandem valores quæsi eliciuntur

$$\begin{aligned}
 J &= 0,81144 & L-J &= 9,909256 \\
 M &= 1,25325 & L-M &= 0,098046 \\
 O &= 2,12630 & L-O &= 0,327624 \\
 S &= 0,13490 & L-S &= 9,130012 \\
 T &= 0,10080 & L-T &= 9,003441
 \end{aligned}$$

22

ac $\mathcal{Q} = -0,00005 \dots L\mathcal{Q} = 5,698970$

$$\mathcal{M} = +0,00177 \dots L\mathcal{M} = 7,247973$$

$$\mathcal{D} = +0,52266 \dots L\mathcal{D} = 9,718219$$

$$\mathcal{C} = +0,02355 \dots L\mathcal{C} = 8,371991$$

$$\mathcal{E} = +0,01076 \dots L\mathcal{E} = 8,031812$$

§. 260. Ex his iam pro distantia Lunæ a terra erit

$\mu = \text{Præc.}$

	Log. coeff.	Valor. coeff. totius
$0,00462fk \operatorname{col} 2\mu$	7,664440	0,000275
$0,00773fk \operatorname{col}(2\Phi - 2\pi - r)$	7,853230	0,000424
$0,01210fk \operatorname{col}(2\Phi - 2\pi - 2r)$	8,082808	0,000039
$0,00077fk \operatorname{col}(2\theta - 2\pi - r)$	6,885196	0,000046
$0,00057fk \operatorname{col}(2\theta - 2\pi + r)$	6,758625	0,000034

et pro motu momento

$\frac{d\Phi}{dt} = \text{Præc.}$

	Log. coeff.	Val. coeff.
$0,00932fk \operatorname{col} r$	7,960416	+ 0,000555
$0,01558fk \operatorname{col}(2\Phi - 2\pi - r)$	8,192568	+ 0,000928
$0,02142fk \operatorname{col}(2\Phi - 2\pi - 2r)$	8,330819	+ 0,000069
$0,00142fk \operatorname{col}(2\theta - 2\pi - r)$	7,152288	+ 0,000085
$0,00109fk \operatorname{col}(2\theta - 2\pi + r)$	7,037426	+ 0,000064

§. 261. Pro longitudine autem lunæ frequentes resoluti debent æquationes:

$$\begin{aligned}
 +0,00932 &= \mathcal{Q}' - 0,01766(\mathcal{M}' + \mathcal{S}') - 0,004241(\mathcal{M}' + \mathcal{C}' + \mathcal{E}') \\
 +0,01558 &= 1,026834\mathcal{M}' - \mathcal{M}' + 0,01766\mathcal{O}' + 0,00996\mathcal{S}' \\
 &\quad + 0,004221\mathcal{C}'
 \end{aligned}$$

Ee 3

+

+ 0,02142 = 0,026834 $\mathcal{D}' - \mathcal{D}' + 2,664 + 0,02127 \mathcal{U}'$
 + 0,01064 $\mathcal{U}' + 0,01766 \mathcal{M}' + 0,00996 \mathcal{C}'$
 + 0,00142 = -0,840642 $\mathcal{C}' - \mathcal{M}' - \mathcal{D}'/h' - \mathcal{U}'/g' + 0,02114 \mathcal{U}'$
 + 0,16853 $\mathcal{U}' + 0,02396 \mathcal{M}'$
 + 0,00109 = 1,159358 $\mathcal{C}' - \mathcal{D}'/g' - \mathcal{U}'/h' - 0,35615 \mathcal{U}'$
 + 0,16850 \mathcal{U}'
 hincque prodit pro longitudine vera :

$\Phi = Pr.$		Log. coeff.	Val. coeff.
+ 0,000932 f_k	$\sin \nu'$	7,9699416	+ 115
+ 0,01521 f_k	$\sin (2\Phi - 2\pi - \nu')$	8,182130	+ 187
+ 0,79079 f_k	$\sin (2\Phi - 2\pi - 2\nu')$	9,898060	+ 529
- 0,00121 f_k	$\sin (2\theta - 2\pi - \nu')$	7,083939	- 15
- 0,00082 f_k	$\sin (2\theta - 2\pi + \nu')$	6,913527	- 10

§. 262. Ob inclinationem ergo orbitae lunaris ad ellipticam omnes correctiones huc redeunt, vt sit

I. Pro distantia lunae a terra :

$\# = Praec.$		Log. coeff.	Val. coeff.
+ 0,000084 f	$\cos 2\eta$	5,926350	+ 0,000092
+ 0,001832 f	$\cos (2\Phi - 2\pi)$	7,262918	+ 0,002004
+ 0,000011 f	$\cos (2\theta - 2\pi)$	5,034111	+ 0,000012
+ 0,000001 f	$\cos (4\theta - 4\pi)$	4,011505	+ 0,000001
- 0,00465 f_k	$\cos \nu'$	7,664440	- 0,000275
- 0,00773 f_k	$\cos (2\Phi - 2\pi - \nu')$	7,853230	- 0,000424
- 0,01210 f_k	$\cos (2\Phi - 2\pi - 2\nu')$	8,082808	- 0,000039
- 0,00077 f_k	$\cos (2\theta - 2\pi - \nu')$	6,885196	- 0,000046
- 0,00057 f_k	$\cos (2\theta - 2\pi + \nu')$	6,758625	- 0,000034

H. Pro

II. Pro motu momentaneo :

$\frac{d\Phi}{dt} = Praec.$		Log. coeff.	Val. coeff.
- 0,000169 f	$\cos 2\eta$	6,227887	- 0,000185
- 0,003697 f	$\cos (2\Phi - 2\pi)$	7,567849	- 0,004043
- 0,000227 f	$\cos (2\theta - 2\pi)$	6,356026	- 0,000248
- 0,000005 f	$\cos (4\theta - 4\pi)$	4,698970	- 0,000005
+ 0,00932 f_k	$\cos \nu'$	7,9699416	+ 0,000555
+ 0,01558 f_k	$\cos (2\Phi - 2\pi - \nu')$	8,192568	+ 0,000928
+ 0,02142 f_k	$\cos (2\Phi - 2\pi - 2\nu')$	8,330819	+ 0,000069
+ 0,00142 f_k	$\cos (2\theta - 2\pi - \nu')$	7,152288	+ 0,000085
+ 0,00109 f_k	$\cos (2\theta - 2\pi + \nu')$	7,037426	+ 0,000064

III. Pro longitudine lunae vera

$\Phi = Pr.$		Log. coeff.	Val. coeff.
- 0,000096 f	$\sin 2\eta$	5,984018	- 226
- 0,001823 f	$\sin (2\Phi - 2\pi)$	7,260310	- 411
- 0,000910 f	$\sin (2\theta - 2\pi)$	6,959131	- 205
- 0,0000028 f	$\sin (4\theta - 4\pi)$	5,450835	- 6
+ 0,00932 f_k	$\sin \nu'$	7,9699416	+ 115
+ 0,01521 f_k	$\sin (2\Phi - 2\pi - \nu')$	8,182130	+ 187
+ 0,79079 f_k	$\sin (2\Phi - 2\pi - 2\nu')$	9,898060	+ 529
- 0,00121 f_k	$\sin (2\theta - 2\pi - \nu')$	7,083939	- 15
- 0,00082 f_k	$\sin (2\theta - 2\pi + \nu')$	6,913527	- 10

CAPUT

CAPUT XVI.

EXPOSITIO INÆQUALITATUM LUNAE
HÆTENUS INVENTARUM.

§. 263.

Quas igitur invenimus hætenus lunæ inæqualitates eae primum, si originem earum spectemus, ad sex classes reducuntur. Quatenus enim luna in motu suo a regulis Keplerianis, in quibus quidem motum apogei competimus, recedit, eius errores vel primo a solo lunæ aspectu, seu eius distantia a sole pendunt, seu quod eodem redit, per angulum γ tantum designantur, quibus variatio lunæ continetur. Ad secundam classem refero eas lunæ inæqualitates, quae insuper ab excentricitate eius orbitæ pendunt. Tertia classis eas complectitur inæqualitates, quae ab excentricitate orbitæ solis trahunt. Quarta vero eas, quae per utramque excentricitatem coniunctim determinantur. Quintae porro classis annumeramus eas inæqualitates, quae parallaxin solis involvant, atque errores quatuor ante memoratorum generum implicent. Sexta denique classis suppediat eas inæqualitates, quae praeterea ab inclinatione orbitæ lunaris ad eclipticam pendunt.

§. 264. Quodsi vero ad usum harum inæqualitatum attendamus, prout eae, ad lunam accommodari debent, tum eae in quinque classes commodissime distribuuntur. Primo enim perpendendae sunt eae inæqualitates

qualitates, quarum ope vera distantia lunæ a terra determinatur, ut inde porro tam lunæ diameter apparens, quam eius parallaxis horizontalis assignari possit. Secundo loco formulæ erunt collocandæ illæ, quae motui momentaneo designando inserviunt, ex quibus deinceps motus lunæ horarius accurate exhiberi poterit. Tertium locum occupabunt eae inæqualitates, quae veram longitudinem lunæ ad eclipticam relativam præbeant. Quarto vero positio lineæ nodorum lunæ, seu longitudo nodi ascendentis; ac quinto vera inclinatio orbitæ lunaris ad eclipticam inveniri debet; ut deinde vera lunæ latitudo concludi possit. Manifestum autem est, has inæqualitates plurimum inter se permixtas, ita ut vix vllum habeatur genus, cuius inæqualitates non a reliquis generibus pendeant; cui tamen incommodo facile medela adhibetur.

§. 265. Quamquam numerus inæqualitatum, quas sumus consecuti, tantopere increvit, ut calculus sine maxima molestia expediri nequeat, tamen iam monui, non omnes inæqualitates, quibus motus Lunæ perturbatur, esse designatas, sed potius earum numerum omnino esse incertum. Facile quidem intelligitur, plerasque has præterminatas inæqualitates nullius fere esse momenti, atque sine notabili errore iis supersederi posse: verum tamen sunt inter eas nonnullæ, quae ad plura minuta secunda affurgere videntur, quarum argumenta supra iam innui; ex quo omnino operæ esset pretium in eas omnia cura inquirere. Sed earum inuestigatio tam est lubrica et incerta, ut leuissima omniffo in calculo facta eas maxime

maxime afficiat. Cum igitur in calculo plurimos terminos reicere cogamur, istam investigationem frustra plane suscipere, quamdiu scilicet rem sine approximatione exequi non licet. Cuius defectus eximium habemus exemplum in inaequalitatibus postremo loco inuentis, quae statim atque in negligendo minus largi fuerant, mirum quantum prodierunt immutatae; ac nullum plane est dubium, si calculum adhuc accuratius profsequi liceret, quin valores inueniri notabilem insuper mutationem sine subirent. Imprimis autem aequatio ab angulo $2\Phi - 2\pi - 2r$ seu a dupla distantia apogei a nodo pendens, est suspecta, ac minime pro certa haberi potest, cum leuissima circumstantia eam magnopere perturbare valeat.

§. 266. Si enim in causam inquiramus, cur analysis posterior tam diuersos valores pro his inaequalitatibus suppeditauerit, primo quidem statim patet, neglectum litterarum germanicarum S , M , O , etc. in calculo priori possimum hoc diffectionem produxisse: ingens enim valor litterae O imprimis aequationem ab angulo $2\Phi - 2\pi - 2r$ pendentem tanroperie auxit. Praeterea vero etiam non paruum augmenti haec aequatio inde est nata, quod in calculo posteriori rationem quoque habiturus termini $\cos(2\eta - 2r)$, qui tam in valore $\frac{d\Phi}{dr}$ quam $\frac{d\eta}{dr}$ inesse est deprehensus; vnde tuto colligere licet, si alios quoque terminos similes veluti $2\theta - 2\pi - 2r$, etc. per se sunt misam, in calculum introduxissemus, coefficientien-

ter-
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efficientes terminorum $2\Phi - 2\pi - 2r$ non medicorum inde mutationem subireuros fuisse. Quamobrem plus hinc colligere non possumus, nisi inaequalitatem Lunae ab angulo hoc $2\Phi - 2\pi - 2r$ pendentem minime esse conueniendam, etiam si fortasse tanta non sit, quam inuenimus. Vera autem eius quantitas certius ex observationibus quam ex Theoria colligi posse videtur.

§. 267. Quoniam vero hae inaequalitates omnes ad anomaliam Lunae veram referuntur, antequam eas ad usum adhibere liceat, modum triidi conueniet ad quodvis tempus propositum anomaliam Lunae veram determinandi. Cognita autem excentricitate orbitae lunaris k et motu anomaliae mediae, inde ad quodvis tempus facile anomalia media p colligitur. Verum ex anomalia media p et excentricitate k anomalia vera r definiti debet ope huius aequationis $dp = \frac{(1-k)^{\frac{3}{2}} dr}{(1-k \cos^2 r)^2}$; vnde quidem non difficulter, si nota esset anomalia vera r , vicissim inueniri posset anomalia media p . Calculo enim peracto, si breuitatis gratia ponatur

$$\delta = \frac{1 - \sqrt{1 - k^2}}{k} = \frac{1}{2}k + \frac{1.1.k^3}{2.4} + \frac{1.1.3.k^5}{2.4.6} + \text{etc.}$$

reperitur:

$$p = r + 2k \sin r + 2\delta(k - \frac{1}{2}\delta) \sin 2r + 2\delta^2(k - \frac{3}{2}\delta) \sin 3r + 2\delta^3(k - \frac{3}{2}\delta) \sin 4r \text{ etc.}$$

cuius seriei progressio primo incertum patet. Cum autem k ac proinde δ sit valde paruum, erit satis exacte:

$$p = r + 2k \sin r + (\frac{3}{2}kk + \frac{1}{2}k^4 + \frac{3}{2}k^6) \sin 2r + (\frac{5}{2}k^3 + \frac{1}{2}k^5) \sin 3r + (\frac{7}{2}k^4 + \frac{1}{2}k^6) \sin 4r + \frac{9}{2}k^5 \sin 5r + \frac{11}{2}k^6 \sin 6r$$

§. 268. Data ergo anomalia media Lunae p , eius anomalia vera r elici debet ex hac aequatione

$$r = p - 2k \sin r - (\frac{3}{2}kk + \frac{1}{2}k^4 + \frac{3}{2}k^6) \sin 2r - (\frac{5}{2}k^3 + \frac{1}{2}k^5) \sin 3r - (\frac{7}{2}k^4 + \frac{1}{2}k^6) \sin 4r - \frac{9}{2}k^5 \sin 5r - \frac{11}{2}k^6 \sin 6r$$

eius quidem ope, si cognita fuerit excentricitas k , calculus non solum expedietur. Quoniam enim termini sinus involuentes sunt admodum parvi, in his statim poni poterit $r = p$, unde valor verior pro r eruetur, qui deinde iterum in his terminis adhibitus, inflorem valorem pro r suppeditabit. Argue hoc modo post aliquot operationes versus tandem valor pro anomalia vera r obtinebitur. Interim tamen, quo ite calculus facilius perfici queat, aequatio haec ita potest transformari, ut loco sinuum anomaliae verae r , sinus anomaliae mediae p introducatur; id quod sequenti modo praestabitur.

§. 269. Ponatur brevitatis gratia:

$$\frac{1}{3} + \frac{1}{3}kk + \frac{1}{3}k^4 = a; \quad \frac{1}{5} + \frac{1}{5}kk = b; \quad \frac{1}{7} + \frac{1}{7}kk = c$$

ut sit:

$$r = p - 2k \sin r - ak \sin 2r - bk^3 \sin 3r - ck^5 \sin 5r - dk^7 \sin 7r$$

ac ponatur denuo

$$2k \sin r + ak \sin 2r + bk^3 \sin 3r + ck^5 \sin 5r + dk^7 \sin 7r = R$$

$k^2 \sin 3r$

$n6r$

p , eius

$\sin 3r$

$\sin 6r$

k , calculi terminis statim eruetur, in modo d' anomaliae calculi potest sinus quendi

$k = y$

$\sin 6r$

$= R$

et

et cum sit $r = p - R$ habebitur:

$$\begin{aligned} \sin r &= (1 - \frac{1}{2}R)R + \frac{1}{2}R^3 + \frac{1}{24}R^5 \text{ cof } p \\ \sin 2r &= (1 - 2R)R + \frac{1}{2}R^3 \text{ cof } 2p \\ \sin 3r &= (1 - \frac{2}{3}R)R \text{ cof } 3p \\ \sin 4r &= (1 - 8R)R \text{ cof } 4p \\ \sin 5r &= \sin 5p - 5R \text{ cof } 5p \\ \sin 6r &= \sin 6p \end{aligned}$$

negligendo scilicet terminos, qui ipsius k potestates sex vel superiores continent.

§. 270. Evolutio autem huius calculi sit maxime prolixæ, si quidem ad sextam potestatem ipsius k ascendere velimus. Facile autem expedietur, si ad quartam substituamus, cum autem reperitur:

$$r = p - (2k - \frac{1}{2}k^3) \sin p + (\frac{5}{2}kk - \frac{1}{2}k^4) \sin 2p - \frac{1}{2}k^3 \sin 3p - \frac{1}{2}k^5 \sin 4p$$

quæ expressio satis accurate pro quantis anomalia media p conveniensem anomaliæ veram indicabit. Calculus autem, si excentricitas k constet, facili negotio absolvetur. Formula hæc quoque ita representari poterit, ut sit

$$r = p - 2k(1 - \frac{1}{3}kk) \sin p + \frac{1}{2}kk(1 - \frac{1}{3}kk) \sin 2p - \frac{1}{2}k^3 \sin 3p - \frac{1}{2}k^5 \sin 4p$$

Hinc igitur tabulam constructi conveniet, unde pro quavis anomalia media proposita ipsa respondens anomalia vera excerpri queat.

§. 271. Cum autem inventa fuerit anomalia vera r , longitudo Lunæ regula Kepleriana invenienda, quam

Rf 3

supra

supra (206) per § indicavimus, ita exprimitur, ut sit

$$e = C + 4,0085272 p$$

$$-1,0085272 \left(2k \sin^2 \frac{1}{2} k (1 + \frac{2}{3} k^2 + \frac{7}{15} k^4) \sin 2p + \frac{1}{3} k^2 (1 + \frac{2}{3} k^2) \sin^2 p \right) + \frac{1}{3} k^2 (1 + \frac{2}{3} k^2) \sin 4p + \frac{2}{15} k^2 \sin^2 p + \frac{7}{15} k^4 \sin 6p$$

vbi $C + 1,0085272 p$ exhibet longitudinem Lunae mediam; quae si vocetur e , atque in coefficientum partibus minimis pro k scribatur valor proximus 0,0545, erit

Mag. coeff. val. in min. sec.

$$e = 3 - 2,0170544 k \sin p - 0,30471822675 k^2 = 6^{\circ}, 17', 55''$$

$$- 0,756770 k k \sin 2p - 9,878964 k^2 = 464 = 7,44$$

$$- 0,336551 k^3 \sin 3p - 9,527051 k^4 = 11 k$$

$$- 0,15786 k^5 \sin 4p - 9,198282 k^6 = 3$$

unde patet superfluum futurum fuisse, si superiores expressiones ultra quartam potestatem ipsius k extendere voluissent.

CAPUT

ut sit

$$\frac{1}{2} k k \sin 2p$$

nae me-

um par-

0,0545,

n. sec.

, 17', 55''

7,44

res ex-

tendere

CAPUT

CAPUT XVII.

INVESTIGATIO ELEMENTORUM

MOTUS LUNAE

§. 272.

Inventis iam per Theoriam hisce inaequalitatibus, quibus motus Lunae perturbatur, aequam eas ad computum astronomicum accommodare liceat; elementa, quae in eas ingrediuntur, per observationes determinari oportet. Primo scilicet ad datam epocham cum longitudo Lunae media, tum eius anomalia media, ac locus nodi medius constitui debet, ut eadem res inde ad quodvis aliud tempus assignari queant. Deinde quoque ex observationibus verus valor eccentricitatis lunaris colligi debet, a quo potissimum quantitas praeparatum inaequalitatum pendet. Excentricitas autem orbitae solaris pro factis certa haberi poterit, cum sit $e = 0,0168$. Lunae vero excentricitas tam propere iam constat, ut inde sine errore ad quamlibet anomalias mediam vera factis exacte assignari possit. Eius enim in anomalia vera error aliquot minorum primorum committitur, inaequales Lunae inde non ultra aliquot minuta secunda afficiuntur.

§. 273. Quodsi autem factam quasvis Lunae observationes ad hanc finem adhibere velimus, ob tam ingentem inaequalitatum numerum, investigatio elementorum maxime molesta redderetur. Quocirca ex obser-